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Dr. GREGORY's ELEMENTS OF CATOPTRICS AND DIOPTRICS.

To which is added,

- I. A Method for finding the *Foci* of all *Specula* as well as *Lens's* universally. As also for *Magnifying* or *Lessening* a given Object by a given *Speculum* or *Lens* in any assign'd Proportion, &c.
- II. A *Solution* of those *Problems* which are left undemonstrated.
- III. A particular Account of *Microscopes* and *Telescopes*, from Mr. *Huygens*.

WITH

An INTRODUCTION shewing the Discoveries
made by *Catoptrics* and *Dioptrics*.

By W. BROWNE, A. M. & Med. Pract.

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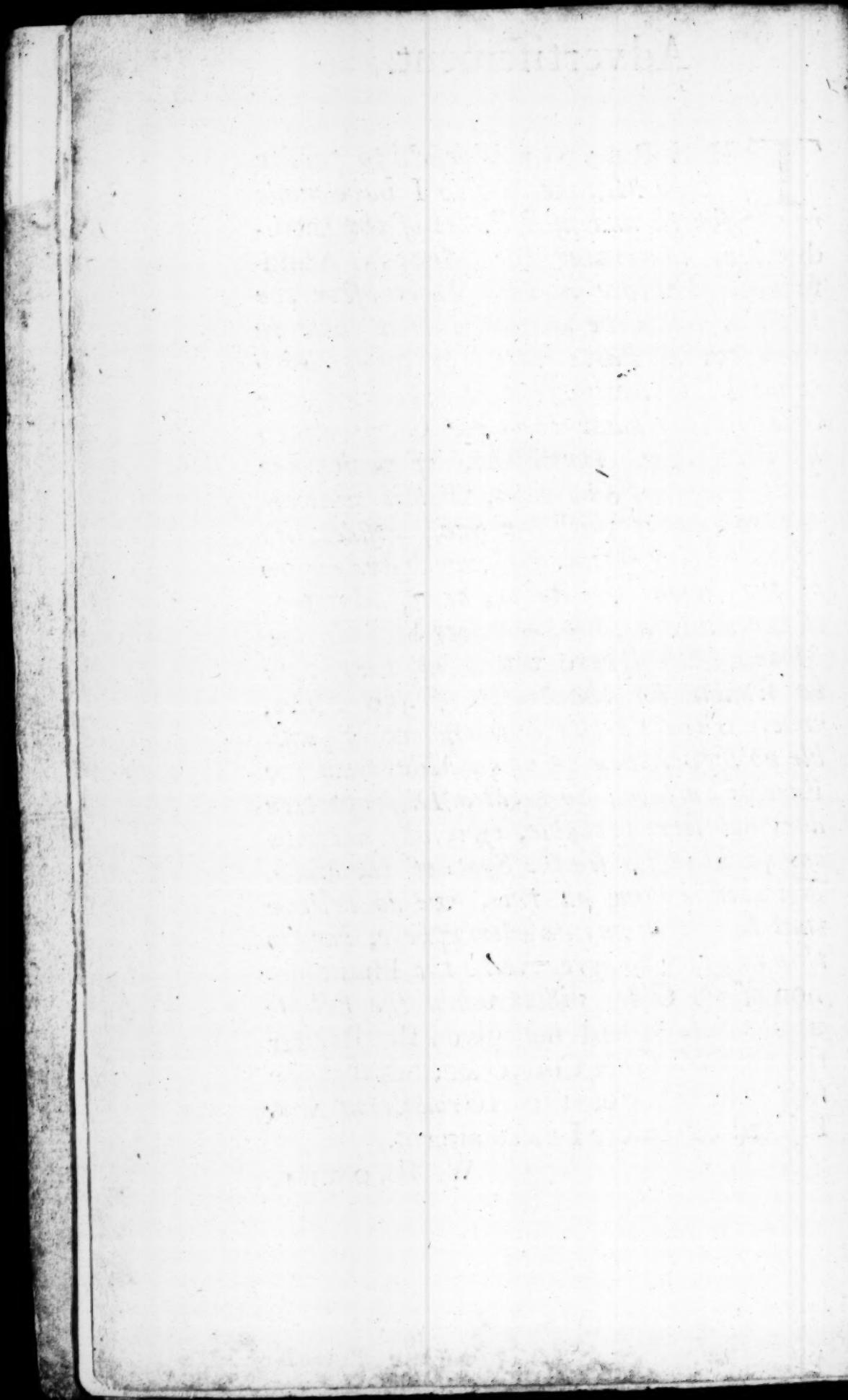


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THE READER is desir'd to pardon some Mistakes which I have made in the fourth and fifth Pages of the Introduction, concerning the Moons's Atmosphere. I wrote it two Days after the Eclipse, when the Accounts that came in from several Parts seem'd to agree with what I laid down; but it appears by the Observations made upon the Glory which appear'd about the Moon, of which we have since had Accounts, that it was really concentric with the Sun; besides, its breadth being between a seventh and eighth of the Moons Diameter, or $7\frac{1}{2}$ Minutes; if we allow a Mile for every Second, the Moon's Atmosphere wou'd at that rate be sensible for the height of 450 Miles, whereas the Earth's Atmosphere is sensible no higher then 45 or 50 Miles from the Earth's Surface: So great a Disproportion does not seem probable, especially since we can so distinctly see the Spots of the Moon. Notwithstanding all this, we do believe that the Moon has an Atmosphere, but too thin and low to agree with the Phænomenon of the Glory visible when the Eclipse was total. I had not given the Reader the trouble of this Correction, but that the first half Sheet of this *Introduction* was Printed off before I cou'd alter it.

W. BROWNE.





INTRODUCTION,

Shewing the *Discoveries* made by
CATOPTRICS and DIOPTRICS.



T is a great Encouragement for those who would take the Pains to perfect themselves in any Science, to be first inform'd of what Service it will be to understand it: And since there is no part of Learning of so real and general benefit to Mankind, as this of *Catoptrics* and *Dioptrics*, it is but Justice both to the Subject and Reader, to give some account of the many wonderful Discoveries which we owe entirely to this Science.

The *Sight* of Man is of it self confined to very narrow Views, and though it takes in a great part of the Creation at once, yet

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all is represented in Miniature and imperfectly. The naked Eye sees only so much of external *Objects* as is sufficient to move the principal Passions, and give notice of what more immediately concerns the Safety and Happiness of the Animal. What is more than this, was left as a Subject for our Curiosity, upon which we might exercise those Faculties which are bestow'd by our bountiful Creator for this very End, of searching into the astonishing Mechanism of all his Works, and from thence enlarging our Idea of his Greatness. *Objects* plac'd at a great distance, whether upon the Surface of our Earth, or in the Heavens, are seen under so small an Angle, that their Parts are not to be distinguish'd one from another; and by this means those distant charming Scenes of Nature were hid from us, which the Study of *Catoptrics* and *Dioptricks* has since laid open to our View. This noble part of Knowledge teaches us how by a due Position of *Glasses* ground into certain Figures, we may enlarge the Diameters of the Heavenly Bodies, and all such *Objects* to which we are allow'd no nearer approach, in what proportion we please, and view them as perfectly and distinctly as if we could summon them before us, and command them to the End of our *Telescope*. This has brought us into

INTRODUCTION. vii

into a perfect Acquaintance with those surprizing parts of the Creation, which are far separate from this Globe of ours, and with which we are allowed no Commerce but looking. We can now perceive the *Sun* to be a vast Globe of Fire, and by the different *Phases* of all the *Planets*, that He is the Fountain of all their *Light*. The Surfaces of most of them appear like so many *Maps* of Land and Water, and there are few now but allow both them and the *fix'd Stars* some nobler use than to twinkle upon us o' Nights. By fixing upon some remarkable *Spots* upon their Surfaces, and observing how they shift their Position, and in what time they again return to the same place, we determine the Motion of these Bodies round their *Axes*, and the Time in which that Revolution is perform'd. Several *secondary Planets* or *Satellits*, which were too small for the naked Eye, are now discern'd to move round *Jupiter* and *Saturn*, as the *Moon* round our *Earth*; and about the last of them is seen the particular *Phanomenon* of an *Annulus* or *Ring*. Nor is the Discovery of these *Satellits* meerly speculative, but of prodigious use and advantage; for their *Eclipses* have determin'd the *Velocity of Light*, as *Romer* has shewn, and are so frequent as to be the most constant appearance

the Heavens afford us at present, for the Solution of the great and valuable Problem of the *Longitude*. The Distances, Magnitudes, and Motions of all the Heavenly Bodies, and even the Irregularities of the *Moon*, have by this means been so nicely observ'd, and by the Power of Numbers reduc'd within some few Tables for common and easy Use, that their Places for any determinate Instant of Time to come are now to be predicted as easily, and almost as exactly as we could wish.

The nice Prediction Dr. *Halley* gave of the late *Solar Eclipse*, total at *London*, a rare sight in our part of the *Globe*, is such an Instance of the great Perfection to which we are arriv'd in these Matters, as has amaz'd those unthinking Gentlemen, who were only to be rowz'd out of their Security in Ignorance by the apprehensions of *Doomsday*. And here, in a particular manner, let me congratulate the learned World upon the long look'd for Discovery of the *Lunar Atmosphere*, which the *Telescopic Observations* upon this noble *Phænomenon* have at last demonstrated : For besides the remarkable Dimness of the *Sun's* Light as the *Moon* was coming on, in comparison with that which was observ'd as she went off ; the Reason of which might very well be, that the preceding part of the *Moon's Atmosphere*

mosphere which first came over the *Sun*, having enjoy'd a *Lunar* Day of 15 of our Days, must be full of Vapours rais'd to a great Height by the constant Action of the *Solar* Heat, and so give a considerable Interruption to his Rays in their passage to us, whilst that which last left the *Sun's* Disc having enjoy'd a *Lunar* Night of the same length, must have its Vapours condensed and sunk upon the Surface of the *Moon*, and so become clear and pellucid ; besides this, the *Ring of Light*, which was visible round the *Moon*, during the total Darkness, was observ'd to be concentrical with the *Moon* her self, even by those who saw the total Darkness little more than momentary, which is a plain Demonstration of a *Lunar Atmosphere*. The prodigious Distances of the fix'd Stars beyond that of any of our Planets, is, besides their little or no *Annual Parallax*, plainly deducible from the *Telescope* ; for the longest that ever was made, and which perhaps represents their *apparent Diameters* one or two hundred times larger than the Truth, has been so far from magnifying them, that by cutting off those irregular Rays which hinder us from distinguishing the true Termination of their Orbs, it makes them appear something lessen'd ; besides that Mr. *Huygens* has given a Method of eyen computing these

these Distances by means of the *Telescope*.

It is now reckon'd no absurd Notion to conceive these *fix'd Stars* as so many *Suns*, probably at as great distances one from another as they are from us, and every one their *System* of inhabited *Planets* circling round it: And perhaps the number of those which we see, counted by *Hevelius* to be 1888, may bear little or no proportion to those others that may be dispers'd thro' the vast Regions of the *Universe*, at such distances from our little *Ball*, that no assistance can ever help us to a sight of them. A Notion that gives, surely the most just and noble Sentiments that the Mind of Man can entertain of an Almighty Author! That the *milky Way* in the Heavens, which we behold in a clear Star-light Night, is nothing else but a continued Cluster of such *fix'd Stars*, is a Truth of which we are assur'd by the *Telescope*. And to the same help it is we owe all we know of those Heavenly Bodies call'd *Comets*; their Distance, Magnitude, and Motion round the *Sun* in such *Eccentrical Orbits*, as come some of them very near to right Lines. To what a surprizing height this *Cometical Astronomy* has been carried by the present Age, notwithstanding the Observations we have been able to make upon these Bodies are so few, and

and those made by our Predecessors so imperfect, may be seen in the Writings of those incomparable Astronomers, Sir *Isaac Newton*, Dr. *Gregory*, and Dr. *Halley*. That in every clear Morning and Evening we see the *Sun* for some time before he rises, and after he sets, is a *Paradox* only to be unriddled by *Dioptrics*; and if we would know the true place of any Heavenly Body elevated not many Degrees above the *Horizon*, the same Science tells us that here seeing is not believing, but that we must correct our *Eyesight* by a *Table of Refractions*. It is true the *Ratio* of Refraction of the *Atmosphere* very near the *Horizon* does not observe a constant Rule, because there happens a very great variety in the Accumulation of Vapours about those Parts: But then this Variation depends pretty regularly upon the Position of the *Sun* above or below the *Horizon*, and the different State of the Weather; and if in the Morning or Evening we see the lower parts of a distant Tower or Mountain thro' a *Telescope* fix'd in Position, we shall find the upper parts of the same Tower or Mountain in the same place, if our Observations be made nearer Noon, and just at Noon the same Object will be seen lowest of all, as the accurate Mr. *Huygens* has observ'd; and this difference is greater in cold and moist than

in hot and dry Weather, and tho' not in a proportion always certain, yet constant enough for Physical Purposes. The *Crepusculum* or *Twilight* is determin'd from the Rays of the Sun below the *Horizon*, first refracted at their entrance into the Earth's Atmosphere, and then reflected from that part of it near our *Horizon*, or rather from the contiguous Surface of the *Aether*, as from a *Concave Speculum* : And the height of the *Atmosphere* has been attempted from this Theorem by *Varenius*, but the Air being a *Medium* of different Density, and consequently of different Refraction at different Distances from the Earth, refracts the Rays of the *Sun* into *Curves*, and makes that Solution less exact. In short, without the assistance of *Telescopes*, *Astronomy* could have come to nothing, and our Observations on the Heavens had gone little further than foretelling a fine Morning from the setting of the *Sun*, or a Shower of Rain from the course of the Clouds. These Instances are sufficient to shew that all the noble Discoveries of the Heavens, of which the present Age may so justly boast, are deriv'd from the Knowledge of *Catoptrics* and *Dioptrics* ; and whatever Improvements are hereafter to be made, can be expected from no other Fountain.

I shall now descend to a Prospect no less amazing, which the same Science opens to

us in the minute parts of the Creation. The Difficulty which hinder'd the naked *Eye* from examining the smallest Particles and subtle Texture of those Bodies that are always under our Command, was, that when such *Objects* are brought near enough the *Eye* to have their least Parts subtend a sensible Angle, they become without the limits of *distinct Vision*. For as long as the *Pupil* of the *Eye* can, by the Circular Fibres of the *Uvea*, be contracted in proportion as the *Object* is brought nearer, the *Cones* of *Rays* from each *Point* may still be look'd upon as *Cylinders*, and will consequently be brought to a Point in the *Focus* of the *Eye*, which is at the *Retina*, and still make *distinct Vision*: But this Contraction of the *Aperture* of the *Pupil* holding no nearer than about four Inches from the *Eye*, if the *Object* is brought nearer than this, the increas'd Magnitude is of no farther service, because the *Rays* from each *Point* must be now consider'd as *diverging*, and will consequently after refraction at the *Eye* be made to converge to a *Focus* beyond the *Retina*, or *Focus of parallel Rays*, and so make *confused Vision*; and the nearer the *Object* approaches, the farther is its *Image* projected beyond the *Focus* of the *Eye*, and becomes so much the more confused. *Dioptrics* teaches us to remedy this Inconveniency

two ways; the first is by looking thro' a hole prick'd in a thin *Plate*, suppose of *Lead*, whose *Aperture* must be so much the smaller as the *Object* is nearer, for this supplies the place of a farther Contraction of the *Pupil*: But because this lessening the *Aperture*, excludes a great many *Rays* from each *Point*, and so diminishes the brightness of the *Image*, and that in a *duplicate Ratio* of the *Diameter* of the less'd *Aperture*, the same Science has also pointed out to us the more curious Invention of the *Microscope*. By means of this, we discern the admirable Range of the *constituent Particles* of all such Bodies as come within our nearer view and acquaintance. The *Cuticula* or outward Skin of the Humane Body is found to be composed of several *Strata* of *Scales*, lying one over another in different Numbers, according to its different thickness in different places: Between these *Scales* the *miliary Glands* dispersed over the Surface of the whole Body, are seen to send out their *Excretory Ducts*, thro' which we perspire; and about one of these *Scales* the *Microscope* reckons near 500 such *Ducts*, and that one Grain of Sand will cover 250 such *Scales*; so that one Grain of Sand will cover 125000 *Orifices* of these *Excretory Ducts*. A Discovery that must make us bless our selves, and stand

stand astonish'd at the *Infinity* of the *Creator*, when the Creature is so much beyond our *Comprehension*! The inquisitive Mr. *Lewenhoek* has oblig'd the World with a prodigious Number of such surprizing Truths, which the curious Reader will find among his Writings. The extream Ductility and Minuteness of the Particles of *Gold* is no less wonderful; for a piece of Silver gilt with *Leaf Gold*, and drawn into the finest gilt Wire, whose Diameter is $\frac{1}{386}$ of an Inch, and the thickness of the Skin of *Gold*, (as Dr. *Halley* has, from the Specifick Gravities of the two Metals, computed it) not above $\frac{1}{134500}$ of an Inch, discovers not the least Particle of Silver through the Pores of this Skin of *Gold*, tho' view'd by the *Microscope*. The Particles of the Dust which flies like Smoke out of the *Fungus Pulverulentus*, or *Puff-Ball* when burst, are discern'd to be perfect Spherules of an Orange Colour, something transparent, and their Diameters not above $\frac{1}{10}$ of that of an Hair; so that a Cube of an Hairs breadth would contain 125000 such Spherules. The *Circulation of the Blood*, that noble Discovery of our Immortal *Harvey*, is now made visible in the transparent parts of Animals, such as the Fins and Tails of Fishes, and the Feet of Frogs; and the *Anastomosis* of the Arteries and Veins put out of

[b 2]

Question.

Question. It is no less instructive than curious to behold the different Organization of the lesser *Species* of Animals, as the regular Armour of the *Flea*, the jagged Proboscis of the *Tick*, and the Bristles of the *Mite*; and in a *Louse* as he he stirs his Legs, you see the Motion of the *Muscles* of his Body, whose *Tendons* seem all to be united in a longish dark Spot in the middle of his Breast, and the like Motion is observable in the *Muscles* of the several Articulations of the Legs, and in those of the Head as he stirs his Horns, there also appears a great variety of Branchings of Blood-Vessels, and the pulse regularly beating in several Arteries, and even the *Peristaltick Motion* of the *Intestines*, continued from the Stomach all the way to the *Anus*, which is also to be seen in the *Flea*, and several sorts of transparent *Maggots* and *Caterpillars*. Besides these, the *Microscope* has presented us with an infinite variety of little *Animals*, with which the naked Eye can have no Acquaintance. They are observable in different Shapes and Sizes about the green Weeds growing in Water, in several Aromatick Infusions, and in the standing Water in the Hollow of the *Cabbage* and *Teazle*, but in such Numbers in that which drains from an Horse Dung-Hill, that they appear sometimes as thick as Bees in a Swarm, or Ants on an *Hillock*, and must be diluted with fair Water

Water to separate their different *Species*. The *Animalcula* in the *Semen virile* are of all, the Subjects most worthy our Notice and Admiration, because from this little shapeless Creature, we have reason to believe that the glorious Frame of Man himself arises; and this the rather, because in the Seeds of Plants and Trees the *Microscope* discovers the future Plant and Tree already form'd, and the *Semen Masculinum* of other Animals, as Bucks, &c. are found to be furnish'd with its *Animalcula*: Where it is to be remark'd, that sometimes the Viscidity of the *Semen* hinders the Success of Observations of this kind, and must in such Cases be diluted with a little warm Water. This Theory of Generation is handsomely and at large explain'd in the *Philos. Transactions* by Dr. *Garden*. The use of *Microscopes* has found that loathsome catching Distemper the *Itch*, to be occasion'd by the Depredations made upon the Skin by a certain *Species* of voracious *Animalcula*, which are describ'd in the *Phil. Trans.* by a Foreigner, in a Letter to Dr. *Mead*; and indeed seems to promise the finishing Hand to the Science of *Medicine*: For if we can once by a sufficient Number of Experiments, determine the different change of the Texture of the Blood in every different Distemper from that which it enjoys in its natural

natural and healthful State, and by mixing the smallest Particles of several Medicines with it, find out those which will again reduce it to that natural State, there seems to be nothing more wanting to the *practic Part*; and if the true *Mechanic Theory* of all these different Changes be ever to be known whilst we live in this Cloud of Flesh, I'm sure we must have the *Data* for it from the *Microscope*. The Method of estimating the Magnitude of *Microscopical Objects* seen by a single *Lens* only, being so easy that any one who knows ever so little of *Plane Trigonometry* will easily hit of it himself, is not mention'd in this Book; besides that it is already given by Dr. *Keill* in his *Physical Lectures*; where he shews that an *Animalculum* placed at the the Distance of $\frac{1}{2}$ of an Inch before a single *Lens*, and seen thro' it under an Angle of one Minute in length, is nearly $\frac{3}{1000}$ of an Inch long, and if its Figure were Cubical, the Magnitude of it would be $\frac{27}{1000000}$ of a Cubick Inch. From whence he concludes with a great deal of Reason, that what some Philosophical People dream of Angels, may very well be applied to these *Animalcula*, that when they have a mind to be merry, several thousand Couples of them may lead up a Country Dance upon the Point of a Needle.

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I am sensible I need say no more in Recommendation of the Subject: And for so much of the Book as is Dr. *Gregory's*, the very Name of the Man gives it sufficient Reputation. But I am conscious that part which I have attempted to add, stands in need of some Name to recommend it with which the World is much better acquainted than with mine; and for that reason I have obtain'd the Favour of making use of those of Mr. *Jones* and Mr. *Desaguliers*. Men against whose Judgment in these Matters their Approbation of the following Papers is the only possible Objection; and whose Names can never fail of meeting with that Esteem which they deserve, when fix'd to any thing of their own, however they may happen to be treated for appearing in this Place to recommend what is mine.

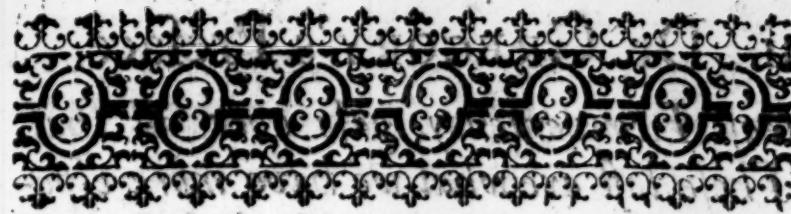
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1715.

W. BROWNE.

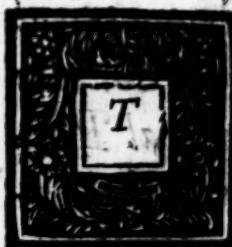


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Dr. G R E G O R Y's
P R E F A C E.



HESE Elements of Catoptrics and Dioptrics, which were Eleven Years ago read publickly in Lectures, in the University of Edinborough; I have composed for the Use of Young Students; in such a manner, that nothing but EUCLID's Geometry is required towards the understanding them: For tho' I have likewise demonstrated from higher Principles, why Spheres and Conoids observe the same Laws; both in reflecting and refracting Light; yet those who are sollicitous only about the Properties of Plane and Spherical Surfaces; may, without the

ii Dr. GREGORY's Preface.

least Inconveniency, pass over all that. These last are what we have more especially consider'd, as also such Optical Instruments as are made by a Combination of them; that is, whose Effects arise either from a single Lens or Speculum, or from severall combined together. I have, as KEPLER did before me, made use of some Postulata, that come not quite up to Geometrical Strictness, but yet are of great Service in resolving Questions in Natural Philosophy, which wou'd otherwise be extremely intricate. If these Elements be found capable of instructing such as are less conversant in Optics, I shall have my End.

THE

THE ELEMENTS OF CATOPTRICS AND DIOPTRICS.

INTRODUCTION.

THOSE Questions concerning the Nature of Lucid Bodies, and of Light, which usually cost Philosophical Writers so much Pains and Trouble, we have, after the Example of Mathematicians, omitted. For if they, who by their Inventions, have so much improv'd this Science, had employ'd all their Time in enquiring into the absolute Nature of its Object, and

iv *INTRODUCTION.*

the most hidden Causes of its *Phænomena*, not contented with deducing after a Geometrical Manner from those more simple and easily observ'd Properties of Light, others less obvious; Optics had fallen much short of that Perfection to which they are now arriv'd. Therefore whether Light be the Action of the Lucid Body driving on those Bodies that lie next it, which likewise drive on others next to them, and so on of the rest, none of them in the mean time singly moving any considerable Space; or whether it consists, which is much more likely, of Corpuscles projected with a very great Velocity from the Lucid Body thro' the circumambient Spices; or whether it be of a quite different Nature, and such as may hereafter, or perhaps never, be perfectly discover'd; yet we may easily be allow'd to assume this Property of it, which is simple enough, and confirm'd by Experiments, That from every Lucid Point, Rays are every way propagated in an Orb, and, in a Medium that is homogeneal, are diffused in right Lines (such being the *shortest*) after the same uniform Tenour.

But if those Rays meet with a Medium differently affected, whose Parts either strike them back, or diffuse them more or less, than the Parts of the former Medium did,

INTRODUCTION. v

did, they will then suffer an *Inflection*, by which general Name, I wou'd, with other Authors, understand their Reflection, as well as Refraction. For Light striking upon a Surface, that absolutely denies it Entrance, but yet hinders not its being diffused after the same manner as before, will all of it return back the easiest way it can find, diffusing it self still as at first; this is called *Reflection* of Light, and the Science which treats of the Laws it observes according to the different Incidence of Rays upon Bodies of different Figures, is call'd *Catoptrics*. But if the Medium, upon which the Light strikes, allows indeed a Passage to its Rays, but then so as that they must be either more or less diffused than before, every Ray will be inflected from the right Line, in which it was before disposed to proceed, and this Inflection is call'd *Refraction*; and the Science which demonstrates the Laws and Effects of it, is call'd *Dioptrics*.

The *Radiant*, is that from every Point of which Rays are propagated.

Parallel Rays, are such as are equidistant from one another.

Diverging Rays, are such as, if produced both ways, meet on the side contrary

vi INTRODUCTION.

to that towards which they move.

Converging Rays, are such as, if produced, meet on the same side towards which they move.

It must be observ'd, that this *Parallelism*, *Divergency* and *Convergency*, is to be understood of Rays proceeding all from the same Point.

The *Focus*, is that Point, in which Rays proceeding from the same Point of the Radiant, being produced, do meet; whence the *Focus of Parallel Rays* is look'd upon as infinitely distant.

The *Angle of Incidence*, is that, which is made by the *Incident Ray*, and a Right Line perpendicular to the *Inflecting Surface* at the *Point of Incidence*.

The *Angle of Reflection*, is that, which is made by the *Reflected Ray*, and the same *Perpendicular*.

The *Angle of Refraction*, is that, which is made by the *Refracted Ray*, and the same *Perpendicular* produced.

The two following Propositions we have assumed for *Axioms*, because they agree

INTRODUCTION.

gree both with Geometrical Reasoning and Experiments.

AXIOMS.

1. A Ray of Light falling perpendicularly upon an Inflecting Surface, either proceeds directly forward, or is reflected back upon it self. For since the Direction of the Ray to the Inflecting Surface, is, of all that can be drawn from the Radiant Point, either the *least*, if the Inflecting Surface be a Plane, or perhaps, where this Circumstance is wanting, the *greatest*, and in both Cases a determinate, and *only one* of its kind; the Ray will still persist in the same Direction, either proceeding forward, or returning backward. For there are innumerable Right Lines inclin'd to this *only one* in any given Angle, no one of which can consequently claim to its self the Direction of the Ray with greater Justice than the rest.

2. If a Plane be supposed, produced thro' the Incident Ray, and a Right Line perpendicular to the Inflecting Surface at the Point of Incidence, the Inflected Ray will likewise be found in the same Plane: or, which is the same thing, every Inflection is made in a Surface that is perpendicular to the Inflecting Surface. And

iii INTRODUCTION.

this Surface shall, according to ALHAZEN, be call'd the *Plane of Inflection*. For since this Plane is, either the *least* or *greatest*, of all that can be produced thro' the Radiant Point, to the inflecting Surface, and consequently an *only one*, the Proposition is demonstrated after the same manner as the former: and indeed if we more closely consider it, we shall find the former to be only a particular Case of this latter.

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THE ELEMENTS OF CATOPTRICS.

PROP. I. THEOR. I.

I F a Ray of Light be reflected from a plane Surface, the Angle of Reflection will be equal to the Angle of Incidence. [Plate I. Fig 1.]

Tho' the Demonstration of this Theorem belongs more properly to Physicks, and a Mathematician might very well take it for granted, as sufficiently prov'd by Experiments; yet in Compliance to Custom, we shall here give a Demonstration of it, and that such an one, as shall have no Dependance upon any Party of Philosophers whatever.

Let

10 The Elements of Catoptrics.

Let A B signify a Ray proceeding from the Radiant Point A, and falling upon the Plane reflecting Surface, G C at B. I say its reflected Ray B E will be such, that the Angle of Reflection P B E (B P being supposed perpendicular to G C) will be equal to the Angle of Incidence P B A.

From the Radiant Point A, let fall A C perpendicular to the Surface G C. All the Rays comprehended between A B and A C, and which possess, in the reflecting Plane, the Right Line B C (the common Section of the Plane of Inflection with the reflecting Plane) wou'd, if the Medium G C d e were passable and homogeneous with the former E B C D, be diffused in the same constant Tenour; that is, at the Distance B e beyond G C, they wou'd be diffused into d e, A B e and A C d being supposed Right Lines. And since, by Supposition, the plane reflecting Surface B C, neither increases nor diminishes the Diffusion of those Rays (for if it were otherwise, which no Experiments yet give us any reason to suspect; 'tis plain the Angle of Reflection as well as of Refraction wou'd, for the same Reasons, become unequal to the Angle of Incidence: For the due inflecting of which Rays, CARTESIUS has contriv'd such Conoids as are proper, Geom. lib. 2.) but only strikes them back upon the same Medium in which they were before;

fore; that is, upon a Medium where the Rays are dispersed or diffused after the same manner as at first; it is plain, that after they are reflected from the Surface GC , the Rays will be dispers'd after the same manner, as they woud have been, if they had never met with the Surface CG ; but, that being removed, had still remain'd in the same Medium. In which case, when the Distance is encreas'd by the Excess BE , the Rays woud be dispers'd thro' the Space ed . Wherefore 'tis plain, that at the Distance BE equal to Be , they will be dispers'd thro' the Right Line DE , equal to the Right Line de . But the Ray AC is reflected upon it self, wherefore DE is equal to de , and the Figure BCD every way similar and equal to the Figure $BCde$; and indeed is the very same reflected, or revolv'd half way about the fix'd Right Line BC . Therefore if from the equal Angles EBC , eBC , the right Angles PBC , ABC be taken away, the Angles EBP , eBp will remain equal. But ABP is equal to eBp , therefore the Angle of Incidence PBA , is equal to the Angle of Reflection PBE .

Q. E. D.

S C H O.

12 The Elements of Catoptrics.

scilicet in quodvis nequicunque recta
sit linea in SCHOLIUM.

In laying down this Law of Reflection, whereby the Angles A B P, E B P, are made equal, great Respect is had to the Maximum and Minimum. For the Sum of the Right Lines A B, B E is a Minimum; that is, less than the Sum of any others drawn from the same Points A and E, to any other Point of the Plane G C: and so vice versa. For it is requisite that the Course which the Ray (reflected from the Plane G C) takes from A to E, shou'd be the shortest of all. Because Nature ever acts by the most easy and expeditious Methods.

But if the Reflection be from a curve Surface, the aforemention'd Sum of Rays (or perhaps their Difference) is sometimes a Maximum. For Mathemati- cians know how near the Relation is be- tween a Maximum and a Minimum, be- tween the Sum and the Difference, and how easy the Transition is from one to the o- ther.

COROL. I.

Hence the Angles E B G, A B C (which are likewise by some call'd the Angles of Incidence and Reflection) are also equal: for they are the Complements of the for- mer to right Angles.

COROL.

C O R O L . 2 .

If the Reflected Ray be look'd upon as an Incident Ray, the Incident Ray will vice versa, be its Reflected Ray.

C O R O L . 3 .

If the Angles P B E, P B A, or the Angles G B E, C B A be equal, B E will be the Reflected Ray belonging to A B.

P R O P . II . P R O B L . I .

T H E Focus of Rays falling upon a plane Speculum being given, to find the Focus of those Rays after their Reflection from the Surface of the Speculum. [Plate I. Fig 2.]

From the given Focus A, draw the Right Line A C perpendicular to the plane Speculum C F, which produce to a, that C a may be equal to C A; a will be the Focus required. Let A D be an incident Ray, join a D, and produce it.

Because in the Triangles A C D, a C D the Sides A C, a C, are equal, and the Angles A C D, a C D also equal, because both right, and the Side C D is common to both, therefore the Angle A D C is equal to a D C; but a D C is equal to E D F; therefore A D C,

E

14 The Elements of Catoptrics.

E D F are equal. And consequently D E, by *Corol. 3. Prop. I.* is the Ray reflected from the plane *Speculum C F*, belonging to A D. The same may be shown of any other Ray proceeding from the Point A, and falling upon the *Speculum*. From whence 'tis plain, that the Rays diverging from A, after their Reflection from the plane *Speculum C F*, will diverge from the Point a: that is, those Rays, whose *Focus* before Reflection was A, will after their Reflection have a for their *Focus*. *Q. E. D.*

SCHOLIUM.

Since the Eye in any Position, as at O, will receive the Rays thus reflected, after the same manner, as if they really had proceeded from the Point a, is it plain the Image must appear in that Place: Because the Rays diverging from the Point a, affect the Eye after the same manner, as if that were the primary Radiant.

COROLL. IV.

From this *Prop.* and *Corol. 2. Prop. I.* it follows, that the Rays E D, O B converging towards the *Focus* a, will after their Reflection from the *Speculum C F*, converge towards the *Focus* A: and that such a Position of the Plane C F, may be easily

The Elements of Catoptrics. 15

easily assign'd, as shall make the Angle O B A equal to any given Angle; which is done by taking O B F equal to half the Complement of the given Angle to two right ones.

C O R O L. 2. In a plane *Speculum*, the Image of any radiant Point is seen in that Place where the Reflected Ray O B that passes thro' the Centre of the Eye, meets with the perpendicular A C, let fall from the radiant Point upon the *Speculum*. Whence, tho' every part of the *Speculum*, except B, were cover'd or taken away, the Image will nevertheless be visible: and if that be cover'd, and all the rest open, it will not be seen at all.

C O R O L. 3.

The Images of Objects that are inclined to the Plane of the *Speculum*, are inclin'd to the plane of the *Speculum* after the same manner as the Objects themselves; and therefore in a plane Horizontal *Speculum*, vertical Objects appear inverted.

This and what follows will be easily understood, if the Object be conceiv'd as made up of several radiant Points; and the Image of every one be attended to: For of all these together the Image of the Object consists.

C O R O L.

16 The Elements of Catoptrics.

Corol. 4. *A radiant Plane and its Image made by a plane Speculum, are equal and similar Figures, but not similarly placed. They differ as the Right Hand from the Left, or as a Figure engraven upon a Copper Plate, does from the Impression of it taken upon Paper. For they will fit if they come together.*

Class A C. set Hillside Logie
Corolla. *meets May 1st*

Because the Right Line AB is equal to AB , AO will be equal to AB and BO together; that is, the Distance of the Image from the Eye, is equal to the Incident and Reflected Ray taken both together.

Corol. 6.

Whatsoever has been said of the Image of any Object or primary Radiant, holds true also of the Image of another Image. From whence arises that Multiplication of Images, made by two or more plane *Specula* duly posited; in which it is principally to be observ'd, that the Distance of any Image from the Eye, is equal to a Ray propagated from the primary Radiant, thro' all the intermediate Reflections to the Eye.

LEMMA

to the Inflection of a Ray. **L** E M M A.

A Ray of Light is inflected by a curve Surface after the same manner, as it wou'd be by a plane Surface, touching the Curve in the Point of Incidence.

Let PQ be any curve Surface, (for here, in the Figure, as likewise in all Cases hereafter, the common Sections of the inflecting Surface, and the touching Plane, with the Plane of Inflection, are used for the inflecting Surface, and the touching Plane themselves: because every Inflection is perform'd in the Plane of Inflection, as is demonstrated in *Axiom 2.* And for the same Reason in the Room of a solid Figure, we make use of a plane one described in the Plane of Inflection; which saves us a great many Words) upon which the Ray AB falls at B . Now since the Ray is of a Thickness not considerable, the Particle, in the Curve Surface B , upon which it falls, will be extremely small. But the Inclination or Direction of the Curve Surface at B , is the same with the Inclination of the Plane DE touching it in that Point: Wherefore the Inflection which depends upon the Direction, is likewise the same, whether it be conceiv'd, as occasion'd by the Particle B in the Curve Surface PBQ , or in the

18 *The Elements of Catoptrics.*

plane one D B E. For as to the Inflection of the Ray A B, it matters not how the rest of the Surface is bent, if the Inclination of the small Particle B, upon which it falls, does but remain the same.

It wou'd be easy to demonstrate the same thing, as the Antients did, from hence, that the Angle of Contact D B P or E B Q, is less than any rectilineal one; and that a Plane may be found so inclined to the Plane D B E, that the Difference between the Inflections that are made by them both, shall be less than any given Inflection. For from hence it will follow, that the Inflection of the Ray A B, made at the Particle B of the Curve Surface, is no way different from that which is made at the Particle B of the plane Surface D B E, touching the Curve Surface in B.

PROP. III. PROB. II.

TO find the Focus of Parallel Rays falling upon a given Spherical Speculum, (or to find the Image of a vastly distant radiant Point) with respect to an Eye placed in the Axis of the Speculum, which is parallel to the incident Rays.

The Elements of Catoptrics. 19

Thro' A the Centre of the *Speculum*, draw the right Line A B parallel to the incident Rays, meeting the *Speculum* in B. Bisect A B in C. I say C is the *Focus* required.

In this, as well as in all the following Propositions, we suppose the Point D to be extreamly near the Point B. And this is necessary, in order that the reflected Ray belonging to the incident one E D, may meet with the Eye, which, by Supposition, is placed in A B, or the same produced: For the reflected Rays belonging to those that fall more remote from the Point B do, after their diverging from their *Focus*, pass beyond the Pupil of the Eye; and consequently contribute nothing towards seeing the Image. Besides, of those Rays that enter the Pupil of the Eye, they that fall most directly, or nearest the Middle of the Pupil; that is, that are reflected from those Points that are nearest to the Point B, conduce more towards seeing the Image, than those that enter the Eye near the Extremes of the Pupil: Because those that fall most directly, and close to one another, move the Sense more forcibly, than those that fall more oblique and scatter'd. For which Reasons, we need have respect only to those that fall nearest to the *Vertex* B; which holds good in all the Propositions following.

20 The Elements of Catoptrics

Let $E D$ be one of the parallel incident Rays. Draw $A D, C D$, and produce them. Because the Point D almost coincides with B , $C D$ will be nearly equal to $C B$: But by Construction, $C A$ is equal to $C B$; therefore $C A, C D$ are equal; and therefore the Angle $C A D, C D A$ are likewise equal. But the Angle $C A D$ equal to $E D A$, which is the Angle of Incidence; for the right Line $A D$ is perpendicular to the Surface of the Sphere. Whence the Angle $C D A$ is equal to the Angle of Incidence of the Ray $E D$. Therefore, by Corol. 3. Theor. I. $D C$ is the reflected Ray belonging to the incident one $E D$. Moreover, the Angles at the Vertex, $E D A, e D O$, and $C D A, N D O$ being equal, the reflected Ray $N D$, belonging to the Ray $e D$ parallel to $B A$, and falling upon the convex Surface, will, if produced backward, go to C . And what is demonstrated of any one Ray $E D$ taken at pleasure, is true of all the rest in the same Circumstances. Wherefore such Rays as are parallel to $A B$, and which when reflected conduce to Vision, if they fall upon a Concave Sphere, are, after Reflection, collected in C ; and from thence again diverging, make the Image to be seen in that Place, by an Eye placed in the *Axis*. The reflected Ray belonging

to the very *Axis* it self does also, as it were, diverge from the Point C, in the Middle between B and A: for the same thing happens to it, as to any other Ray reflected from B D, that cuts the right Line A B in a Point equally distant from the Centre A, and the Point of Incidence. But Rays that fall upon the Convex Side do, after Reflection, diverge from C, and make the Image appear in the Point C, (or which is all one, in a Point whose Distance from C is less than any given one) to an Eye placed in A B produced. *Q. E. D.*

C O R O L.

Hence, and from *Corol. 2. Theor. I.* it follows, that Rays diverging from C, and reflected from the Concave Surface, or converging towards C, and reflected from the Convex Surface, will be parallel to the right Line A B, joining the Centre of the Sphere, and the Point C.

S C H O L I U M. Fig. 5.

If with the Vertex B, the *Axis* B G, and a Parameter equal to the right Line F B, a Parabola be described, it will be the *least* of all that can be circumscribed about the given Circle F B; or the Circle F B will be the *greatest* of all that can be inscribed

22 The Elements of Catoptrics.

within that Parabola, by Corol. I. Prop. XX. lib. I. *Vincentii Viviani de Maximis & Minimis.* And this Circle and Parabola will, at the Point B, have the same Degree of Curvity, (to use KEPLER's Words, in Cap. VIII. Prop. XX. *Paralipom. in Vitellionem*) and are there most intimately united. For as the Contact of Lines is equivalent to two Intersections, and is really no way different from two Intersections infinitely near one another, as Mathematicians know very well, and for the same Reason Surfaces that mutually touch one another, have the same Power in inflecting Rays that fall upon the Place of Contact, and in producing other Physical Effects, as is shewn in the foregoing *Lemma*: So this more intimate Union, and which is equivalent to four Intersections, and consequently the most intimate that can be between the Circle and conick Section, (as the Contact between this last and a right Line is of all the most intimate) this, I say, will, in Physical Effects, that depend upon Surfaces generated by the Revolutions of these Lines, produce a farther Equipollency. For as Rays that are parallel to GB, are by the concave Parabolick Conoid BD, reflected exactly to its *Focus C*, which is distant from the Vertex B, by a fourth part of the Parameter of the generating Parabola; so the same Rays reflected

lected from the concave Sphere, are collected very nearly in the same Point: and the same Speculation holds in several other Physical Matters.

PROB. IV. PROB. III.

THE Focus of Diverging Rays being given, whose Distance from a given concave Spherical Speculum, is greater than the fourth part of its Diameter; To find the Focus of those same Rays after their Reflection from the foremention'd Speculum, with respect to an Eye placed in its Axis. Fig. 6.

Thro' A the Centre of the Sphere, and the given Focus E, draw a right Line, meeting the Spherical Surface in the Vertex B; this I call the *Axis of Radiation*; and suppose the Eye to be placed in it somewhere or other. In this right Line take the Point C such, that B C may be to C A, as B E is to E A. I say C is the *Focus* required.

Let any Ray proceeding from E, fall upon the Concave Surface of the Sphere at the Point D, near enough the Vertex B, (for we have nothing to do with those that fall more remote, because after Reflection they affect not the Eye placed in E B, as has before been shown) in which Case E D will be nearly equal to E B, and may

24 *The Elements of Catoptrics.*

in Physical Matters be taken for it. Draw AD, CD . After the same manner CD is equal to CB ; from whence by Construction CD will be to CA , as DE to EA . And inverting the Proportion, CD is to DA DE , as CA to EA . Wherefore by *Eucl. Elem. VI. 3.* the Angles ADE, ADC are equal; but ADE is the Angle of Incidence of the Ray ED , and consequently DC (by *Corol. 3. Theor. I.*) is the Ray ED reflected from the concave *Speculum*. And since the Ray ED is taken at Pleasure, it is plain the *Focus* of all the Rays diverging from E , after Reflection, will be in C , with respect to an Eye placed in the *Axis* EB . *Q. E. D.*

C O R O L.

Hence likewise will be given the *Focus* C , of Rays ED converging, in the foremention'd Circumstances, towards the given *Focus* E , and reflected from a given convex Spherical *Speculum*.

S C H O L I U M. Fig. 7.

Rays diverging from E , are reflected from the Concave Sphere BD converging to C , for this reason; because the Circle BD described on the Centre A (by whose Rotation the Sphere is generated) has the same Degree of Curvity with an *Ellipsis*, described upon any *Foci* C and E (found

as

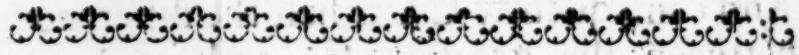
as in the *foregoing Prop.*) thro' B, and generating an oblong Spheroid, by which, as is commonly known, Rays diverging from one of the *Foci*, are from the concave Surface reflected, converging towards the other of them.

Draw the lesser *Semiaxis* H K, join the right Lines C K, E K, each of which are equal to the *Semi-axis* B H. From the Centre H upon C K, let fall the Perpendicular H L.

It is plain L K is half the *Latus Rectum* belonging to the *Axis* B G, because it is a Third proportional to the right Lines C K and H K. (Which is also true, if from H the Point of Intersection of the greater *Axis*, with the right Line K H bisecting the Angle comprehended by the right Lines drawn from the *Foci* to any Point of the Curve, a right Line be let fall perpendicular upon either of the foresaid right Lines.) And consequently CL is half the difference between the greater *Axis* and its *Latus Rectum*; and is also a Third proportional to the right Lines, K C, C H, that is, to the greater *Semi-axis*, and half the distance of the *Foci*. Again, (by the Construction in the *preceding Prop.*) B C is to C A, as B E to E A; from whence by changing and inverting the Proportion, B E will be to B C (E G) as A E to A C. Therefore B E + E G. B E - B C : : A E

26. The Elements of Catoptrics.

$E + A C$. $A E - A C$, that is, $B G$ is to $C E$, as $C E$ to $A E - A C$. And taking the halves of them, $B H$ is to $H C$, as $H C$ to $H A$; That is, $H A$ is a Third proportional to the greater *Semi-axis*, and half the Distance of the *Foci*; whence it is equal to half the difference between the greater *Axis* and its *Latus Rectum*. Wherefore $A B$ is half the *Latus Rectum*. And consequently the Circle $B D$ is (by Corol. I., Prop. XX. lib. I. Vinc. Viviani) the greatest of all that can touch the Ellipsis $B G$ on the inside at B . Wherefore (as is shown in the *Scholium* of the preceding Prop.) it will in B have the same Curvity with the foremention'd Ellipsis. From whence it comes to pass, that a Sphere thence generated, performs the same Thing in reflecting very nearly, which a Spheroid generated by the Ellipsis does exactly.



PROP. V. PROB. IV.

THE Focus of diverging Rays being given, whose Distance from a given Concave Speculum, is less than the Fourth Part of the Diameter of the Speculum; To find their Focus after Reflection from the foremention'd Speculum, with respect to an Eye placed in the Axis. Fig. 8. Thro'

Thro' A, the Centre of the Sphere, and the given *Focus* E, draw a right Line, meeting the Spherical Surface in B. In this take the Point C such, that BC may be to CA, as BE to EA. I say C is the Focus required.

Let ED be any one of the incident Rays, draw AD, CD, and produce them; draw likewise the right Line, ER, parallel to the right Line CD.

Since the Arch BD is extremely small, ED, EB, and GD, GB will be equal. Wherefore ED is to EA, as CD to CA; that is, (because DC, ER are parallel) as ER to EA; therefore the right Lines, ER, ED, and consequently the Angles, ERD, EDR are equal. But EDR is the Angle of Incidence of the Ray ED, and ERD is equal to its alternate NDA, wherefore (by Corol. 3. Theor. I.) DN is the reflected Ray belonging to the incident one ED. And since ED is taken at pleasure, it is plain that all the Rays proceeding from E, after they are reflected from the Concave Spherical Surface, will, if they be produced backwards, meet in C, or will have their *Focus* in C. Q. E. D.

CAROL.

C O R O L.

Hence, and from *Corol. 2. Theor. 1.* may be found the *Focus E*, of the Rays *N D*, converging towards a given *Focus C*, and reflected from a given *Concave spherical Surface B D*.

P R O P. VI. P R O B. V.

THE Focus of Diverging Rays being given; To find their Focus after Reflection from a given Convex spherical Speculum, with respect to an Eye placed in the Axis. Fig. 9.

Thro' *A*, the Center of the Sphere, and the given *Focus E*, draw a right Line meeting the spherical Surface in *B*: In this take the Point *C* in such manner, that *A C* may be to *C B*, as *A E* to *E B*. I say *C* is the *Focus* required.

Let *E D* be any incident Ray proceeding from *E*, draw *C D* and produce it. To this thro' *E* draw *E R* parallel, meeting the right Line *A D*, produced in *R*. The Arch *B D* being evanescent, for the Reasons already given, *C D* will be equal to *C B*,

C B, and E D to E B; and therefore A C is to C D, as A E to E D, but because of the equi-angular Triangles A C D, A E R, A C is to C D, as A E to E R. Therefore A E is to E D, as A E to E R. Wherefore E R is equal to E D. Therefore the Angle E R D, or its equal N D R, is equal to the Angle E D R, that is, to the Angle of Incidence of the Ray E D. From whence it follows, that D N is the reflected Ray of the incident E D. And since E D is taken any how, it is plain all the Rays diverging from the *Focus* E, and entring the Eye, whose Position is given, will, after their Reflection from the convex spherical Surface B D, diverge from the *Focus* C. *Q. E. D.*

C O R O L.

Hence may be found the *Focus* E, towards which the Rays N D (before Reflection towards a given *Focus* C, whose Distance from the *Speculum* is less than the Fourth Part of its Diameter,) do, after their Reflection from a given convex *Speculum*, converge.

S C H O L I U M. Fig. 10.

It had been easie to have explain'd and demonstrated these Two last Propositions and

30 The Elements of Catoptrics.

and their Corollaries, in the same Words, and with different Figures, but this wou'd have bred Confusion to Beginners. However, one *Scholium* will serve for both, to show that the Circle B D has the same degree of Curvity, with the Hyperbola described up. on the *Foci* C and E, thro' the Point B, by the Rotation of which Curve round its *Axis*, is generated the Surface of an Hyperbolick Conoid, performing exactly the proposed Reflection.

Draw the second Diameter K M. To the joined right Line B K, erect K L perpendicular, meeting the *Axis* G B in L, and H L will be a Third proportional to the right Lines, B H and H K, and consequently equal to half the *latus Rectum* belonging to the *Axis* B G. And since, from the Nature of the *Foci*, the Rectangle G C B is equal to the square of the right Line K H; adding to both the square of B H, the squares of the right Lines B K and H C will be equal: Wherefore B K, H C are equal. And because of the Rectangular Triangle B K L; B L, or the Sum of half the Transverse *Axis*, and half the *latus Rectum*, is a third proportional to the right Lines H B, B K; or to the right Lines H B, H C: That is, to half the Transverse *Axis*, and half the Distance of the *Foci*. Again (by the Construction in the Two preceding Prop.) B C is to C A; as B E

to

to E A. Wherefore B E is to E G, as A E to A C. Therefore B E -- E G. B E + E G :: A E - A C. A E + A C, that is, B G is to C E, as C E to A E + A C. And taking the halves of them, B H is to C H, as C H is to H A: That is, H A is a third proportional to half the Tranverse Axis, and half the distance of the Foci. Therefore H A is equal to B L, or to the sum of half the Tranverse Axis, and half the latus Rectum. Wherefore A B is half the latus Rectum, and consequently the Circle B D is (by Crol. 1. Prop. 20. lib. 1. Vinc. Viviani) the greatest of all that can touch the Hyperbola B G on the inside at B. Therefore, as has been shown already, it is equally curve in B with the Hyperbola. And hence it is, that the Surface of the Sphere, generated by the Circle, performs *very nearly the same Thing* in reflecting, which the Surface of the Hyperbolick Conoid does *exactly*: That is, it changes the Focus of the Rays E, into C; or the Focus C into E, as has been shown in Prop. V. and VI.

PROP. VII. PROB. VI.

THE Focus of Rays falling upon a given spherical Speculum being given; To find the Focus of the same Rays after Reflection, with respect to an Eye (even any where out of the Axis) given in Position.

Hitherto we have supposed the Eye placed in the *Axis of Radiation*, that is, in a right Line drawn through the radiant Point, and the Center of the reflecting Sphere, both because most optical Instruments are made after this manner, and because the Image seen by the Eye so placed, is much more lively and distinct than any other, because it is form'd by Rays least scatter'd, and most exactly reflected, and upon this Account challenges to it self alone the Title of an Image.

But that it may appear that the Method before used in constructing Physical Problems, and demonstrating them when constructed, is more universal, and sufficient to determine the Image, seen by the Eye, however placed: Let the reflecting Sphere be signified by its greatest Circle B D, whose Plane passes thro' the radiant Point E, and O the Center of the Eye. It is required to find the Image of the Point E
made

made by the reflected Rays, with respect to an Eye placed at O.

It is plain in the first Place, that the Image will be found somewhere in the reflected Ray passing thro' O. To find that Ray, this Problem must be solved, by which having two Points E and O, and (in the same Plane) the reflecting Circle given; it is required to find such a Point in the Circumference of the Circle, that a Ray falling upon it from either of the given Points shall be reflected to the other of them. And this is *Prop. XXXIX. Lib. V. Opticae Alhazeni*, to which he has pre-mised 7 or 8 *Lemmata*, and is now commonly called *Alhazen's Problem*. The Problem is in its Nature *solid*, and not to be constructed universally without the Intersection of a Conick Section with the given Circle. The Construction of this Problem has been publish'd by several eminent Geometers, *Barrow*, *Slusius*, &c. but most elegantly by the most noble *C. Huygens* in the *Philosophical Transactions*, No. 98. We proceed therefore to determine exactly the Point it self C, where the Image is seen in the right Line O B, (drawn as those famous Men have directed).

Produce E B, O B, till they again meet the Circle in P and R. Bisect the right Lines B R, B P in S and A, and divide S B in C so, that S C may be to C B, as

D

A E

34 *The Elements of Catoptrics.*

AE to E B. I say the Image by the Eye in O will be seen in C; and not in the meeting of the reflected Ray D N with the right Line E Q, joining the Radiant Point and the Center of the Sphere, as Euclid in *Theor. 17. and 18. Catoptrice*, and others wou'd have it.

Let the Ray E D fall upon the Point D, very near the Point B; For only those Rays, that fall after this manner do, after their Reflection, enter the Pupil of the Eye placed at O. Those that fall at a greater Distance, after Reflection, pass beside the Eye, and conduce nothing at all to Vision. Wherefore in investigating the Point C, where the Image is form'd, D must be supposed to coincide with B: In which Case $ED = EB$, $CD = CB$, $AD = AB$, and $SD = SB$. Moreover, because O B is the reflected Ray of E B, the Angles A B Q, S B Q will be equal; and the Point D approaching near to B, and at last coinciding with it, the Angles A D Q, S D Q coinciding with the former A B Q and S B Q (at least being very little different from them) will in that Case be likewise equal. But farther, the Point D approaching to B, the ultimate Angles, or the small and evanescent ones, D A B, D Q B, D S B will also be equal: For the Circle drawn thro' Q and the coinciding Points D and B; that is, described

bed on the Diameter BQ , and consequently touching the Circle $B P Q R$ on the inside, passes thro' the middle Points of the right Lines $B P$, $B R$, howsoever drawn from the Point B : And consequently the Angles DAB , DQB , DSB , in this Case, are in the same Segment of the Circle passing thro' the Points D , B , S , Q and A . These Things being laid down, which follow from the Coincidence of the Point D with B ; since, by Construction, AE is to EB , as SC to CB , AE will be to DE , as SC to CD . And since in Triangles, the Sides are as the Sines of the Angles subtended by them; the Sine of the Angle ADE , or MDA , will be to the Sine of the Angle DAB , as the Sine of the Angle CDS is to the Sine of the Angle DSB : Because therefore the Consequents are equal, the Antecedents will likewise be equal, and consequently the Angles to which they relate, namely ADM and CDS are equal. But it has been shown before, that the Angles ADQ , SDQ are equal: Therefore the Angles QDM , QDC , and those opposed to them at the vertex LDE , LDN are equal. Wherefore ND is the reflected Ray of the Incident one ED . And the same may be shewn of the other Rays, that meet with the Eye at O ; namely, that their reflected Rays will diverge from the Point C .

36 *The Elements of Catoptrics.*

Wherefore to the Eye receiving these Rays only, the Image of the radiant Point E, made by Reflection, will appear at C. *Q. E. D.*

C O R O L. 1.

Hence it follows, that Rays converging towards the Point C, will after their Reflection from a convex Spherical *Speculum*, converge towards E, and there form the Image, for an Eye placed in the right Line BE, any where beyond E. As likewise that the Image of the Radiant Point C, made by Reflection from a concave Spherical *Speculum*, will, to an Eye placed in the right Line EB produc'd, as at A, appear in E.

C O R O L. 2.

But if the Radiant Point E be vastly distant, then the *Ratio* of the right Line AE to EB becomes a *Ratio* of Equality, and by Construction SC is equal to CB: That is, the Image of the vastly distant Radiant Point situated in AB produced, made by a convex or concave *Speculum*, with respect to an Eye placed any where in BR, or the same produced, will be in the middle Point of the right Line BS.

The Eye being placed in the *Axis* of Radiation, that is, the right Lines BA and BS

BS coinciding with the right Line BQ, this same Construction serves, and degenerates into the Constructions of the four preceding Propositions.

But the spherical Surface being changed into a plane one, then this Construction will be changed into that of *Prob. I.* For because the right Lines AE, SC are in that Case infinite, but their difference remains still finite, the *Ratio* of them will be a *Ratio of Equality*: Therefore EB and BC, which are proportional to them, are equal. From whence it is plain, that this one Construction of this *Probl. VI.* contains all the former ones, as being more simple, which is frequent in Geometry.

S C H O L I U M.

The fore-mention'd Properties do therefore belong to the Circle BDR, because it has the same Curvity, at the Point B, with a conick Section described thro' B upon the *Foci* C and E, which, as is known, reflects the Rays diverging exactly from C. The Equality of Curvity in the foresaid Case, is from thence manifest: That the Segment of the Diameter of the Section (produced if need be) cut by the equicurve Circle, is equal to the *Latus Rectum* of that Diameter: And *vice versa.*

PROP. VIII. THEOR. II.

THE Image of a radiant Plane Surface, made by a Spherical Speculum, is also a Plane Surface. Fig. 12, 13.

Hitherto we have determin'd the Image of any radiant Point with respect to an Eye, placed, either in the same *Axis* of the *Speculum* with the radiant Point, or out of it. From whence it is easy to determine the Images of radiant Bodies, because made up of radiant Points. But to the more easy Determination of it, as likewise to many other Things belonging to Practice, the *Theorem* before us will not a little conduce: In which we make use of that Image of any Point, which is seen by an Eye placed in the same *Axis* with the radiant Point. For we speak not here of those Images, that are seen by the Eye in any other Position (which yet may be determined by the help of the preceding Prop.) because they are only secondary and less considerable: Especially since such an Image of a right Line has been already sufficiently consider'd by the famous Barrow, in *Leet. Opt.* XVI. and XVII.

Let any Spherical *Speculum* be signified by its greatest Circle BD, whose Center

is

is A; and the radiant Plane by the right Line F E. From A upon F E, suppose A E drawn Perpendicular, meeting the *Speculum* in the Vertex B. Find, by some one of the preceding *Prop.* applicable to the Case, the *Focus* C of those Rays after Reflection, whose *Focus* before Incidence was E; thro' this draw the Plane C T parallel to F E. I say, that the Image of the Plane F E made by the *Speculum*, will be placed in the Plane C T.

From any Point F of the radiant Plane, to A the Center of the *Speculum*, suppose the Ray F A produced, meeting the *Speculum* in D, and the Plane C T in T. The reflected Ray of the Incident one F D, is, by Construction, C D; but let B H be supposed the reflected Ray of F B, meeting the Ray F D in H.

Because the Angle F A E is, by Supposition, small, the Arch D B will be also small, and almost degenerate into a small right Line. And the Circumference of a Circle described on the Diameter B F, will pass thro' the Points D and E; because of the right Angles F E B, F D B. Whence the Angles B F D, B E D, in the same Segment, are equal. And since the Angles opposed at the vertex A are equal, or the same, F B A, E D A will be likewise equal: Therefore the Angles A B H, A D C, that are by *Prop. I.* equal to these,

40 The Elements of Catoptrics.

are likewise equal. Whence the Triangles A B H, A D C are equiangular, and consequently similar. Wherefore A B is to A H, as A D to A C; and since A B is equal to A D, A H will be equal to A C. But because of the smallness of the Angle C A T, A T is reckon'd equal to A C, therefore A H is to be reckon'd equal to A T. But H is the Image of the Point F in the radiant Plane; whence its Position is in the Plane C T. The same may after the like manner be shewn of any other Point in the Plane F E: From whence the Proposition is manifest.

But if the Object exposed to the *Speculum* be vastly distant, the Images of each of the radiant Points will, by *Prop. III.* be in the middle Points of the Semidiameters of the *Speculum* tending towards them: That is, the Image of a distant Object, will form a Spherical Surface, concentrical with the *Speculum*. But because the radiant Body, by Supposition, is seen under a very small Angle, so small a Portion of the Spherical Surface as is possess'd by its Image, scarce differs from the Plane Surface to which A B is Perpendicular. Wherefore the Proposition holds in all Cases; For the Demonstration takes Place in any other Case, as well as in those two express'd in the Figures.

COROL.

COROL.

If the Angle E A F be too great, the right Line A H will be sensibly less than A T ; whence the Image of an Object seen under too large an Angle, from the Center of the *Speculum*, made by a concave *Speculum*, will be sensibly concave : and that which is made by a convex *Speculum* will appear convex.

SCHOLIUM.

Concerning the Images of radiant Surfaces, made by Spherical Surfaces, it will not be improper to take notice of some few things. 1. If the Object expos'd to the concave *Speculum* be distant from it more than by its Semidiameter, the Image of it C T, will be distant from the *Speculum*, more than by a fourth part of its Diameter, but less than by its Semidiameter ; that is, it will appear hanging in the Air between the Object and *Speculum*, and likewise inverted ; that is, the upper Parts will appear undermost, and the right on the left. If the Object recedes from the *Speculum*, the Image will approach to the *Speculum* ; but if the Object approaches to the *Speculum*, the Image will recede from it, till meeting at last with the Object, it will coincide with it in the

42 The Elements of Catoptrics.

the Center. After the same manner the Image of the Radiant **C T**, will be **F E**, whose Properties are evident by what has been just now said. But if the Distance **B E** be so far increas'd, that the Object be vastly distant, the Distance of the Image **C T** from the *Speculum*, will be equal to a fourth part of its Diameter. But if the Sun be the radiant Body; in the room of its Image will be excited a Burning (if the *Speculum* be considerably larger than the Image of the Sun) because of the Sun's Rays being closer compacted in that Place. But if a Lucid Body be placed in the Middle, betwixt the *Vertex* and Center of the *Speculum*, its Image made by the *Speculum* at a vast Distance, will enlighten Objects vastly distant.

2. If **C B** the Distance of the Radiant **C T**, from the concave *Speculum* **B**, be less than a fourth part of its Diameter, its Image will be **E F** (by *Prop. IV.* and *VIII.*) plac'd beyond the *Speculum*, and erect as in a plane *Speculum*, and which approaches to the *Speculum*; as the Object approaches to it, and so on the contrary. After the same manner the Image of the Radiant **F E**, made by a Convex *Speculum*, will be **C T**, whose Properties, from what has been said, are easily detected: For as the Radiant approaches to the *Speculum*, the Image likewise approaches to it; but when

when that recedes to a Distance even infinite, the Image will stop at the middle Point, between the *Speculum* and its Center.

3. Suppose the Rays of each Point of the Radiant were so inflected, as to be about forming the Image C T; but are hinder'd from it by the Intervention of the convex *Speculum* B D. From what has been said above it is plain, that these Rays after Reflection from the *Speculum*, will make the Image F E, whose Properties, Situation, and Figure, with respect to the Object C T (if it may be call'd an Object) are easily found by the foregoing Rules.

In the Preceding, and all other Cases, the *Ratio* of the Image to the Object is given, because they appear under the same, or equal Angles from the Center of the *Speculum*. But this will be more plain by the following *Theorem*.

PROP. IX. THEOR. III.

THE Radiant and its Image made by the *Speculum*, are seen from the Vertex of the *Speculum* under equal Angles.

Fig. 14, 15.

From A the Centre of the *Speculum*, upon the Radiant H F, let fall the Per-

pen-

44 *The Elements of Catoptrics.*

pendicular $A E$ (which, by *Prop. VIII.* will be likewise perpendicular to the Image $f h$) meeting the *Speculum* in the *Vertex* B . Join the right Lines $B F$, $B H$, $B f$, $B h$. I say the Angles $H B F$, $h B f$, are equal. Join $F f$; this will pass thro' the Center A , because by Supposition the Image of any Point is plac'd in the *Axis* of Radiation. After the same manner the Points $H A$, and h are plac'd in the same right Line. From the Nature of the Image, $B E$ is to $E A$, as $B C$ to $C A$: therefore $B E$ is to $B C$, as $E A$ to $C A$. But because of the equiangular Triangles $A E F$, $A C f$; $E A$ is to $C A$, as $E F$ to $C f$. Wherefore $B E$ is to $E F$, as $B C$ to $C f$. And since the Angles $B E F$, $B C f$ are right, the Angles $E B F$, $C B f$ will be equal. After the same manner $E B H$, $C B h$ are equal: Whence $H B F$, $h B f$ are equal. *Q. E. D.*

COROL.

The radiant Line $F E H$, and its Image $f C h$, are to one another as their Distances from the *Vertex*, $B E$, $B C$. And if the Radiant be a Surface, it will be to its Image, in a duplicate Proportion of those same Distances. From whence, if the Distance of the Radiant, and its Image from the *Speculum* be given, the Proportion

tion of the Image to the Radiant is also given. In like manner, if the Image of the Radiant **C** made by several *Specula* be given, we shall have the Proportion of the primary Radiant to the last Image. It follows farther, that the Line **F H**, and its Image *fh*, will be cut in the same Proportion by the right Line **B E**, joining the *Vertex* and *Center*: or that **F E** is to **E H**, as *f C* to **C h**.

PROP. X. PROB. VII.

To find such a Position of the radiant Body, that its Image made by the Speculum may be equal to any given Figure similar to the Radiant: Or, which comes to the same thing, that the Radiant and its Image made by the Speculum, may be in a given Ratio.

Fig. 16.

Let the Speculum **B D** be given, whose Semidiамeter is **B A**, and let the given Ratio be **B A** to **A M**. If the Radiant be a Line, bisect **B M** in **E**; and by some of the preceding Propositions, find the *Focus* **C**, corresponding to the *Focus* **E**. I say **C** is the Place of the radiant Line. Because from the Relation of the Points **C** and **E**,

B C

46. The Elements of Catoptrics.

BC is to CA , as BE to EA ; BC will be to $BC + CA$, as BE to $BE + EA$: that is, because BE is equal to EM , BC will be to BA , as BE to AM : and inverting the Proportion, BC is to BE , as BA to AM . But (by *Cord. Prop. IX.*) BC is to BE , as the radiant Line placed in C , to which BC is perpendicular, is to its Image in E : therefore the radiant Line in C is to its Image made by the given *Speculum* BD at E , as BA to AM ; that is, in the given *Ratio*. *Q. E. D.*

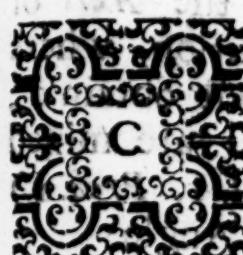
But if the Radiant be a Surface, in the room of AM in the preceding Construction, you must take a right Line, to which BA is in a *Ratio* subduplicate of that, which it bears to AM .



THE



THE ELEMENTS OF DIOPTRICS.



ATOPTRICS being finish'd, we proceed to Dioptrics, by the Assistance of which we are still furnish'd with more Instruments, and such as are fitter for Use, and admitted to a nearer View of the Secrets of Nature. For Glafs is easier brought to a due Figure than Metal, and preserves it longer: Nor does a *Lens* suffer so much Loss in its Polish from any Injuries of the Air, as a *Speculum* does. And it is difficult to make a *Speculum*, that shall reflect such strong and close Rays as a Glass *Lens* transmits.

But besides these Physical Difficulties, the Causes of which are not to be sought from Dioptrics, a *Lens* is preferable in Practice

48 The Elements of Diöptrics.

Practice to a *Speculum* for the following Reason: because a Fault in a *Speculum* produces an Error in the reflected Ray six times greater than an equal Fault in a *Lens*, when the Ray passes out of Air into Glass; and four times greater when the Ray passes out of Glass into Air. For by *Theor. I.* the Error of Incidence produces an Error twice as great in the reflected Ray, but near thrice less in a refracted Ray out of Air into Glass; and twice less in a refracted Ray out of Glass into Air, as will appear at *Schol. 2. Prop. XI.*

In Dioptrics, as before in Catoptrics, we suppose the Eye placed in the *Axis* of Radiation: because this Position, where only *Lenses* are used, is easiest obtain'd; but where *Specula* and *Lenses* are mixed together, that *Axis* (by *Corol. 1. Prob. I.*) may be inclin'd in any given Angle.

PROP. XI. THEOR. IV.

A Ray of Light, at a plane Surface of a Medium of different Density, is so refracted, that the Media remaining the same, the Sinus rectus of the Angle of Incidence will, in all Inclinations, bear the same Proportion to the Sinus rectus of the refracted Angle.

Fig. 17.
It

The Elements of Dioptrics. 49

It is sufficient for a Mathematician to have proved this *Theorem* by Experiment, so that from this laid down, he may demonstrate the Properties of a given Figure in refracting, or investigate a Figure whose Laws of refracting are given: Yet, for the same Reasons that induced us to demonstrate a like Catoptrical *Theorem*, we shall give the Demonstration of this.

For this Purpose, different Persons have used different and quite contrary *Postulata*. *Cartes* taking up the same *Theorem* (tho' in other Cases, he affirms the Propagation of Light to be instantaneous) wou'd have it in his *Dioptricks*, that a Ray of Light is carried with a greater Celerity thro' a denser *Medium*, as Water or Glass, than thro' one less dense, as Air. This appear'd too gross to *Barrow*, *Fermat*, and others, who went into a contrary and more probable Opinion. Moreover, *Barrow*, and *Maignanus* believ'd, that some certain thickness of the Ray of Light, however small (which *Cartes* neglected as inconsiderable) was necessary to demonstrate this *Theorem*. We have followed the Method of Geometers, and composed a Demonstration far enough from any Sect, and depending only upon a very simple Property of Light, which we assumed in the beginning.

E

Let

50 The Elements of Dioptrics.

Let the Light flowing from the Point A, and comprehended in the Plane of Inflection between A B and A D, fall upon the right Line B D, the common Section of the Plane of Inflection with the plane Surface of the *Medium* B C E D of different density. (We may justly consider this by it self, because the Rays which compose it, and their refracted Rays, are, distinctly from the other circumfused Light, contain'd within the fore-mention'd Plane by *Ass. 2.*) Let one of the extreme Rays A D be Perpendicular to the right Line B D, whose refracted Ray D E (whatsoever be the density of the *Medium* B E) will consequently go directly on in the same right Line with A D: But let the other A B containing a certain Angle with A D, meet the right Line B D, which separates the *Media*, in B, thro' B draw K B F parallel to A E. Let the foresaid Light contain'd betwixt the right Lines A B, A D be supposed to be propagated in the *Medium* D G, betwixt the right Lines D E and B G continued *in infinitum*: That is, let B G be the refracted Ray of A B, whether the Light passes out of a rarer *Medium* into one more dense, as in *Fig. 17.* or out of a denser one into a rarer, as in *Fig. 18.* I say, the Sine of the Angle A B K, will, in any Inclination of the Ray, have the same Proportion to the Sine of the Angle G B F.

On

On the Center B and with any distance BG, describe an Arch of a Circle meeting the Ray BG in G, from which let fall G N Perpendicular to AN, intersecting the right Line K F in F.

The Light which proceeding from A is diffused into the right Line BD, will, if BG be put for the Length of the refracted Ray, be diffused within the new *Medium* DG into the right Line GN: For the Spaces into which the foremention'd Light is diffused, must be computed by these right Lines parallel to the right Line BD, because at BD it loses its former degree of Diffusion, and after that enjoys a new one. The right Line BD which divides the *Media* remaining the same, suppose the *Medium* DC, in which AB is refracted into BG, were taken away, and another *Medium* DC of a different Density substituted in its Place, in which the Ray AB is refracted into BC, *Ex. gr.* meeting the Circle in C, thro' which draw the right Line CE parallel to BD, meeting AN in E, and KF in M. The right Line CE will be the space into which the Light is diffused within the *Medium* DC, the length of the refracted Ray BC continuing still the same. If from the right Lines GN, CE into which the Light is diffused, you take away the right Lines FN, ME which depend not at all upon the

52 The Elements of Dioptrics.

Facilities of the *Media*, that is, their *readiness in diffusing Light*, (For the length and Position of the Ray A B remaining the same, the foremention'd right Lines F N, M E remain the same and invariable, howsoever the Facilities of the *Media* D G, D C are augmented, diminished, or even annihilated.) There will remain G F, C M for the genuine Effects of the *Media* D G, D C respectively. Therefore since all Things that respect the Refraction of the Ray A B from the *Medium* D K, to *Media* of different Densities D G and D C, excepting only the Densities of those *Media*, are made the same, (namely the *Medium* D K out of which the Light passes the same, the Angle of Incidence A B K of the Ray A B the same, and the length of the refracted Ray within the second *Medium* D G, or D C the same) the foremention'd right Lines G F, C M will have the same Proportion with the Facilities of the *Media* D G, D C which produced them: For all Effects are proportional to their Causes.

Now if the *Medium* D C, in which the Ray A B is refracted into B C, be the same with the *Medium* D K, in which Case the refracted Ray B C (if it may be call'd so) of the Ray A B will be in a direct Line with the Incident one A B; G F will be yet to C M as the Facility of the *Medium* D G

DG to the Facility of the *Medium* DC, that is, by Construction, to the Facility of the *Medium* DK. Moreover CM is the Sine of the Angle CBM, or the Angle ABK, which is the Angle of Incidence of the Ray AB, and GF is, to the same *Radius*, the Sine of the Angle GBF, that is, of the refracted Angle of the same Ray: Wherefore the Sine of the Angle of Incidence is to the Sine of the Angle of Refraction, whatsoever be the Inclination of the Ray AB to the refracting Plane, (for the Angle DAB to which ABK is equal, is taken at Pleasure) as the Facility of the *Medium* DK, to the Facility of the *Medium* DG, inflecting the Ray AB into BG. But the *Media* being supposed the same, the Densities of the *Media*, and their Facilities arising from thence, and consequently the *Ratio* of these will remain the same: Therefore in every Inclination of the Ray, the *Ratio* of the Sine of the Angle of Incidence, to the Sine of the Angle of Refraction, remains the same. *Q. E. D.*

This *Theorem* is thus shortly demonstrated in the Geometrick Phrase. The Position of the Incident Ray remaining the same, the refracted Ray remains the same, whatsoever be the Length of the Incident Ray AB: Therefore no respect is to be had to its Length, or the right Line AB is to be taken for nothing. Therefore the

54 *The Elements of Dioptrics.*

right Line FN , which bears a certain Proportion to AB , vanishes at the same Time. Whence the right Line GN , into which the Light is diffused at a given distance BG within the *Medium DG*, is changed into the only and determinate right Line GF , which is consequently appointed by Nature, for the Measure of the Facility of the *Medium DG* to which it owes its rise. *Q. E. D.*

But if in the passage of a Ray from a denser *Medium* into a rarer (Fig. 18.) its Inclination be such, that GF which is to CM in the Proportion of the *Facilities* of the given *Media*, shou'd exceed the *Semi-diameter BG*; then the refracted Ray of AB will be nothing (after the manner of an impossible Case in a *Geometrical Problem.*) But that which shou'd have been the refracted Ray, will not enter the rarer *Medium DG*, but will be reflected from its Surface, according to the Law of *Theor. I.* in which Case the Law of this present *Theorem* (as being more universal) does nevertheless take place.

The converse of this Proposition is sufficiently manifest. Namely, That the refracted Ray of any Incident is truly assign'd, when the Angle of Incidence is to the Angle comprehended by the assign'd right Line, and another drawn Perpendicular to the refracting Surface, in the Proportion

portion of the *Facilities* of the given *Media*, which is the *Measure* of the *Refraction* between them: For the refracted Ray of the Incident proposed, can be no other than that assigned.

The Invention of the preceding *Theorem*, which is the principal one in *Dioptrics*, is commonly attributed to *Cartesius*, tho' it was known to *Willebrord Snellius*, who was dead Ten Years before *Cartes*'s *Dioptrics* was publish'd: For *Vossius* at pag. 36. of his Treatise, *De Natura & Proprietate Lucis*, published at *Amsterdam*, in the Year M DC LXII, tells us, That it appear'd from *Snellius*'s Papers which himself had seen, that he had found out, that the Proportion between the Secants of the Angles, which are the Complements of the Angle of Incidence, and the Angle of Refraction to right Ones, is constantly the same: But it is known that the Secants of Angles are reciprocally, as the Sines of their Complements to right Ones; Because the *Semi-diameter* is a mean Geometric Proportional between them.

Since we have happen'd to speak of Secants, it is worth the taking Notice how near the inquisitive *KEPLER* was towards finding out this *Theorem*, who at *Prop. V. and VI. Paralipom. in Vitellionem*, lays down these Secants for the respective Measure of Refractions.

56 The Elements of Dioptrics.

COROL. 1.

If BG be the refracted Ray of the Incident AB , every Thing else remaining the same, BA will be the refracted Ray of the Incident GB : For the *Facilities* of the *Media*, by which the Sines of the Angles are govern'd, remain the same.

COROL. 2.

From the Demonstration of this *Theorem* it follows, that if the *Ratio* of A to B , measure the Refraction between the *Media* A and B , and the *Ratio* of A to C the Refraction between the *Media* A and C ; the *Ratio* of B to C will be the measure of the Refraction between the *Media* B and C . Which is also manifest, by supposing the thickness of the common *Medium* A , between the parallel Planes of the other Two to vanish: For by this means, the Angles remaining the same, that part of the Ray, propagated thro' the Three *Media*, which is in the *Medium* A , vanishes.

COROL. 3.

Hence it likewise follows, that if the Position of the Ray AB remaining the same, the

the Position of the inflecting right Line BD be changed by any Angle, suppose E, the Position of the refracted Ray BG, will also be chang'd by an Angle e , which is to E, as the difference of I and R to I; (the *Ratio* of the Quantities I and R being supposed the same with the *Ratio* of the Sine of the Angle of Incidence, to the Sine of the Angle of Refraction, and the Angles so small, that they may be look'd upon to have the same Proportion with their Sines) and the Angles e and E are situated on the same Sides of the right Lines BG, and BD respectively when I exceeds R, but on contrary Sides when R exceeds I. After the same manner, from this Proposition may we judge of the change the Angle GBF will undergo, when the Angle ABK is changed by any other Cause whatever.

SCHOLIUM I.

The foresaid Law of Refraction is confirm'd from the Wisdom of Nature, always acting by the most easy and expeditious Methods; which we before found in *Catoptrics*, to be a legitimate *Axiom*. Namely, to the Light proceeding from A to G, such a Point of Incidence B is assign'd, that it may perform its Journey ABG in the *least Time* possible. Which Point

58 The Elements of Dioptrics.

Point of Incidence may be found by the help of any Method (Fermat's for Instance) which determines a *Maximum & Minimum*, but most easily after the manner following.

PROB.

TWO Points being given in Media of different Densities, and the plane Surface dividing the Media, whose Densities are likewise given, being given in Position: To find such a Point in the foresaid Surface, that a moveable Body proceeding from one of the given Points, thro' the Point sought to the other of them, may take up the least Time in its Journey. Plate II. Fig. I.

Let the given Points be A and G, and let MN be the Surface dividing the *Media* of different Densities, or rather the common Section of the plane of Inflection continued thro' A and G, with the foresaid Surface dividing the *Media*: For (by Ax. 2.) the Point sought will be found in that. Let the right Lines I and R express the Velocities of Light in the *Media* DK, CF respectively.

Because the Velocity of a moving Body being given, the Time is as the space run, and the Space being given, the Time is

as the Velocity reciprocally ; if neither being given, the Time will be in a compound Proportion of the Velocity reciprocally, and the Space directly. Therefore the Time in which the right Line A B, is run, will be express'd by $R \times A B$, and the Time in which B G is run by $I \times B G$; and consequently the Time in which the whole Journey A B G is perform'd, which is made up of both of them, will be express'd by the quantity $R \times A B + I \times B G$, which must be a *Minimum*.

From the Position of the Points A and G, and of the right Line M N being given, the right Lines A D, G C, and D C which we shall call a , c , and d respectively, are given in Magnitude. Call C B x . Whence $B D = d - x$, $A B = \sqrt{d^2 - 2dx + x^2 + a^2}$, and $B G = \sqrt{c^2 + x^2}$. Therefore $R \times \sqrt{d^2 - 2dx + x^2 + a^2} + I \times \sqrt{c^2 + x^2}$ must be a *Minimum*.

Now this is (as Mathematicians know very well) when the *Fluxion* of $R \times \sqrt{d^2 - 2dx + x^2 + a^2} + I \times \sqrt{c^2 - x^2} = 0$. But by the *Method* of that most celebrated Geometer Sir Is. Newton, for finding the Fluxions of as many Fluents as are involv'd in any given Equation, (to which also the famous Leibnit's Differential Calculus relates) which you may see at Cap. XCV. Vol. II. Operum Mathemat. Wallisi, The Fluxion

60 The Elements of Dioptrics.

xion of $\frac{R \times \sqrt{d^2 - 2dx + x^2 + a^2} + I \times \sqrt{c^2 + x^2}}{2\sqrt{d^2 - 2dx + x^2 + a^2} + 2\sqrt{c^2 + x^2}}$ is $\frac{R \times -2dx + 2xx}{2\sqrt{d^2 - 2dx + x^2 + a^2}} + \frac{I \times 2xx}{2\sqrt{c^2 + x^2}}$. Therefore making this equal to nothing, we have

$$\frac{I \times x}{\sqrt{c^2 + x^2}} = \frac{R \times d - x}{\sqrt{d^2 - 2dx + x^2 + a^2}}$$

And substituting in the room of their Values, the right Lines themselves express'd in the Scheme, it becomes $\frac{I \times CB}{BG} = \frac{R \times BD}{AB}$.

And if $BG = AB$, then $I \times CB = R \times BD$, or $BD : BC :: I : R$.

Therefore if the right Line DC be divided in B , so that DB may be to BC in the given *Ratio* of I to R , the Light will perform its Journey from A to G (or backward from G to A) in the shortest or least Time possible, by going along ABG . But DB and BC are the Sines of the Angles BAD and BGC , that is, of the Angles ABK , GBF , namely, of the Angle of Incidence, and the Angle of Refraction. Wherefore that the Time of the passage from A to G (or from G to A) may be the *least*, the Ray must so fall, that the Sine of the Angle of Incidence, may be to the Sine of the Angle of Refraction in the *Ratio* of I to R , that is in the *Ratio* of the Velocities of Light in the fore-said *Media*, or (as is shewn before in the

the Demonstration of Prop. XI.) in the *Ratio* of the *Facilities* of the same *Media*, upon which the *Velocities* of the Body moving thro' them depend.

Since there are in the Plane dividing the *Media*, an infinite number of such Points (two of which are in the right Line MN) that the Light in passing from A to G (or from G to A) thro' any of them shall take up any given Time exceeding the *least*; there can be no Reason assign'd, why it shou'd pass thro' one of those Points rather than another: Therefore it will pass thro' none but the only and determinate one of its kind B.

SCHOLIUM 2.

When a Ray of Light passes out of Air into Glass, we observe the Sine of the Angle of Incidence to be to the Sine of the Angle of Refraction, as 3 to 2 in round Numbers: And in passing out of Air into Water, I is to R (which Symbols we shall use for the future, to express the *Ratio* of the Sine of the Angle of Incidence to the Sine of the Angle of Refraction, which is the measure of Refraction) as 4 to 3. And on the contrary, in the Passage of a Ray out of Glass into Air, I will be to R, as 2 to 3, by *Corol. 1.* From whence the Reason is plain, why a Ray in Glass, striking upon a Surface of Air,

more

62 The Elements of Dioptrics.

more obliquely than in an Angle of Incidence of about 42 Degrees does not enter the Air, but is reflected from its Surface: For the right Line $G F$ which is Sesquialteral of the Sine of 42° , exceeds the *Radius*, in which Case, The Light shall be reflected as has been observ'd above.

Moreover, because in the Passage of a Ray out of Air into Glass, I is sesquialteral or $1\frac{1}{2}$ of R , and in its Passage out of Air into Water, I is $1\frac{1}{3}$ of R ; in its Passage out of Water into Glass (by Corol. 2.) I will be $1\frac{1}{3}$ of R .

And from Corol. 3. it follows, that the Error of the refracted Ray, is but subtriple of the Error of the Surface of Glass by which it was occasion'd, and on the same Side with it, when the Ray passes out of Air into Glass, but subdouble of it, and on the contrary Side, when the Ray passes out of Glass into Air.

PROP. XII. THEOR. V.

LET BD be a Surface dividing Media of different Densities, upon which let a Ray fall proceeding from E , and be refracted into DN , the Surface dividing the Media remaining the same, but the Media

dia being transposed, I say the refracted Ray DC of the Ray MD being placed in the same right Line with ED, will be in the same right Line with the former refracted Ray ND.

Fig. 2.

Draw the right Line ADO Perpendicular to the Surface BD in D.

Since DN is the refracted Ray of ED, (by the foregoing,) the Sine of the Angle EDA will be to the Sine of the Angle NDO in the *Ratio* of I to R, and if the *Media* be transposed, the *Ratio* of I to R will be the Measure of the Refraction out of the *Medium* BDM, into the *Medium* BDE: That is, supposing MD to be the incident Ray, its refracted Ray will be such, that the Sine of the Angle MDO shall be to the Sine of the Angle ADC, as I to R: That is, by what has been just now shewn, as the Sine of the Angle ADE to the Sine of the Angle NDO. Therefore since the Angles EDA, MDO are equal, and consequently their Sines, the Sines of the Angles ADC, NDO will be likewise equal, and consequently the Angles themselves. Wherefore ND and DC lie in the same right Line. *Q. E. D.*

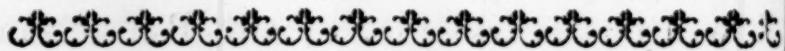
C O R O L.

If Rays diverging from E after Refraction at the Surface BD, diverge from

64 The Elements of Dioptrics.

from the Point C, the *Media* being transposed, and the Surface dividing the *Media* remaining the same, Rays converging towards E, will, after their Refraction at BD, converge towards the Point C.

Therefore all that can be demonstrated of Rays diverging, may equally be applied to converging ones. Wherefore for the future, we shall only speak of Rays diverging, leaving those converging to be determin'd by this *Theorem*.



PROP. XIII. PROB. VIII.

FROM the Focus of diverging Rays being given, to find the Focus of the same Rays after their Refraction, at a plane Surface of a Medium of different Density, with respect to an Eye placed in the Axis. Fig. 3, 4.

Thro' E the given *Focus* of diverging Rays, draw the right Line EB Perpendicular to the plane Surface BD of a *Medium*, either denser or rarer; in this take the Point C such, that CB may be to EB, as I to R. I say C is the *Focus* sought.

Let fall any Ray ED. Thro' D draw AO parallel to EB. Join CD and produce it to F.

If

If the Ray ED be near enough to the right Line EB (for we speak here only of those Rays that fall near the *Axis*, since these only do, after Refraction, affect the Eye placed, by Supposition, in the *Axis* of Radiation produced, those that fall more obliquely passing by the pupil of the Eye, and conduced nothing at all towards discerning the Point E, as we have before observ'd in *Catoptrics*) ED will be nearly equal to EB, and CD to CB. Wherefore CD is to ED, as I to R: But CD is to ED, as the Sine of the Angle BED, or its equal EDA, to the Sine of the Angle BCD, or its equal ODF: Therefore the Sine of the Angle EDA, is to the Sine of the Angle ODF, as I to R. But the Angle EDA is the Angle of Incidence of the Ray ED: Therefore ODF is the respective Angle of Refraction, that is, DF is the refracted Ray of the Incident ED. And the same may be shown of any other Ray diverging from E, whence C is the *Focus* required. *Q. E. I.*

C O R O L. I.

From hence and from *Corol. I. Prop. XI.* it follows, that the same Things being supposed as before, Rays in the *Medium* BDO converging towards C, will, after Refraction, converge towards E.

F

And

66 The Elements of Dioptrics.

And if the *Media* were transposed, it follows from hence and from *Prop. XII.* That Rays in the *Medium BDO* converging towards *E*, will, after Refraction, converge towards the Point *C*; and Parallel Rays will, after Refraction at a plane Surface of any *Medium*, still remain Parallel.

C O R O L . 2 .

If *BDA* be Water and *BDO* Air, *CB* will be to *EB* as 3 to 4: From whence it happens, that Water appears a fourth Part less deep, than it really is. But if *BDA* be Glass, *EB* will be $1\frac{1}{2}$ of the right Line *CB*.

S C H O L I U M .

What is demonstrated of any Incident Ray, namely, that its refracted Ray *DC* produced backwards, till it meets with *EB*, is to the Incident *ED*, as *I* to *R*, is also true of the perpendicular Incident *EB*: For its refracted Ray *BC* produced backwards, tho' as to its Position it be the same with the Incident, yet as to its length it is to it in the very same *Ratio* of *I* to *R*. And because this Ray passing thro' the Center of the Eye, and the others that are nearest it, are the only ones that affect the

the Sense, from thence it is, that with respect to the Eye, C is the *Focus* of the Rays diverging from E, and refracted at B D. This Consideration extends likewise to what follows.

PROP. XIV. PROB. IX.

TO find the Focus of parallel Rays falling upon a Spherical Surface of a Medium of different Density, after Refraction. Plate II. Fig. 5, 6.

Let the Spherical Surface whose Center is A, be expres'd by the Circumference of the Circle B D, thro' A draw the Axis A B parallel to the Incident Rays, meeting the Circle in the Vertex B; in which take the Point C such, that B C may be to A C, as I to R. I say C is the *Focus* required.

Let any Ray fall at D, draw the right Lines A D, C D and produce them. The Sine of the Angle B A D, or its equal E D O, is to the Sine of the Angle C D A (or its Complement to two right Ones) as C D to A C, or (the Arch D B vanishing for Reasons before often mention'd) as C B to A C, that is, by Construction, as I

68 *The Elements of Dioptrics.*

to R: But EDO is the Angle of Incidence of the Ray ED, therefore DC is its refracted Ray. After the same manner it may be demonstrated, that any other of the parallel Rays will, after Refraction, pass thro' C. Therefore C is their Focus. *Q. E. D.*

C O R O L L A R I E S.

Hence and from *Corol. I. Prop. XI.* and *Corol. Prop. XII.* it follows.

1. That Rays in Air parallel to the *Axis*, after Refraction at a *Spherico-convex* Surface of Glass, converge to a Point whose distance from the Vertex is equal to three Semi-diameters of the Sphere.
2. That Rays in Glass diverging from a Point distant from a *Spherico-concave* Surface of Air by three Semi-diameters, do, after Refraction, become parallel to a right Line drawn thro' the radiant Point, and the Center of the Sphere.
3. That Rays in Glass parallel to the *Axis*, after Refraction at a convex spherical Surface of Air, diverge from a Point, whose distance from the Vertex is equal to the Diameter of the Sphere.
4. That Rays in Air converging towards a Point distant beyond a concave Glass by the Diameter of the Sphere, after Refraction at the concave Surface of Glass,

Glaſs, become parallel to a right Line drawn thro' the Center of the Sphere, and the fore-mention'd Point.

5. That Rays in Air parallel to the *Axis*, after Refraction at a concave Surface of Glass, diverge from a Point at three Semi-diameters distance from the Vertex.

6. That Rays in Glass converging to a Point three Semi-diameters distant beyond a *Spherico-convex* Surface of Air, after Refraction, become parallel to a right Line drawn thro' the Center of the Sphere and the fore-mention'd Point.

7. That Rays in Glass parallel to the *Axis*, after Refraction at a *Spherico-concave* Surface of Air, converge to a Point distant from the Vertex, by a Diameter of the Sphere.

8. That Rays in Air diverging from a Point at a Diameters distance from a Sphere of Glass, after Refraction at a convex Surface of Glass, become parallel to a right Line drawn thro' the fore-mention'd Point, and the Center of the Sphere.

SCHOLIUM. Plate II. Fig. 7.

The generating Circle of the Sphere BD, has the same degree of Curvity with an *Ellipsis* passing thro' B, by whose Rotation a *Spheroid* being made of a denser *Medium*, collects Rays in the ambient *Medium* parallel

70 The Elements of Dioptrics.

rallel to the right Line A B, exactly at C. For we all know, that BG, the greater *Axis* of that *Ellipsis*, ought to be to EC the Distance of the *Foci*, as I to R. Supposing then the same Construction with that at *Schol. Prop. IV.* BG is to EC as KC to HC, that is as HC to CL, that is as KC+HC to HC+CL, or to HC+CK—LK, that is as BC to BC—half the *Latus Rectum*, belonging to the *Axis* BG, (for LK has, at *Schol. Prop. IV.* been shewn equal to half the fore-mention'd *Latus Rectum*) but by the Construction of this *Prop.* the Situation of the Point C is such, That BC is to AC, as I to R, that is, as BG to EC: Therefore BC is to AC, as BC to BC—half the *Latus Rectum*, belonging to the greater *Axis*: Wherefore AC is equal to BC—half the fore-mention'd *Latus Rectum*: And consequently AB is equal to half the *Latus Rectum*. Wherefore the Circle BD is equally curve in B with the *Ellipsis* BKG (by *Corol. 1. Prop. XX. lib. 1. Vinc. Viviani de Maximis & Minimis.*) Whence it is, that a Sphere generated by the fore-mention'd Circle, performs in refracting very nearly the same Thing, which an oblong *Spheroid* generated by the Rotation of an *Ellipsis* does exactly; namely, that Rays in a rarer ambient *Medium*, parallel to the right Line BG, after

after Refraction at its Surface, may converge to the Point C.

After the same manner, by the help of Schol. Prop. VI. it may be shewn, that the Circle BD is at the Vertex equally curve with an *Hyperbola* passing thro' B, and generating a Conoid, which being of a rarer *Medium*, does so refract Rays in a denser ambient *Medium* parallel to the right Line AB, that they shall afterwards diverge from the Point C: Or, which returns to the same Thing (by Prop. XII.) being of a denser *Medium*, does so refract Rays in that same denser *Medium* parallel to the right Line AB, at their entrance into the rarer ambient *Medium*, that they shall afterwards converge to the Point C.

PROP. XV. PROB. X.

THE Focus of Rays diverging and falling upon a given Spherical Surface of different Density, to find the Focus of the same Rays after Refraction. Plate II. Fig. 8, 9, 10, 11, 12, 13, 14, 15.

Let the Spherical Surface be express'd by the Circumference of its greatest Circle BD. Thro' the given Focus and the

72 The Elements of Dioptrics.

Center of the Sphere, draw the right Line EA meeting the Circumference in B, in which take the Point C such, that the *Ratio* compounded of the *Ratio* of EA to AC, and of CB to BE, may be equal to the *Ratio* of I to R. I say C is the *Focus* required.

We shall make one Demonstration serve for all the eight principal Cases of this Prop. express'd by so many several Figures, the four first of which, suppose I greater than R, the others less, and which differ from one another, according as the given *Focus* is situated on this, or that side of the *Focus* of parallel Rays; or according as the refracting Surface is Convex or Concave.

Let fall any Ray at pleasure ED proceeding from E. Join AD, DC, and thro' C draw a right Line parallel to AD, meeting the right Line ED in H.

The Point D almost coinciding with B, CD and CB, as likewise ED and EB are almost equal. Wherefore the *Ratio* of I to R, is equal to the *Ratio* of EA to AC, and the *Ratio* of CD to ED together; that is (because of AD and CH parallel) to the *Ratio* of CD to ED, and the *Ratio* of ED to DH together, or to the *Ratio* of CD to DH: But CD is to DH as the Sine of the Angle DHC, or its Complement to two right Ones ADH or

or EDO, is to the Sine of the Angle DCH, or of the Angle ADC, or in some Cases its Complement to two right Ones. Moreover, the Angle EDO is the Angle of Incidence of the Ray ED: Therefore as I to R, so the Sine of the Angle of Incidence of the Ray ED, to the Sine of the Angle ADC. Wherefore (by Prop. XI.) DC is the refracted Ray belonging to the Incident one ED; and since ED is taken at Pleasure, 'tis plain the *Focus* of all the Rays proceeding from E, will be C. *Q. E. I.*

If the *Media* proposed were Air and Glass, and the Ray pass'd out of Air, into Glass (as in the four first Figures) the *Focus* C will be very easily found: Namely, if the third Part of the right Line EA be to AB, as EC to BC. For trebling the Antecedents, EA is to AB (or AD) as $3 EC$ to BC (or DC): But EA is to AD, as the Sine of the Angle EDO to the Sine of the Angle DEA, and $3 EC$ is to DC as thrice the Sine of the Angle EDC (or HDC) to the Sine of the Angle DEA: Wherefore the Sine of the Angle EDO is triple of the Sine of the Angle HDC, and consequently (in these very small Angles which have the same *Ratio* with their Sines) $1\frac{1}{2}$ of the Sine of the Angle ADC. Therefore DC (or D lying in the same right Line with it) is

74 *The Elements of Dioptrics.*

is the refracted Ray of E D passing out of Air into Glass.

When the Ray passes out of Glass into Air (as in the four last Figures) the *Focus* will be found from the same Principles, by taking the Point C such, that half the right Line EA may be to AB, as EC to BC.

COROL. 1.

From this *Prop.* and XII. and *Corol. 1.* *Prop. XI.* it will be easy, from the given *Focus* of Rays converging, and falling upon a Spherical Surface of a *Medium* of different Density; to find their *Focus* after Refraction.

COROL. 2.

Hence from the *Foci* E and C, and the Vertex B being given, we may find A the Center of the Sphere, and thence the refracting Sphere it self; By taking the Point A such, that EA may be to AC in a *Ratio* compounded of the *Ratio's* of BE to BC and I to R. Or from the *Foci* and Center being given, may be found the Vertex. The like Problem concerning Spherical *Specula*, may (by the Correspondent *Prop.* in *Catoptrics*) be solved with equal ease, and more elegant Geo-

Geometrical Constructions from thence deduced.

COROL. 3.

The Geometrical Construction of this Prob. X. is easily deduced from what has been premised. *Plate II. Fig. 16.* Every Thing else remaining as before, thro' B and A draw the right Lines BQ, AN parallel, meeting the right Line EN drawn from E as you please in F and N: Make AN to AM as I to R. Join the right Line FM, and produce it 'till it meets the *Axis* EA in C. This will be the *Focus* required.

Join CN, meeting the right Line BF in Q; to this thro' F draw parallel the right Line FH intersecting the right Line EA in H.

For EA. EB :: (EN. EF :: EC : EH : : EC-EA. EH-EB ::) AC. BH. & *permutando* EA. AC :: EB. BH. Therefore CB. BE, + EA. AC :: (CB. EB, + EB. BH :: BC. BH :: BQ. BF :: AN. AM ::) I. R. Whence C is the *Focus* required.

This same Construction will serve (only changing the order of the Points E, B, A, C, M and N) if R exceed I, or if the Rays converge, or are parallel, in which Case it will be changed into that of *Prob.*

76 *The Elements of Dioptrics.*

IX. Or if the Concavity of the refracting Surface look towards the *Focus E*, or if *BD* be plane, in which Case it changes into that of *Prob. VIII.* For when the right Line *AMN* (divided as above) is infinitely distant from *BQ*, the right Line *FM* will be the same in Position with a right Line joining the Point *F* with a Point dividing *BE* after the same manner as *AM* is divided in *N*, supposing *BE* to be homologous to *AM*; because the Tangents of Angles are reciprocally, as the Tangents of their Complement to a right Angle.

After the same manner, if of the four Points *A*, *B*, *C*, and *E*, any other three be given, *Ex. gr.* *E*, *B*, and *C*, the fourth *A* will be determin'd; that is, the Spherical Surface passing thro' the given Point, and changing the *Focus E* into *C*, will be determin'd, & vice versa.

S C H O L I U M.

Since there is no Spheroid or Conoid generated by the Rotation of a Conick Section round its *Axis*, and consisting of a *Medium* of different Density from the ambient *Medium*, which can exactly change a given *Focus* of Rays into another given one by Refraction at a single Surface: It follows, that this Property of performing the Thing proposed pretty nearly, does therefore belong

long to a Spherical Surface; because the greatest Circle of that Sphere has the same degree of Curvity with *Cartes's Curve*, (Lib. II. Geometriae) by whose Rotation are made Surfaces, which separating *Media* of different Density, answer the Problem exactly. But if the Condition of that equicurve Circle, and of the given *Foci* lying on the same side of the Center be such, that the *Semi-diameter* of the Circle be a middle proportional between the Distances of the *Foci* from the Center of the Circle. And one of these *Distances* be to the *Semi-diameter* in a *Ratio*, measuring the Refraction between the given *Media*, that Curve of *Cartes's* is chang'd into the Circumference of a Circle. In this Situation of the *Focus* a Sphere of given Density, generated by the Rotation of a Circle, will so refract Rays proceeding from one of the *Foci*, that they shall all afterwards diverge precisely from the other of them.



PROP. XVI. PROB. XI.

ONE Focus of a given Lens, being given, to find the other.

A *Lens* is a Transparent Body, of a different

78 The Elements of Dioptrics.

ferent Density from the ambient *Medium*, and terminated by two Surfaces either Spherical, or Plane and Spherical. A right Line Perpendicular to both its Surfaces is call'd the *Axis* of the *Lens*. The Points where the *Axis* intersects each Surface are call'd the *Vertices*, either the *Vertex of Incidence*, or the *Vertex of Emergence*, according as it lies in that Surface upon which the Rays first fall, or out of which they again emerge. The *Thickness* is the distance between the *Vertices*.

The Terms being thus explain'd, the *Focus* required may, by means of a Calculation (which in Practice is to be prefer'd to the nicest Constructions) ground-ed upon the foregoing Propositions, be easily determin'd, and Canons (which be-cause of their Difficulty to be remember'd are here omitted) establish'd by making the Calculation general. As if from the *Focus* before Incidence being given, it were required to find the *Focus* after Emer-
gence: First, the *Focus* of the Rays after their Refraction at that Surface of the *Lens*, upon which they first fall must be found; and this is done by *Prop. XIII.* if the Surface of the *Lens* be plane; but if it be Spherical, and the Rays parallel by *Prop. XIV.* and by *Prop. XV.* if they be diverging or converging. And having thus got the *Focus* of the Rays after Refraction at this

this first Surface, that is, while they are passing thro' the *Lens*, which is likewise call'd the *Focus of Transition*, after the same manner their *Focus* after Refraction at the second Surface of the *Lens*, or rather at the Surface of the ambient *Medium* contiguous to this second Surface, will be found ; that is, their *Focus* after Emerson from the *Lens*. *Q. E. I.*

If there were more than one *Lens*, we must proceed after the same manner with every one of them.

By the like Method, from the *Focus* made by the help of one or more given *Lens*'s being given, the *Focus* before Incidence is found, or from the optical Machine being given, the distance of the Object is determin'd.

C O R O L.

If the Geometrical Construction of this Problem be desired, it is easily deduced from *Corol. 3.* of the foregoing *Prop.* by assuming the Construction there given for one Surface of a *Lens*, and repeating it for the other. *Fig. 17.* For the right Lines *EN*, *BF*, *AN* and *FM* being drawn according to the Directions of that *Corol.* 'Tis plain the right Line *FM* will tend towards the *Focus* of Transition of those Rays, whose *Focus* before Incidence was *E* : If therefore this

80 *The Elements of Dioptrics.*

this meets with the parallel right Lines, drawn thro' *a* the Center of the hindmost Surface, and *b* the Vertex of Emersion, in *n* and *f*, and *a n* be taken to *am*, as I to *R* at the Egress out of the *Lens*, (that is, as *R* to *I* at the Ingress into the *Lens*) the joined right Line *mf* will meet the Axis *Bb* in *e*, the *Focus* required after Refraction at both the Surfaces of the *Lens*. For the right Line *mf* is in the same Condition, with respect to the Emersion of the Rays from the *Lens*, in which *MF* was with respect to their Immersion.

In like manner, if of the Six Points *A, B, E, a, b, e*, any other five be given, the sixth may be determin'd: For Example, from the two *Foci*, the Thickness of the *Lens*, and one Surface being given, the other Surface may be discover'd.

This is the Construction, which *Barrow* receiv'd from a Friend, and placed without any Demonstration at the End of *Lect. XIV.*

PROP. XVII. THEOR. VI.

IF a plane Radiant Surface sends out Rays upon any plane or spherical refracting Surface, the Rays proceeding from every

every Point of the Radiant Surface will, after Refraction, have their respective Foci very nearly in one and the same Plane, parallel to the Radiant Plane. Plate II. Fig. 18.

Let any refracting Surface be express'd by BD , whose Center is A , and a plane Radiant Surface by the right Line EF . From A upon EF , let fall the Perpendicular AE meeting BD in B . Find C the Focus of those Rays after Refraction, whose Focus before Incidence was E , thro' which draw the Plane CT parallel to the Plane FE . I say, the Foci of Rays proceeding from every Point of the Plane EF (or the principal Image of the Plane EF , which is made up of the Images of every one of those Points, with respect to an Eye placed in the *Axis* of Radiation: For we take no Notice here of the secondary Image, seen by an Eye in any given Situation, which *Barrow* has consider'd in his three last *Lect.*) will all be posited very nearly in the Plane CT . From the Point F taken at Pleasure in the radiant Plane, to the Center A draw the right Line FA , meeting the refracting Surface at D , and the Plane CT in T . DC will be the refracted Ray of the Incident ED ; and suppose the refracted Ray of the Incident FB to be BH , meeting the right Line

82 *The Elements of Dioptrics.*

Line DT in H . Because the Angle EAF is supposed but small, the Arch BD is to be look'd upon as a right Line, and a Circle described on the Diameter BF will pass thro' the Points D and E , because of the Angles BEF , BDF being right. Whence the Angles EBF , EDF (namely the Angles of Incidence of the Rays FB and ED) contain'd in the same Segment, are equal, and consequently the Angles of Refraction AH , AC will also be equal. Therefore by reason of the Angles at A equal, the Triangles BAH , DAC are equiangular, and BA is to AH , as DA to AC , and since BA is equal to DA , AH will be likewise equal to AC : But because the Angle EAF or TAC , is very small, AT is very nearly equal to AC , and therefore AH , AT may be look'd upon as equal: That is, the *Focus* of the Radiant Point F is situated very nearly in the Plane CT . And since F is taken any how, the same will hold true of all the Points of the Plane EF , namely, that their *Foci* will be in the Plane CT . Which is demonstrated after the same manner in any other Case whatever. *Q. E. D.*

C O R O L.

Hence it follows, that the Image of the radiant Plane EF , to which the *Axix* of

of the *Lens* is Perpendicular, is also a Plane parallel to the former Plane: For the Image constituted in the Plane C T may be consider'd, as a plane Surface sending out Rays upon the second Surface of the *Lens*. But if the Angle E A F be too great, so as that A T shall much exceed A C, and can by no means be look'd upon as equal to it, in that Case 'tis easy, after the manner of Corol. Prop. VIII. to determine whether the Image of the Plane E F will be convex or concave. For Example, it will be concave towards A, if B D the refracting Surface be of a denser *Medium* and convex, or of a rarer *Medium* and concave, & vice versa.

Prop. XVIII. THEOR. VII.

THE Image appears from the Vertex of *Emersion* under an Angle equal to that, under which the Object appears from the Vertex of *Incidence*. Plate II. Fig. 19.

Let G B the *Axis* of the *Lens* be supposed produced, and standing perpendicular on the radiant Plane F E H at E, which (by Corol. of the fore-going Prop.) is also Perpendicular to its Image f Ch. Join

84 The Elements of Dioptrics.

the right Lines BF , BH , Gf , Gh . I say the Angles FBH , fGh are equal. Out of the innumerable Rays proceeding from the Point F , and after Refraction at the *Lens*, again united at f the Image of this Point, chuse two, one of which FB meets the *Lens* at the Vertex of Incidence B , and being there refracted, tends to the *Focus of Transition* of that Point, and being again refracted at L , is directed towards f : The other FD being first refracted in D , tends straight on to f , till it emerges out of the *Lens* at the Vertex of Emersion, where being again refracted, it proceeds towards f .

DGB is the Angle of Incidence of the Ray DG , and CGf its Angle of Refraction; and LBG is the Angle of Incidence of the Ray LB , which (by *Corol. 1. Theor. IV.*) wou'd be refracted into BF , and EBF is its Angle of Refraction. Because the right Lines B , G (BG the Thickness of the *Lens* being neglected) become equal, the Sines of the Angles DGB , LBG which are proportional to these, and consequently the Angles themselves DGB , LBG , and therefore their Angles of Refraction, likewise CGf , EBF will be equal. After the same manner CGh , EBH are found equal: Therefore the Angles FBH , fGh are equal, which

which is shown after the same manner in any other Case whatever. *Q. E. D.*

COROL.

Hence it follows, that a radiant Line is to its Image made by a *Lens*, as the distance of that from the Vertex of Incidence, to the distance of this from the Vertex of Emergence, or (the Thickness of the *Lens* being neglected) as their Distances from the *Lens*. But if the Radiant be a Surface, their homologous Lines will still remain in the same Proportion; but the Radiant will be to its Image in a duplicate Proportion of those Distances. Whence it will be easy to determine the Proportion which the *last Image* (which is immediately seen by the Eye) of an Object, made by the Mediation of one or more *Lens's*, bears to the Object it self.

From this *Prop.* it follows likewise, that a Radiant Line and its Image, are cut in the same Proportion by the *Axis* (of the *Lens* produced,

G 3 PROB.

THE POSITION OF A RADIANT IN THE AXIS
OF THE LENS BEING GIVEN, TO DETERMINE
ITS IMAGE MADE BY A GIVEN LENS, WITH RESPECT
TO AN EYE PLACED IN THE AXIS OF THE LENS.

PROP. XIX. PROB. XII.

Let $PCQD$ be the *Lens* proposed, in
whole *Axes* CD produced at Pleasure,
suppose the *Radiant* to be placed, to the
extreme Points of which the right Lines
 CE , CF do tend. Let b be the *Focus* of
Rays proceeding from that Point of the
Radiant which lies in the *Axes*, after Re-
fraction at both the Surfaces of the *Lens*,
found by Prop. XVI. Thro' which draw
the Plane ebf , to which CD is Perpen-
dicular. Then thro' D the right Lines Df ,
 De being drawn parallel to CF and CE ,
'tis plain (from the two foregoing Prop.)
that f is the *Image* of that Point to which
the right Line CF is directed, and e is
the *Image* of that to which CE is direct-
ed. From whence it is manifest, that the
Image fb e will be seen by an *Eye* placed
any where in CD produced beyond b , and
receiving the *Rays* from the respective Points
of the *Image* diverging.

But

But it is to be observ'd, that every Point of the Image \mathfrak{f} \mathfrak{b} \mathfrak{e} does not, like the primary Radiant, send forth Rays every way and into all Parts; but the Rays of each Point constitute a Cone opposed at the Vertex to that Cone, which has the fore-mention'd Point for its Vertex, and the Lens that refracts the Rays for its Base: Whence from the Situation of the Eye, and the Diameter of its Pupil being given, it will presently be known, whether the Eye will receive the Rays of any given Point, that is, whether it will see that Point. Which Consideration must always take Place likewise, in Vision of Images made by a Speculum.

For Example, if the Radiant be vastly distant, and the Surface $\mathbf{P} \mathbf{C} \mathbf{Q}$ plane, $\mathbf{P} \mathbf{D} \mathbf{Q}$ convex, and the Lens made of Glass with Air all around it, $\mathbf{D} \mathfrak{b}$ will (by Corol. 3. Prop. XIV.) be equal to the Diameter of the Sphere $\mathbf{P} \mathbf{D} \mathbf{Q}$: For the Rays being parallel to $\mathbf{B} \mathbf{C}$, pass unrefracted thro' the plane Surface $\mathbf{P} \mathbf{C} \mathbf{Q}$, upon which they fall perpendicularly; and the Image in respect to the Radiant is inverted.

But if $\mathbf{P} \mathbf{C} \mathbf{Q}$ be convex, and $\mathbf{P} \mathbf{D} \mathbf{Q}$ plane, $\mathbf{C} \mathfrak{b}$ will (by Corol. 1. Prop. XIII. and Corol. 1. Prop. XIV.) be equal to the Diameter of the Sphere $\mathbf{P} \mathbf{C} \mathbf{Q}$, together with a Third Part of $\mathbf{C} \mathbf{D}$ the Thickness of the Lens. And neglecting the Thick-

88 The Elements of Dioptrics.

ness of the *Lens*, as is usually done in the object *Lens* of a Telescope, the Distance of the Image of a vastly distant Radiant, from a Plano-convex *Lens*, is equal to the Diameter of the Sphere.

If the Plano-convex *Lens* *PQ* were of Water, *D b* wou'd, by *Prop. XIV.* be sesquialteral of the Diameter.

If the *Lens* *PQ* be of Glass, and both its Surfaces equally Convex, the Thickness being neglected, *D b* will (by *Corol. I. Prop. XIV.* and *Prop. XV.*) be equal to the *Semi-diameter* of either Sphere. In an entire Sphere of Glass, the Image of a very distant Radiant, will be at the distance of a Quarter of the Diameter behind the Sphere; for in this Case, the Thickness of the *Lens* cannot be neglected: In one of Water, this Distance equals the *Semi-diameter*.

In all these and the like Cases, if the Sun be that distant Radiant, and the *Lens* be notably broader than the Image of the Sun, in the Place of the Image a Burning will be excited, and that more vehement than from a Concave *Speculum*, if the Image of the Sun be equally distant from a *Lens* and *Speculum* equally broad, because of the greater loss of Rays at this, than at that. If a Lucid Body be put in the fore-mention'd Place, the Image of it will be cast at a very great Distance, and

and will enlighten those Parts that are vastly remote.

If the Radiant be not vastly distant as before, but yet more remote from the *Lens*, than the Place of the Image of a vastly distant Radiant; besides the Appearances just explain'd, namely, that the Image will be inverted, &c. if the Radiant approaches towards the *Lens*, the Image will recede, and *vice versa*; till the Radiant comes to the place of the Image of a very distant Radiant, and then its Image will become vastly distant. All which may be seen in a darken'd Chamber, receiving no Light but thro' a convex *Lens*, upon which Radiants at different Distances cast their Rays. The Place of the Image of any Radiant is known, by its being painted most distinctly upon a white unpolished Plane, situated in the Chamber. Nor is there need of subjoining any Thing more of this Experiment, which is now very common; or of that other, founded upon the same optical Principles, in which by the pellucid Colours of a Picture painted upon Glass, and transmitting the close Rays of a Flame, an Image is shown painted upon a white unpolish'd Plane.

If the Radiant be nearer the convex *Lens* than the Image of a vastly distant Radiant, then its Image will be form'd not on the opposite,

opposite, but on the same side of the *Lens*, and its Place according to what has gone before, from the Place of the Radianc being given will be determin'd. This Image is always erect, and greater than the Radianc: And as the Radianc approaches to the *Lens*, the Image likewise approaches to it, and in like manner they both recede from it at the same Time, but the Image more swiftly.

Fig. 21. If the convex *Lens* were chang'd into a concave One, the same Construction remains, and by a Calculation, (according to the *Corollaries* of *Prop. XIV.*) *abf* the Image of a very distant Radianc, made by a Plano-concave *Lens* *PQ*, will be found erect, and on the same side of *D* with the Radianc, and distant from it by the *Diameter of the Sphere PDQ*. Now if the Surface *PCQ* be also Spherical, and the Radianc from being distant becomes near, its Image will be determin'd by the same Construction, if by *Prop. XV.* the Image of the radiant Point placed in the *Axis of the Lens* be first found. In which Case, besides what has been just now said, it is also to be observ'd, that the Image approaches to, or recedes from the *Lens* at the same Time as the Radianc does, but more slowly, as will be plain to any one, who will give himself the trouble of a Calculation.

Let

Let the Rays of any Radiant, inflected after such a manner as to be ready to form the Image $e b f$ if nothing hinder'd, be conceiv'd to be intercepted by a Plano-concave *Lens* $P Q$; upon whose concave Surface $P D Q$, described on a Diameter equal to the right Line $D b$, they first fall: The same Construction as before, determines the Image made after the Refraction of these Rays at the foremention'd *Lens*. For (by Corol. 4. Prop. XV.) the Rays converging towards b , are so refracted from the concave Glass, as to become parallel to the *Axis*, that is, they constitute a vastly distant Image, to whose extreme Points (by the foregoing *Theor.*) the right Lines $E C$, $F C$ or $e D$, $f D$ are directed, and which is consequently given in Position. And in this Case, the chief Thing to be observ'd is, that to an Eye placed about B , the distant Image appears inverted, in respect to the Image that wou'd have been made at $e b f$, without the Intervention of the *Lens*. It will be easy to proceed in all other Cases, according to these Examples. Q. E. F.

PROP.

92 The Elements of Dioptrics.

PROP. XX. PROB. XIII.

To find such a Position of the Radiant with respect to a given Lens, that the Image made by the Lens may be equal to a given Figure, which is similar to the Radiant: Or, which is the same Thing, that the Radiant may be to its Image made by the Lens, in a given Proportion. Plate II. Fig. 22.

Let a *Lens* of Glass be given, for Example, let A B be the *Semi-diameter* of the first Surface, and C B the *Semi-diameter* of the second Surface. Draw C M at pleasure, making any Angle with C A. Let the Proportion of M D to D C be given, namely, that which the homologous Lines of the Radiant, and its Image bear to one another. Join A M, to which thro' B draw B E parallel, meeting the right Line C M in E; M E taken twice will be the sought Distance of the Radiant from the *Lens*. If D m be taken in the same right Line, but on the other side of the Point D, equal to the right Line D M, and join A m; B e drawn parallel to this, will cut off another m e, which will likewise satisfy

tisfy the Problem: For the Radiant placed at the distance of twice me from the *Lens*, is to its Image (but then it will be made on the contrary side of the *Lens*, with respect to the former Image) as DM to DC .

By Corol. Prop. XVIII. The Radiant Line is to its Image made by a *Lens*, as their Distances from the *Lens*: But if twice ME or twice me be the distance of the radiant Point from the *Lens*, the distance of the Image of that Point from the *Lens*, will be to the former distance respectively, as MD to DC : As will be plain to any one, who will undertake a Calculation according to Prop. XV. For if AB be call'd a , BC b , the Thickness of the *Lens* C , MD r , and z the sought distance of the Radiant from the *Lens* of Glass about which Air is circumfused; we shall have $zz =$

$$\frac{6abz - 2acz \mp 6arz \times 2crz \mp 4acr}{3a - c + 3b}$$

If the Thickness of the *Lens* be neglected (which is done in the Construction of the Problem) the sought Distance of the Radiant from the *Lens* will be equal to

$$\frac{2ab \times 2ar}{a + b}$$

In

In the Scheme referr'd to, the *Lens* is made convex on both sides; but the same Construction will serve for any *Lens*, since from the variety of *Lens's*, only the order of the Points A, B and C is chang'd. Another Construction of this Problem is easily deduc'd from *Corol. Prop. XVI.*

Prop. XXI. Prob. XIV.

WITH two given Spherical Lens's or Specula, or one Lens and one Speculum, to make an Optical Machine, which to an Eye seeing at a given Distance, shall distinctly represent a given near Object under a given Angle, the Distance of the Eye from the last Lens or Speculum being likewise assign'd.

Since the Eye is a Machine made on purpose, that the Images of exterior Radiants may be distinctly painted upon its Bottom, (which is made concave for this end, as *Corol. Theor. VI.* requires,) all the rest of its Apparatus conduced only to its Motion or Security, which are necessary to be provided for; it is plain, that a given Eye can only see distinctly at a given Distance from the Object. Now if the Eye cou'd place

it

it self at pleasure, at such a Distance from any Object as is necessary to distinct Vision, (all other Things, as the Degree of Light, &c. being as they shou'd be) Vision wou'd always be distinct. And tho' there be no Eye so stiff, as to see only at a determinate Distance, but can according to the flexibility or mobility of the Parts with which every Eye is endued, apply it self to Objects placed at different Distances, and change its Figure according to the Distance given; so as to be no longer look'd upon as the same given Eye, but various and mutable, as Occasion requires: Yet since this mobility is confined within certain Limits, and there are a great many Objects to which we cannot at Pleasure come near enough to be within those Limits; 'tis plain, there will be need of an Optical Machine to see them distinctly. But any Spherical *Lens* or *Speculum* will be sufficient for this Purpose; since by its Assistance the Image of any Radiant (to which we cannot come so near as we wou'd) may be brought near us, (as is plain from what has been before demonstrated) and then we shall be able to view it, since the Eye is supposed moveable at Pleasure with respect to any Thing near.

But because, besides distinct Vision, our Occasions sometimes require us to look into the more minute Parts of an Object; and

and it is found by Experiment, that an Object seen under a less Angle than of one Minute, is consider'd by the Observer as a Point, and its Parts not at all to be distinguish'd one from another: It often happens, that when the Object is brought nearer the Eye, that the Particles to be observ'd may be seen under a sensible Angle, and greater than the fore-mentioned one, the Object it self becomes too near the Eye, and is out of the Limits required for distinct Vision. This Inconveniency, if it be the only one, may be remedied by the Assistance of any given *Lens* or *Speculum*, by Prop. X. or XX. where the Image of an Object made by a given *Speculum* or *Lens* is represented in any given Measure.

But if both the fore-mention'd Inconveniences urge at the same Time, they are not to be removed without the help of two *Lens's* or *Specula*, or one *Lens* and *Speculum*. Having shown therefore, what Assistance the Sight may receive from a single given *Lens* or *Speculum*, we shall proceed to Machines made by two combined together; or to construct the Problem proposed universally. *Plate II. Fig. 23, 24.*

Let R be the given Object, S the given Angle in which it is to be represented, D the given Distance requisite for distinct Vision, and L the given Distance of the Eye from

from the *Lens*. Make the Triangle AOB, in which the Angle at the Vertex O is equal to the given S, and OE Perpendicular from the Vertex upon the Basis equal to the given D. If we take the middle Point of the Object placed in the *Axis* of the *Lens* or *Speculum*, which in Practice is very convenient, the Triangle AOB must be made Isosceles. Take OV equal to L. At V place either of the *Lens's* or *Specula*, having its *Axis* in the right Line OE.

By Prop. VI. or XVI. having one *Focus* E of a *Lens* or *Speculum* given, find the other e; that is, that the Rays whose *Focus* before Incidence is e, may have E for their *Focus* after Inflection at the *Lens* or *Speculum*. Thro' e draw the right Line aeb parallel to the right Line AEB, meeting the right Lines V_a, V_b drawn thro' V parallel to the right Lines VA, VB, in a and b.

'Tis plain from Prop. IX. or XIX. that if aeb be the Radiant, AEB will be its Image: Wherefore if the given near object R, and another *Lens* or *Speculum* be placed after such a manner (by Prop. X. or XX.) that the Image of the Object R made by this *Lens* or *Speculum*, may obtain the Situation and Magnitude aeb; the Microscope required is made. For aeb is the Image of the Object R, the

98 The Elements of Dioptrics.

Image of which Image, seen by the Eye placed in O, is A B B. Now this appears under the Angle A Q B equal to the given S, and at the Distance O E equal to the given D, and consequently distinct, and the Distance of the Eye from the last Lens or Speculum in V, is equal to the given E. Q. E. F.

If the given Eye be an old Man's, every Thing else remaining as before, the Right Line O E becomes infinite, and the Point e is found by Prop. III. or XIV.

If the Object proposed were vastly distant, the Problem wou'd be impossible. But if the Angle S were not given, a Telescope from the remaining data might thus be made. In the right Line given in Position, tending towards the propos'd distant Object, straight forward from O take O E, O V equal to D and L: And in V place one of the Lens's or Specula, having its Axis in V E. By Prop. VI. or XVI. one Focus E of the Lens V being given, find the other e, which being supposed the Focus of Rays before Inflection at V, E may be their Focus after Inflection. In the right Line O E let the other Lens or Speculum, having its Axis in the same, be so placed, that the Image of the distant Object made by it, may be situated in the right Line a e b perpendicular to O E, and the Telescope required is made. For the first

first Image of the distant Object is in aeb : And the Image of this Image is, by Construction, in AEB , whose distance from O is equal to the right Line $O\dot{E}$ or the given D , and consequently distinctly seen, and the Distance of the Eye from V is equal to the given L . But a distant Object being given, by Prop. XIX. its Image made by a given *Lens* or *Speculum*, and likewise the Image of that Image made by the given V , will be given, and consequently the Angle under which this last is seen by the Eye in a given Position.

But if O be an old Man's Eye, the Angle AOE , because of AB being in this Case vastly distant, is equal to AVB or Vb . Therefore the Angles under which a distant Object with and without a Machine appears, are as the Distances of the *Lens's* or *Specula* from the common *Focus*: For small Angles are almost as their *Cotangents*.

S C H O L I U M.

It will be very convenient that the *Lens* or *Speculum*, which immediately receives the Rays of the Object, and forms its first Image, (and is therefore call'd the *Object Glass*) be as perfect as possible. For the Errors or Defects of this *Lens*

180 The Elements of Dioptrics.

or *Speculum* affect the Image made by it; and since this Image acts the Part of an Object, to be seen thro' the *Speculum* or *Lens* V, (which is nearest the Eye, and therefore call'd the *Ocular Glass*) its Defects, that is, the Defects or Errors of the Object Glass by which it is form'd, will be greater and more sensible, by how much the Image AEB is greater than aeb : That is, by how much more perfect (the Object Glass remaining the same) the Machine is made. But the Errors of the Ocular *Lens* or *Speculum* V, are equally sensible, whatsoever be the Image aeb , or the Object *Lens* or *Speculum* by which it is produced: That is, the Ocular Glass V remaining the same, its Defects are equally apparent and discoverable, to whatsoever Degree of Perfection the Machine, by changing the Object *Lens* or *Speculum*, which forms the Image aeb , be brought. For the Eye being given, the given Oculars *Lens* or *Speculum* V is always at the same Distances from the Image aeb , doing the Office of an Object and its Image AEB, and consequently shows the same Defects.

PROP.

PROP. XXII. PROB. XV.

WITH Three or more given Spherical Lens's or Specula, to make a Machine, which to a given Eye shall distinctly represent a proposed distant Object under a given Angle, the Distance of the Eye from the last Lens or Speculum being assign'd.

By the help of any one of the given Lens's or Specula form the Image of the distant Object, and with the two others, by the foregoing Prop. make a Microscope that shall represent it in the Conditions proposed, and the Telescope required is made.

In like manner with these Lens's or Specula may a Microscope be made, and then by adding a fourth, another Telescope: In all which we have the Proportion which the Image seen by the Eye bears to the Object, or which the Angle under which that is seen, bears to the Angle under which this is seen without the Machine: And consequently the Powers of a Machine in promoting Vision, are by Corol. Prop. IX. and XVIII. easily estimated.

PROP. XXIII. PROB. XVI.

TO make a Spherical Lens of Glass, whose Thickness is given, that shall to an Eye seeing at a given Distance, represent a given Object placed at a given Distance, under a given Angle distinctly, the Distance of the Eye from the Lens being likewise assign'd. Plate III. Fig. 1.

We have hitherto shown what Assistance may be had from given Lens's or Specula, or both, howsoever combined, in order to supply the Defects of Vision; it remains, that we demonstrate the manner of making a Lens for given Uses.

The radiant Line AB , whose Distance from the Lens is VE , is to be represented distinctly under the visual Angle $\angle O\theta$, to an Eye seeing distinctly at the Distance Oe . In the Isosceles Triangle $\angle O\theta$, (whose base is ab , and Height eO) upon the Perpendicular Oe (so situated as to bisect the radiant Line AB perpendicularly in E) take OL equal to the given Distance of the Eye from the Lens. Draw aL, bL , and produce them to D and F . Make LV equal

qual to the given Thickness of the *Lens*. From the right Line *VL* given in Position, perpendicular to the Glass at the Points of Incidence, I draw the refracted Rays in Glass *V₂*, *V₃*, *LH*, *LG* corresponding to those in Air *AV*, *BV*, *DL*, *EL*. Produce the right Lines *HL*, *GL*, till they meet with the right Lines *V₂*, *V₃* in β and α . The right Line ϵ is joined, will cut the Perpendicular *O* perpendicularly in ϵ , because both the Angles of Incidence and of Refraction, on each side of the right Line *Oe* are equal. From the *Foci* *E* and ϵ , and the Vertex *V* being given, find (by *Corol. 2. Prop. XV.*) the Sphere of Glass *KVM* that may refract the Rays in Air diverging from *E*, so as to make them converge towards ϵ . By the fore-mention'd *Corol.* find also *NLP* the Surface of a Sphere of Air passing thro' *L*, that may refract the Rays in Glass converging towards their *Focus* ϵ , so as to make them afterwards diverge from ϵ : The solid Figure *MKNP*, being made of Glass, and terminated by the Spherical Surfaces *KVM*, *NLP* and a Conic Surface whose *Axis* is *VL*, is the *Lens* required. But care must be taken, that the Portions *VK*, *LN* be not too great, because all the preceding Demonstrations hold only true of Rays falling near the Vertex.

By Construction, the Rays diverging from E, do after Refraction at the first Surface KV M, converge to the *Focus* ϵ : Wherefore (by Prop. XVII.) the Image of the radiant right Line A B (made by Refraction at KV M) is in the right Line $\alpha\beta$ given in Position: But the Ray AV, by Construction, after it is refracted at the Surface KV M comes to the Point α ; therefore that Point α is the Image of the Point A, which is produced by the first Surface only. In like manner β is the Image of the Point B, and the right Line $\alpha\beta$ is the Image of the radiant right Line A E B. Moreover (by Construction) the Rays within the *Lens* that converge towards ϵ , after they are refracted at the Surface N L P, diverge from e ; wherefore (by Prop. XVII.) the Image of the future Image $\alpha\beta$, made by Refraction at the Surface N L P, is in the right Line aeb : But the Ray GL within the *Lens*, proceeding directly towards α , after it is refracted, becomes L F, and proceeds from α ; whence the Image of the future Image in α , after Refraction at N L P becomes α : And so likewise the Rays within the *Lens* that converge toward β , after Refraction at N L P diverge from b . That is, aeb is the Image of the Radiant A E B, placed at a given Distance V E from the *Lens*, made by the *Lens* N K M P,

N K M P, and to be seen distinctly by a given Eye in O, because it is at the required Distance O e from the Eye, and is also seen under the given Angle aOb by the Eye O, which is at the given Distance O L from the *Lens*, and V L is the given Thickness of the *Lens*. *Q. E. F.*

The same Construction serves, if the Eye of an old Man be given: For in that Case, aeb is at an infinite Distance, and the former right Lines aL , bL must be drawn thro' L parallel to the right Lines Oa , Ob given in Position, and the Center of the Surface N L P (by *Corol. 6. Prop. XIV.*) is distant from L by a third Part of the right Line L.

But if the Object proposed be a distant one, the Construction will become much more Simple: For the Center of the first Surface K V M (by *Corol. 1. Prop. XIV.*) will be distant from the Vertex V by a third Part of the right Line V.



PRO. XXIV. PROB. XVII.

TO make a concave Spherical Speculum of Glass, whose Thickness and Diameter of Concavity are given, so that the Rays

106 The Elements of Dioptrics.

Rays parallel to its Axis, refracted from both its Surfaces, shall be collected in the same Point of its Axis. *Plate III. Fig. 2.* being upon *Specula* made of Metals, neither receive a due Figure and polishing so easily, nor preserve it so long, it will be convenient to use those of Glass. Let *B* then be the Vertex of the Concave Glass Surface *EBF*, *A* its Center, and *BD* the Thickness of the Lens, being given; we are to find *DC* the Semi-diameter of the Surface *GDH*, so that the Rays parallel to *AB*, and falling upon the Concave *Speculum*, as well those that are reflected from the Surface *EBE*, as those that are twice refracted at the fore-mention'd Surface *EBF* at their Entrance into it and Emission out of it, but reflected from the Concave Surface *GDH*, may all meet in the same Point of the Axis *AB*: Or, which is the same Thing, a *Speculum* of Glass is to be made after such a manner, that the two Images of a distant Object made by the two Surfaces of the *Speculum*, may coincide, and consequently so as to be most powerful in burning, or in forming the Images of distant Objects. *Plate III. Fig. 3.*

From

LEMMA. To 107. Quidam linea etiam

From any Point L of the infinite right Line SM, on one side take LN equal to BD, and on the other side LR equal to twice BD, from N likewise directly forward take NT equal to thrice AB, and TV equal to twice NT, and VM also equal to twice RL. On the Center R, with the *Radius* RT describe a Circle meeting the Perpendicular erected at L in K: Join MK, and at K erect the Perpendicular KS meeting the infinite right Line first drawn, in S. *Fig. 2.* In the right Line DA from D towards A, take DC equal to the right Line LS; and on the Center C with the *Radius* CD describe an Arch of a Circle DG, similar to the Arch BE described on the Center A, and join EG. The solid Figure generated by the Rotation of the plane Figure BDGE about the fix'd right Line BD, and consisting of Glass, is the *Speculum* required. *Q. E. F.*

Fig. 4. But if AB be sufficiently large in respect to the Thickness BD, the Problem will admit of this more expeditious Construction. Divide BD in O, so that the greater Segment BO may be to the lesser OD, as 5 to 4; produce BA to C, so that CA may be equal to OD; and C will be

108 *The Elements of Dioptrics.*

be very nearly the Center of the outward Surface G D H. *Q. E. I.*

We shall omit the Demonstration of these two Constructions, because it is very easily deduced from the following Analytical Calculation founded upon *Prop. V. XIV. and XV.* For if $AB = a$, and $BD = c$, we shall have $DC = \frac{9aa + 18ac + 5cc}{9a + 5c}$.

In like manner may a *Speculum* be made, if the Radiant be in any other Position.

I had determined to have subjoin'd a general Calculation for finding the *Foci* of any *Speculum* or *Lens* universally: But that is abundantly done already for *Lens's* by that excellent Analyst *Edm. Halley*, in *Philos. Transact.* for the Month of *Nov. 1693.* and elegantly applied to particular Cases.

S C H O L I U M.

Hitherto we have shown what Advantages may be expected from Spherical *Lens's* or *Specula*, towards the Construction of Machines: But the different Refrangibility of the Rays of Light, and that in given Rays given, immutable, and annex'd to certain *Colours*, discover'd

ver'd by that admirable Philosopher Sir *ISAAC NEWTON*, has so much disturb'd our Dioptrical Reasonings, that no Exactness can now be hoped for from *Lens's*, even tho' form'd into what Conoidal Figures we please. But since the Law of *Catoptrics* concerning the Equality of the Angles of Incidence and Reflection, is preserv'd inviolable in Rays however heterogeneous, as the same great Man observes; it is better to use a *Speculum* instead of the Object *Lens*, which forming the Image of a distant Object at a considerable Distance, discovers the Errors that arise from the different Refrangi-bility of Rays sensible enough, and not at all to be dissembled, if the Rays falling obliquely be admitted by an Aperature sufficiently large, which is very often necessary: But in smaller *Lens's*, such as the ocular Ones, the Error is so small and insensible, that they may be still safely used.

James Gregory was the first who gave a Specimen of this sort of *Cata-Dioptrical Telescopes* consisting of *Lens's* and *Specula*, in *Optic. Promot. Prop. 59.* which was many Years afterwards given out by Mr. *Cassegrain* a French Man, for his own. The same, upon Physical, as well as Geometrical

cal Accounts alter'd and improv'd, is publish'd by Sir ISAAC NEWTON, in his admirable *Theory of Lights and Colours.*

Since *Specula*, being opaque Bodies, cannot have the same *Axis* with *Lens's* without being perforated at their Vertex, and consequently suffering an irreparable Loss (arising from both these Causes) of those Rays that fall near the Vertex, and are most accurately reflected: The Position of the *Axis* may (by *Carol. I. Prop. II.*) be alter'd as we please, by the help of a plane *Speculum*; and by this Means (besides other Advantages) the Necessity of Perforation is quite taken away, and the Loss of Rays falling near the *Axis* occasion'd by the Opacity, is very much diminish'd by the Obliquity of the second *Speculum*, which is observ'd by the Accurate *Newton* in his *Cata-Dioptrical Telescope*. But if by Reason of Physical Difficulties, in turning and polishing proper *Specula*, we must still continue the use of *Lens's*, perhaps it wou'd be of Service to make the Object *Lens* of a different *Medium*, as we see done in the Fabrick of the Eye; where the Chryalline Humour (whose Power of refracting the Rays of Light

The Elements of Dioptrics. 111

Light differs very little from that of Glass) is, by Nature, who never does any Thing in vain, join'd with the Aqueous and Vitreous Humours, (not differing from Water as to their Power of Refraction) in order that the Image may be painted as distinct as possible upon the Bottom of the Eye. There are likewise other Advantages of the fore-mentioned Artifice in the Animal Eye, which belong not to this Place.

F I N I S.

ADDEN.

Figure figures very well from the
Globe) is by Nature who never does
any thing in vain, for who
ever any Virtue or Humane (not super-
natural) in order to the Improvement
be required as fully as possible upon
the Bottom of the Earth. There are like-
wise other Advantages of the Earth which
belong not to this Figure.

ENIUS

ADDEIN



ADDENDA.



N order to remove every Rub that might lie in the way of a Reader who is to begin with this Subject, and interrupt his Progres thro' this most excellent Treatise ; besides having all along in the Translation explain'd such Passages as the Author's Laconick Style has made something too puzzling for a Beginner, and corrected several considerable Faults of the Press, which had escap'd his Care, I have thought it proper to subjoin a few Things, which may perhaps be of service to obviate some farther Difficulties, and supply some seeming Defects.

B

I. From

I. From Prop. VI. Probl. V.

We may give a Solution of the following Catoptrical Problem, for finding the Focus of any given Speculum universally; which the Author tells us, at the end of his Book, it was once his Design to have done.

The Problem is this. *The Focus, or Point in the Axis of a given Speculum, from whence or towards which Rays proceed, being given, to find the Focus or Point where those Rays will again meet, after they fall upon, and are reflected by a given Speculum.*

Now this Problem being solved in one Case, namely that of Rays falling upon a *Convex Speculum*, and diverging from a certain Point in the Axis of the Speculum, will, *mutatis mutandis*, be applicable to all other Cases whatever, whether the *Speculum* be *Convex*, *Concave*, or *Plane*, and whether the Rays fall *diverging*, *converging*, or *parallel*.

[Plate I. Fig. 9.] Let then BD be a given *convex Speculum*, whose Centre is A, and E a given Point in its Axis, from whence the Rays which fall upon the Speculum diverge; It is requir'd to find the *Focus*, or Point C, in which the Rays ED diverging from the *Focus*, or Point E, do, after reflection from the convex Speculum BD meet. Call EB the given Distance of the Point

Point E from B the Vertex of the Speculum d ; A B the given Radius of the Speculum r ; and B C the Distance of C the *Focus* required from B the Vertex of the Speculum x . By *Prop. VI. Probl. V.* $AC(r-x) : CB(x) :: AE(d+r) : EB(d)$. From whence arises this Equation, $dr - dx = dx + rx$, which gives us $BC(x) = -\frac{dr}{2d+r}$

for our general Rule; which shews that in the Case of *diverging Rays* falling upon a *convex Speculum*, the *Focus* C is always affirmative, and to be taken from the Vertex B directly forward: And that the greater d is, the greater will be the *Focal Distance*, till at last d becoming infinite, and consequently the finite Term $+r$ vanishing, it

will be $x = \frac{dr}{2d} = \frac{r}{2}$ or half the Radius: That

is, the *Radiant* receding from a *convex Speculum*, the *Image* will also recede beyond the Vertex of the Speculum, but so slowly that when the Radiant becomes vastly distant, the Image will have got no farther than the middle Point, between the Vertex and Centre of the Speculum: And *vice versa*, the nearer the Radiant approaches to the Speculum, the nearer the Image approaches to it, till at last they both meet and coincide at the Vertex, According to *Observat. 2. Schol. Prop. VIII.*

If $EB(d)$ be equal to $AB(r)$, the *Focal*

cal Distance BC becomes $= \frac{r}{3}$: but if $EB (d)$ be equal to half $AB (\frac{1}{2}r)$ the *Focal* Distance will in this Case become $= \frac{r}{4}$

Now to apply this Universal Canon to all other possible Cases; the Terms of which it consists remaining always the same, 'tis only changing the Signs + or — according as the Case requires. For in the Case of *converging Rays*, the Point in the Axis of the Speculum towards which they converge, is on the other side of the Vertex of the Speculum, and consequently its given Distance, in respect to the former, is *Negative*, or $-d$. For the same Reason, in the Case of a *concave Speculum*, the Centre of the Spherical Surface lying on the contrary side of the Vertex, the given Radius becomes *Negative*, or $-r$. And in these Cases respectively, wheresoever the quantity d or r occurs, it must have a contrary Sign to what it had before.

If then *converging Rays* fall upon a *convex Speculum*, d being in this Case negative, the Rule will be $\frac{-dr}{-2d+r}$; which

shews that when $2d$ exceeds r , the *Focus* will be still affirmative, but if $2d$ be less than r , the *Focus* will be *negative*, or on the contrary side of the Vertex of the Speculum.

culum. That is, if the Point behind the Speculum, towards which the Rays converge, be at a greater Distance from the Vertex of the Speculum than half the Radius, the *Focus* is still to be taken from B the Vertex directly forward, according to *Corol. Prop. IV.* But if its Distance be less than half the Radius, then the *Focus* must be taken from the Vertex B backwards, according to *Corol. Prop. VI.* When d is equal to

$\frac{1}{2} r$, the *Focal* distance becomes $= \frac{-dr}{o}$

and is consequently infinite. That is, Rays converging to a Point in the Axis of a convex Speculum, at an equal Distance between the Centre and Vertex, will after reflection proceed parallel, according to *Corol. Prop. III.* When d is equal to r , the *Focal* Distance becomes $= r$; that is, Rays converging towards the Centre of the Speculum are reflected by a convex Speculum back again upon themselves.

If *parallel* Rays fall upon a *convex* Speculum, d in this Case becoming infinite, the

Rule will be $\frac{dr}{2u} = \frac{1}{2} r$, which shews that

Rays falling parallel upon a convex Speculum, are collected in a Point at the distance of half the Radius behind the Speculum, according to what has been demonstrated at *Prop. III.* And consequently a vastly distant Radiant will have its Image form'd

in

in this Point; whence the Sun's Beams will be there collected, and heat or burn any thing placed therein.

If diverging Rays fall upon a *concave Speculum*, the Radius being in this Case $-r$, the Rule will be $\frac{-dr}{2d-r}$; which shews

that when d is less than $\frac{1}{2}r$, the *Focus* is *affirmative*, when d is equal to $\frac{1}{2}r$ the *Focus* is *infinite*, and when d is greater than $\frac{1}{2}r$ the *Focus* is *negative*; and when d is equal to r , the *Focal Distance* is $= -r$. That is, if the Point in the Axis of the Speculum from which the Rays diverge, be nearer the Vertex than half the Radius, the *Focus* will still be behind the Speculum, according to *Prop. V.* If it be just at the Distance of half the Radius, the Rays after Reflection, will proceed parallel, according to *Prop. III.* If it be at a greater Distance than half the Radius, the *Focus* after Reflection, will be on the same side of the Speculum with the *Focus* before Reflection, or Point from which the Rays diverge. If it be at the distance of the whole Radius, the Rays after Reflection meet in the same Point from which they first diverge. It is moreover manifest, that the more d exceeds $\frac{1}{2}r$, the less will be the *negative Focal Distance*; but if d be infinite, the *Focal Distance* in this Case $= -\frac{dr}{2d} = \frac{1}{2}r$,

can

can be no less than half the Radius: And on the contrary, the less d exceeds $\frac{1}{2}r$, the greater will be the *Focal Distance*; till at last, d becoming equal to $\frac{1}{2}r$, the *Focal Distance*, in this Case, $= \frac{-dr}{o}$, becomes infinite.

But when d grows any thing less than $\frac{1}{2}r$, the *Focus* becoming *affirmative*, is thrown at a great distance on the contrary side of the *Speculum*; and by how much d is less than $\frac{1}{2}r$, the less will be this affirmative *Focal Distance*. So that a *Radiant* placed at a greater Distance than half the Radius from a *concave Speculum*, the farther it recedes from the *Speculum*, the nearer its *Image* which is on the same side approaches to the *Speculum*, and at the *Centre* of the *Speculum* they meet, and afterwards cross one another, till the *Radiant* becoming *vastly distant*, the *Image* will come within half the Distance of the Radius from the *Speculum*: And *vice versa*, the nearer the *Radiant* approaches the *Speculum*, the farther the *Image* recedes from it, and at the *Centre* of the *Speculum* both meet, and afterwards cross one another, till at last the *Radiant* coming to half the Distance of the Radius from the *Speculum*, the *Image* becomes *vastly distant*: Whence if a lucid Body be placed at the distance of half the Radius from a *concave Speculum*, it will enlighten Places that are *vastly distant*. If the *Radiant* comes nearer the *Speculum* than

than half the Radius, the Image will be cast from before the Speculum to a great distance on the contrary side ; and the nearer the Radiant now approaches to the Speculum, the nearer will the Image likewise approach to it, till at last they both coincide at its Vertex, and *vice versa*, according to *Observat. I. Schol. Prop. VIII.*

If *converging Rays* fall upon a *concave Speculum*, d and r being in this Case both negative, the Rule will be
$$\frac{dr}{-2d-r}$$

which shows that the *Focus* is always *negative*. That is, Rays that fall converging upon a concave Speculum, will always be collected in a Point or *Focus* on this side the Speculum ; according to *Corol. Prop. V.* If d be equal to r , the *Focal Distance* is $= \frac{-r}{3}$; but if d be equal to $\frac{1}{2}r$, the *Focal*

Distance is $= \frac{-r}{4}$.

If *parallel Rays* fall upon a *concave Speculum*, d in this Case becoming infinite, the Rule will be
$$\frac{dr}{2d} = -\frac{1}{2}r$$
. That is,

Rays falling Parallel upon a concave Speculum, are collected in a Point at the distance of half the Radius on this side the Speculum;

culum, as has been demonstrated at *Prop. III.* Whence if the Sun's Beams be receiv'd upon such a *Speculum*, the same will be the burning Point.

If *diverging Rays* fall upon a *plane Speculum*, the Radius r being infinite, the Rule will be $\frac{dr}{r} = d$; That is, the *Focus* of diverging

Rays reflected by a Plane *Speculum*, will be at as great a distance behind the *Speculum*, as the Point from which they diverge is before the *Speculum*; according to *Prop. II.* And consequently the Image of any Radiant, made by Reflection from a Plane *Speculum* will be seen as far behind the *Speculum* as the Radiant is before the *Speculum*; and they will both not only recede from, and approach to the *Speculum* at the same time, but likewise keep an equal pace with one another.

If *converging Rayss* fall upon a *plane Speculum*, d being negative, and r infinite, the Rule will be $-\frac{dr}{r} = -d$; that is, Rays

converging to a Point at a certain Distance behind a Plane *Speculum*, will have their *Focus* at an equal Distance before the *Speculum*; according to *Corol. I. Prop. II.*

If *parallel Rays* fall upon a *plane Speculum*, both d and r being in this case infinite,

finite, the *Focal Distance* $\frac{dr}{2d+r}$ will be also infinite. That is, Rays falling parallel upon a *Plane Speculum*, will be reflected back parallel.

It is worth observing, that the consideration of *diverging Rays* relates to Objects that are near us, and such as we examine with our naked Eye, or by the help of a Microscope: *Parallel Rays* are consider'd when we have to do with Objects vastly distant, and such as we look at thro' Telescopes. And *converging Rays* fall under our Consideration, when the Rays proceeding diverging or parallel from any Object are by one Speculum or *Lens* made to converge, and then intercepted by the Interposition of another Speculum or *Lens*, before they arrive at their Point of Convergence; which is of great use in examining the Effects of Optical Machines, made by a Combination of more than one *Lens* or *Speculum*, and constructing such as are proper for any assign'd purpose, where this Contrivance is often absolutely necessary.

There are two farther *Uses* of this Method, the first is, to determine what degree of *Convexity* or *Concavity* is necessary for a *Speculum* to represent a given Object at a given *Focus*. And this is very easily done from the Equation first found $dr - dx = dx - rx$: For d and x being given; the Radius of Convexity

Convexity $r = \frac{2dx}{d-x}$. Where it is plain,

that if x be greater than d , r will be a negative Quantity, and the Problem impossible for a convex *Speculum*. That is, if it be required to represent a given Object, at such a given *Focus*, whose Distance on the other side of the *Speculum* shall be greater than the Distance of the Object on this side; instead of a convex Glass, we must use a concave, whose negative Radius will be

$= \frac{2dx}{x-d}$. After the same manner the Degree of Concavity is found from the same Equation, only changing the Sign of r from + to - : For d and x being given, we shall have the Radius of Concavity $r = \frac{2dx}{x-d}$.

Where it is manifest, that if d be greater than x , r will be a negative Quantity, and the Problem impossible for a concave *Speculum*. That is, if it be required to represent a given Object, at such a given *Focus*, whose Distance on the other side of the *Speculum*, shall be less than the Distance of the Object on this side, instead of a concave *Speculum*, we must make use of a convex,

whose affirmative Radius is $= \frac{2dx}{d-x}$.

And in both Cases if d be equal to x , then the Radius either of Convexity or Concavity

$= \frac{2dx}{+d-x} = \frac{2dx}{0}$, will be infinite, and

- +

C 2

the

the Problem will be impossible for either a convex or concave Speculum. That is, if it be required that a given Object shall be represented at such a given *Focus*, whose Distance on the other side of the Speculum shall be equal to the Distance of the Object on this side, instead of a *convex* or *concave* Speculum, a *plane* Speculum is the only one that can be used. If d be infinite, the

Radius of Convexity becomes $\frac{2dx}{d} = 2x$;

but the Radius of Concavity $\frac{2dx}{-d} = -2x$,

will have a negative Value, which shews that the Problem is impossible in the Case of a concave Speculum. That is, a vastly distant Object cannot be represented at any given *Focus* behind a concave Speculum; but may by a Convex, whose Radius of Convexity must be equal to twice the given *Focal Distance* behind the Speculum. If x be infinite, then the Radius of Convexity

$= \frac{2dx}{-x} = -2d$, has a negative Value, and

the Problem is impossible for a convex Speculum; but the Radius of Concavity is

$= \frac{2dx}{x} = 2d$: Whence if we would have a

given Object represented at an infinite Distance behind the Speculum, we can make use of none but a concave Speculum, whose

Radius

Radius of Concavity must be equal to twice the given Distance of the Object.

Hitherto we have consider'd the *Focus* as affirmative, that is, behind or beyond the Speculum; but the same Rule, only changing the Sign of x in the Equation $x = \frac{dr}{2d+r}$, will equally hold if we would have the *Focus* negative, or on the same side of the Speculum with the Object. For in the Case of a convex Speculum we shall have the Radius of Convexity $= \frac{-2dx}{d+x}$ always of a negative Value, and consequently the Problem is always impossible for a convex Speculum; but in the Case of a concave Speculum, the Radius of Concavity will be $= \frac{-2dx}{d-x}$, which shews that the Problem is always possible for a Concave Speculum, be the quantities d and x as they will. If we would have $d = x$, then the Radius of Concavity will be also equal to d or x : That is a Concave Speculum, whose Radius is equal to the distance of the Object from the Speculum, will reflect the Image into the same place with the Object.

The other Use of this Method is, from the *Image* and *Speculum* given, to find the *Distance* of the Object from the *Speculum*. That is in the Equation first found

dr

$dr - dx = dx + rx$, x and r being given, we are to find d , which will consequently

be $= \frac{rx}{r - 2x}$, in *convex Specula*, and in

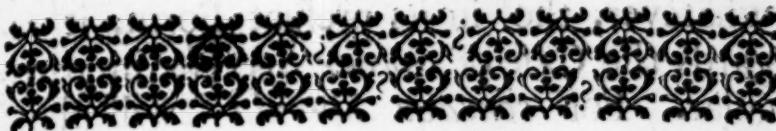
concave $= \frac{rx}{r + 2x}$. Whence 'tis plain that

in *convex Specula*, the Problem will be impossible, when x exceeds $\frac{1}{2}r$, but in *concave Specula* it will always be possible. That is, if the Image is to be at a greater distance than $\frac{1}{2}$ the Radius behind the Speculum, it cannot be made by an Object placed before a convex Speculum at what distance soever: But let the Distance of the Image be behind the *Speculum* be what it will, it may be form'd by an Object exposed at some certain distance before a concave Speculum. If the given *Focus* be negative, or the Image on the same side of the Speculum with the Object, then changing the Sign of x , in the forementioned Equation, we shall have, in the Case

of *convex Specula*, $d = \frac{-rx}{r + 2x}$, and in the

Case of *concave Specula* $d = \frac{-rx}{r - 2x}$. So

that the Problem will always be impossible for *convex Specula*, and only possible for *concave ones*, when $\frac{1}{2}r$ does not exceed x , or when the Image is not nearer the Speculum than by half the Radius,



II. From Corol. Prop. IX.

WE may deduce a Solution of the following Catoptrical Problem, of magnifying or diminishing a given Object by a given Speculum in any assign'd Proportion.

The Problem is this ; To find at what Distance from a given Speculum it is necessary to place an Object, in order that the Homologous Lines of the Image made by the Speculum, may bear any assign'd proportion to those of the Object.

Since it is evident from this Corol. that the Homologous Lines of the Radiant and Image, are to one another as their Distances from the Speculum respectively : It follows, That if b to c express the Proportion which the Homologous Lines of the Object and Image are to bear to one another, b will be to c , as d the Distance of the Object to x the Focal Distance of the Image. Whence if we compare this value of x with that deliver'd in the foregoing Problem, we shall have $\frac{cd}{b} = \frac{dr}{2d+r}$. And consequent-

ly the Distance required $d = \frac{br-cr}{2c}$, will be our general Rule ; and will, *mutatis mutandis*,

tandis, extend it self to all possible Cases whatever, tho' in its present Form, it regards the Case of *convex Specula* in particular. For *concave Specula* 'twill stand thus $d = \frac{cr - br}{2c}$. If the Image be desired

on the same side of the Speculum with the Object, x being negative, in the Case of *convex Specula*, 'twill be $d = \frac{-br - cr}{2c}$;

and in the Case of *concave* $d = \frac{br + cr}{2c}$.

From whence 'tis plain that there is no magnifying an Object by a *convex Speculum*, for c being in this Case greater than b , the Rule for the affirmative *Focus* $d = \frac{br - cr}{2c}$ will have a negative Value,

and that for the negative *Focus* $d = \frac{-br - cr}{2c}$ has always a negative Value:

So that we can only diminish an Object, and make it appear less, by a *convex Speculum*, and that only when the *Focus* is affirmative, or the Image to be represented behind the Speculum. And by a *concave Speculum* there is no diminishing an Object, as long as the *Focus* is affirmative, for b being greater than c , the Rule in that Case $d = \frac{cr - br}{2c}$ will have a negative

tive

tive Value ; so that we can only magnifie an Object, and make it appear greater behind a concave Speculum. But if the *Focus* be negative, and the Image and Object to be both on the same side of the Speculum, the Rule being $d = \frac{br + cr}{2c}$, shews that in this Case a concave Speculum will magnifie or diminish an Object in what proportion we please.

It is to be observ'd, that if the Object be a Right Line, the Proportion b to c will express the Proportion between the Object and Image themselves ; but if the Object be a Plane Figure, the Proportion b to c will be only Subduplicate of that which the Object bears to the Image, as we learn from *Euclid*. So that if b to c be as 2 to 1, the Object and Image will be as the Squares of these Numbers, or as 4 to 1. But it must also be noted, That Painters usually measure the largeness of their Figures by the simple Proportion of their homologous Lines ; so that when they speak of an Human Figure twice as big as the Life, their meaning is that the Homologous Lines of this Figure are twice as great as those of the Life ; or that the Dimensions of every Member in length and breadth are twice as great as those of the Man represented ; tho' properly speaking, the Picture is four times as big as the Life.

D

If

If the out Lines of the Image be desired twice as big as the Life, and the *Focus* affirmative, c being in this case greater than b , the Problem will be impossible for any convex Speculum; but the Object being placed before a concave Speculum at the Distance $\frac{cr - br}{2c} = \frac{r}{4}$, or $\frac{1}{4}$ of the Radius, will have its Image magnified in the Proportion assign'd. If we would have the *Focus* negative, and the Image represented on the same side of the Speculum with the Object, still the Problem will be impossible for a convex Speculum; but if the Object be placed before a concave Speculum at the Distance $\frac{br + cr}{2c} \geq \frac{3r}{4}$, or $\frac{3}{4}$ of the Radius, its Image made by the Speculum will be magnified in the Proportion assign'd.

Thus let the Proportion b to c be what it will, the Rule will always give us the Distance, at which the Object must be placed before the given Speculum, in order to have its Image magnified, or diminished in that Proportion. I shall only add one Instance more, and that is, supposing b and c were equal, and the *Focus* affirmative: In this Case we should have both for convex and

$$\text{concave Specula } d = \frac{\frac{1}{2}br - \frac{1}{2}cr}{2c} = \frac{o}{2c} = o:$$

That is, the Object must be placed at the very

very Vertex of the Speculum, in which Case we know both Object and Image coincide. If the *Focus* were negative, then no convex Speculum will do; and the Rule for *concave Specula* will be

$$d = \frac{br + cr}{2c} = \frac{2cr}{2c} = r : \text{That is, the Ob-}$$

ject must be placed in the Centre of the Speculum.

There are two farther *Uses* to be made of this Method; the first is, The Distance at which the Object is to be placed before the Speculum, and the Proportion in which the Image is to be magnified, or diminished, being given, to find what Degree of Convexity or Concavity the Speculum should have, in order to magnifie or diminish the Image in the Proportion assign'd. That

is, in the Rule $d = \frac{br - cr}{2c}$, d, b , and c be-

ing given, we are required to find r ; which will give us in the Case of a convex Speculum, and an affirmative *Focus*, $r = \frac{2cd}{b - c}$;

and in the Case of a concave Speculum, $r = \frac{2cd}{c - b}$. If the *Focus* be negative, for

convex Specula, the Rule stands $r = \frac{2cd}{b - c}$;

for concave, $r = \frac{2cd}{b + c}$.

From all which it appears, that if c be greater than b , or if the Image be desired greater than the Object, and to be represented behind the Speculum, no Convex will do, but a Concave will, whose Radius is $\frac{2cd}{c-b}$. As likewise, if b be greater than c , or the Image be desired less than the Object, and to be represented behind the Speculum, no Concave will do, but a Convex will, whose Radius must be $\frac{2cd}{b-c}$. If the *Focus* be required negative, or the Image to be on the same side of the Speculum with the Object; the Problem is altogether impossible for a convex Speculum, whether to magnifie or diminish; and always possible for a concave Speculum either to magnifie or diminish.

The other *Uſe* we may make of this Method, is from the Distance of the Object before the Speculum, and the Radius of Convexity or Concavity being given, to find the Proportion b to c , which the Object will bear to its Image made by the given Speculum. That is, in the foremention'd Rule, having d and r given, to find the Proportion of b to c . Whence in the Case of a convex Speculum, and an affirmative *Focus*, 'twill be $b:c::2d+rr$; And in the Case of a Concave, $b:c::r-2d,r$. But if

if the *Focus* be negative, in *convex Specula*, 'twill be $b \cdot c :: 2d - r \cdot r$; and in *concave*, $b \cdot c :: 2d + r \cdot r$.

So then in the Case of a convex Speculum, and the Image behind the Speculum, b will always be greater than c , because $2d + r$ is of necessity greater than r : And in the Case of a Concave, b will always be less than c for the like reason; and if d be equal to $\frac{1}{2}r$, then c will be infinite in respect to b . But if the Image be required on the same side of the Speculum with the Object, in *convex Specula*, 'twill be found always impossible: And in *concave*, possible in all Cases whatever, both of magnifying and diminishing; for if d be greater than r , then b will be greater than c ; if d be less than r , b will be less than c ; and if d be equal to $\frac{1}{2}r$, then c will be infinitely greater than b .





III. At Corol. 3. Prop. XV.

THE Author gives a Construction to find the Focus of Rays refracted at a Spherical Surface, and towards the latter end of that Corol. applies the same to the Case of a plane Surface: Affirming, that in this Case the Right Line FM , which determines the Focus by its meeting the Axis BA produced somewhere in C , will be the same with a Right Line, joining the Point F with another Point taken in BE produced, at such a Distance from B , as to make that Distance bear the same Proportion to BE which AN does to AM . Because the Tangents of Angles are reciprocally as their Cotangents.

[Plat. III. Fig. 5.] That this may be more easily conceiv'd, Suppose the Surface BD plane, and take the Point C upon BE produced, so as that BC may be to BE , as AN

AN to AM, and join FC. We are to shew that the Right Line FM produced, will in this Case meet the Axis AB produced, in the Point C, and there determine the Focus. Join EN, EM, and on the Centre E with any Radius, as EA, describe an Arch of a Circle AG, and draw the Tangents of the Angles AEN, AEM, which will always be as AN to AM, and are in the present Case those very Lines, as likewise their respective Cotangents GH, GK. Since A the Centre of the plane Surface BD, and consequently the Right Line AMN is at an infinite distance from B, the Right Line FM becomes parallel to EM; and consequently if produced beyond F, will meet the Axis AB produced somewhere, suppose at C, so as to make the Triangle CBF similar to the Triangle KGE. And therefore the right Line EH, passing thro' the Vertex of both those Triangles, will cut their Bases CB, GK similarly in E and H, so as to make BC to BE, as GK to GH. But because the Tangents of Angles are reciprocally as their Cotangents, AN is to AM, as GK to GH; therefore BC is to BE as AN to AM. And consequently the Point C, where the right Line FM produced meets the Axis AB produced, is that very Point C taken at first upon BE produced, so as that BC may be to BE as AN to AM. Q. E. D.



IV. At Prop. XVI. Probl. XI.

THE Author recommends the use of Analytical Calculations, for finding the Foci of Lens's, as far better than the very nicest Geometrical Constructions. For which reason it cannot be improper to give the less skilful Reader an Example, that he may see how such Calculations are to be manag'd.

[Plat. III. Fig. 6.] Suppose BD to be the given *Lens*, and E a Point in its Axis, from which Rays diverging fall upon the *Lens*, A the Centre of its first Spherical Surface, and C the Centre of its second Spherical Surface, BD the thickness of the *Lens*, and I to R the *Ratio* of Incidence to Refraction. And it is requir'd to find F the *Focus* of those Rays after Refraction at both Surfaces of the *Lens*. We must first find f the *Focus* of those Rays after their Refraction at the first Surface only, or their *Focus of Transition*. Call EB *d*, BD *t*, AB *r*, CD *e*, Bfx, DFy. By *Proposit. XV.* EA (*d+r*): Af (*x-r*) + fB (*x*): BE (d)

(d) :: I : R : Whence multiplying the Extremes and Means, $R dx + R rx = Idx - Idr$; and $Bf(x) = \frac{Idr}{Id - Rd - Rr}$.

Whence it is plain, that if d be so great in respect to r , that Id exceeds $Rd + Rr$, the *Focus f* is affirmative, and to be taken from B the Vertex of the refracting Surface directly forward, as at *Fig. 6*: If Id is less than $Rd + Rr$ the *Focus f* is negative, and to be taken from B backwards, as at *Fig. 7*; and if Id be equal to $Rd + Rr$, *Bf* becomes infinite, and the Rays proceed parallel.

Thus having found *f* the *Focus*, after the first Refraction, we may by the same means find *F* the *Focus* after the second Refraction. For by the same

Prop. XV. $fC \left(\frac{Idr}{Id - Rd - Rr} - t + \epsilon = \frac{Idr - Idt + Rdt + Rrt + Id\varrho - Rd\varrho - Rrp}{Id - Rd - Rr} \right)$:

$C F (y + \epsilon) + F D (y) : D f$
 $\left(\frac{Idr}{Id - Rd - Rr} - t = \frac{Idr - Idt + Rdt + Rrt}{Id - Rd - Rr} \right) ::$

R : I : (I to R at the Emergence of Rays from any Lens, being as R was to I at their Immersion into it.) Wherefore multiplying the Extremes and Means, we have

$$I^{\cdot}dry - I^{\cdot}dty + IR^{\cdot}dty + IR^{\cdot}rty + I^{\cdot}dey - IR^{\cdot}dey - IR^{\cdot}rey = IR^{\cdot}dry - IR^{\cdot}dty + R^{\cdot}dty + R^{\cdot}rty + IR^{\cdot}dre - IR^{\cdot}dte + R^{\cdot}dte + R^{\cdot}rte; \text{ And consequently, } DF(y) =$$

$$\frac{IR^{\cdot}dre - IR^{\cdot}dte + R^{\cdot}dte + R^{\cdot}rte}{I^{\cdot}dr - IR^{\cdot}dr - I^{\cdot}dt + 2IR^{\cdot}dt - R^{\cdot}dt + IR^{\cdot}rt - R^{\cdot}rt + I^{\cdot}dp - IR^{\cdot}de - IR^{\cdot}rp}$$

Which Equation if we put $p = \frac{R}{I - R}$,
may be abridged, and reduced to $DF(y) =$

$$\frac{Ipdre - Rdpt + Rprt}{Idr - Idt + Rrt + Idp + Rdt - Iprp}$$

And is evidently the same with that, which the famous Dr. *Halley* has given long ago, in the *Philosophical Transactions*, for finding the *Foci of Optical Glasses universally*.

This Calculation being general, will serve for all sort of *Lens's*, be the Matter of which they are made, and the Ambient *Medium* what they will, or what ever be the *Ratio* of *I* to *R*: And tho' it is made for *Lens's* whose *Surfaces* are both *convex*, yet *mutatis mutandis*, it will extend to *Lens's* of any other Figure whatever, whether *double-convex* or *double-concave*, *plano-convex* or *plano-concave*, or *convexo-concave*, which last sort are commonly call'd *Menisci*. For the *Radius* of a *concave* *Surface* being on the contrary side, or *negative* with respect to that of a *convex*, and the *Radius* of a *plane* *Surface* *infinite*: 'tis only changing all

all the Signs \pm or $-$ with which the Radius of the respective Surface, which we would have *concave* instead of convex, is affected in the general Rule ; or making all the Terms infinite, which involve the Radius of the respective Surface, which we would have *plane* instead of convex. So likewise if we would have it extend to other Rays besides *diverging* ones ; the distance of the Point where *converging* Rays meet, from the first Surface of the *Lens*, being on the contrary side or *negative*, in respect to that of *diverging* Rays, and the distance of the Point where *parallel* Rays meet, from the same Surface, being *infinite* : 'Tis only changing the Signs of all the Terms where we meet with d , if the Rays are supposed *converging* ; or making those same Terms infinite if the Rays are supposed *parallel*.

In the Case of a *double-Convex* of Glass, if the ambient *Medium* be Air, I being to R as 3 to 2, we shall have the *Focal Distance* from the second Surface of the *Lens*,

$$y = \frac{6dr\rho - 2dpt \pm 4rpt}{3dr - dt \pm 2rt \pm 3dp - 6r\rho} : \text{ if the ambient Medium be Water, I being to R as 9 to 8, the Rule will be}$$

$$y = \frac{72dr\rho - 8dpt \pm 64rpt}{9dr - dt \pm 8rt \pm 9dp - 72r\rho} \text{ For a double-Convex of Water, and the ambient Medium}$$

Medium of Air, I being to R as 4 to 3, the

Rule is $y = \frac{12dr\rho - 3dpt + 9rpt}{4dr - dt + 3rt + 4dp - 12r\rho}$

And for a *double-convex* of Diamond, in a Medium of Air, I being to R as 5 to 2, the Rule

would be $y = \frac{5dr\rho - 2dpt + 4rpt}{5dr - 3dt + 2rt + 5dp - 10r\rho}$

If the thickness of the *Lens* be neglected, which is generally not considerable, the Terms where t occurs being rejected, the Rule is still farther abridg'd to

$y = \frac{pd\rho}{dr + d\rho - pr\rho}$. Where it is evident,

that if d be so small in respect to r and ρ , that $dr + d\rho$ is less than $pr\rho$, the *Focal Distance*

y will be negative, and $= \frac{pd\rho}{-dr - dp + pr\rho}$

or the Rays after the two Refractions at both Surfaces of the *Lens*, will still proceed diverging from some Point, before the second Surface of the *Lens*; and if $dr + d\rho$ be equal to $pr\rho$, y is infinite, and the Rays after Emersion from the *Lens* proceed parallel. The Error in neglecting t is so small, that if for the ease of the Calculation, we suppose a *Lens* of Glass equally *convex* on both sides, and expos'd to *parallel Rays*; r being in this case equal to ρ , and d infinite, the *Focal Distance* when t is neglected is

$\frac{2dr}{2dr} = r$, and when it is consider'd

$\frac{6rr - 2rt}{6r - t}$; which is only $\frac{rt}{6r - t}$ or nearly

$\frac{1}{6}t$ less than the former. In the case of converging Rays falling upon a *double-convex* of Glass, we have $y = \frac{-2dr\rho}{dr - d\rho - 2r\rho}$, always affirmative: And if the Rays are *parallel*, d being infinite, 'twill be

$y = \frac{2dr\rho}{dr + d\rho} = \frac{2r\rho}{r + \rho}$, which also gives the *Focus* always affirmative, or behind the *Lens*. *Diverging* Rays falling upon a *double-concave*, give $y = \frac{2dr\rho}{dr - d\rho - 2r\rho}$

always negative, as in the case of *converging* Rays on a *double-convex* 'twas always affirmative: But if the Rays are *converging*, it will be $y = \frac{-2dr\rho}{dr - d\rho - 2r\rho}$ affirmative,

when $dr - d\rho$ is less than $2r\rho$, or when the *Focus* of *diverging* Rays collected by a *double convex* is negative, and *vice versa*; If the Rays are *parallel*, 'tis

$y = \frac{2dr\rho}{dr - d\rho} = \frac{2r\rho}{r - \rho}$, always negative. A *plano-Convex* Glass, the *plane* Surface being exposed to *diverging* Rays, gives,

r being infinite, $y = \frac{2dr\rho}{dr - 2r\rho} = \frac{2d\rho}{d - 2\rho}$, affirmative or negative, according as d is greater

ter or less than 2ρ ; if exposed to *converging*

Rays, $y = \frac{-2dp}{-d - 2\rho}$, always affirmative;

if to *parallel* Rays, $y = \frac{2dp}{d} = 2\rho$, so that the

Image of a vastly distant Object is always form'd by a *plano-Convex Lens*, the plane side being turn'd towards the Object, just at the Distance of the Diameter of the second Surface behind it. A *plano-concave Lens* exposed on the *plane* side to diverging

Rays, gives $y = \frac{-2dr\rho}{dr + 2r\rho} = \frac{-2d\rho}{d + 2\rho}$, al-

ways negative; to converging Rays,

$y = \frac{2d\rho}{-d + 2\rho}$, affirmative, when the *Fo-*

cus of diverging Rays on a *plano-Convex* is negative, and *vice versa*; to *parallel* Rays,

$y = \frac{-2dp}{d} = -2\rho$, so that the Image of

a vastly distant Object, is always form'd by a *plano-Concave* at the distance of the Diameter before the second Surface, as it is by a *plano-Convex* behind. A *Meniscus* expos'd on the *concave* side to *diverging* Rays, gives

$y = \frac{-2dr\rho}{-dr + d\rho + 2r\rho}$, affirmative only

when d and r are so great in respect to ρ , that dr exceeds $d\rho + 2r\rho$: To *converging*

Rays, $y = \frac{2dr\rho}{dr - d\rho + 2r\rho}$, affirmative or ne-

gative,

gative, according as $dr + 2rp$ is greater or less than dp : If to parallel Rays $\frac{2drp}{dr + dp} = \frac{2rp}{r + p}$, affirmative, if the Radius of *concavity* is greater than the Radius of *convexity*; and negative if less; and infinite if equal, for the Effects of the first Surface are, in that case, exactly destroy'd by the second, and the Rays suffer'd to proceed still parallel.

It is to be observ'd, that if the thickness of the *Lens* is neglected, as inconsiderable, the *Focus* of all sorts of Rays falling upon any *Lens* will be exactly the same, upon which soever Surface of the *Lens* they are first receiv'd. But if the thickness of the *Lens* be consider'd, there will be some difference in the *Focal Distance*, according as you turn this or that Surface of the *Lens* towards the Rays. And this difference is easily found from the general Rule: For upon turning the other Surface of the *Lens* towards the Rays, p becomes r , and r changes to p ; by which means the Rule will give us the *Focal Distance* in both Cases, and subtracting one from the other, we find their Difference. Thus, if, to abridge the Rule, we suppose the Rays parallel, d being infinite, we have for a *double-Convex* in one case $y = \frac{Iprp - Rpt}{Ir - It + Ip - Rt}$, and up-

on

on turning the *Lens*, $y = \frac{Ippr - Rrt}{Ip - It + Ir + Rt}$; wherefore subtracting one from the other, according as p or r is greatest, we shall have the difference in *double Convex*, occasion'd

by turning the *Lens* $= \frac{Rrt - Rpt}{Ir - It + Rt + Ip}$

or in Glass $\frac{2rt - 2pt}{3r - t + 3p}$. And this is applicable to *Lens's* of any other Figure, by changing the Signs + or - of those Terms, where we meet with r or p , or making them infinite, according as the respective Surfaces are *concave* or *plane*. Thus in the case of a *plano-convex*, r being infinite, the Difference arising upon turning the *Lens*,

becomes $\frac{Rrt}{Ir} = \frac{R}{I}t$; or in Glass $\frac{1}{3}t$, in Water $\frac{2}{3}t$, and $\frac{5}{3}t$ in Diamond. Which shews that the *Focal Distance* is greater by $\frac{2}{3}t$ when the *plane* side of a *plano-Convex* of Glass is turn'd towards a *vastly distant Object*, than when the *convex* side is turn'd to it. After the same manner, the negative *Focal Distance* of a *plane Concave* will be greater by $\frac{2}{3}t$, when the *plane* side is turn'd towards the *vastly distant Object*, than when the *concave* side is turn'd to it. In *double-concaves*,

caves; where the *Focus* is always negative,

the Difference is $\frac{\pm 2pt \mp 2rt}{3r - t - 3p}$, according

as p is greater or less than r . In *Menisci*, the *Focal Distance*, whether affirmative or negative, being always greatest when the *concave* Surface is turn'd towards a vastly distant Object, the Difference is

$\frac{-2rt - 2st}{3r - t + 3p}$ when the *Foci* upon turning either side fall both one way, and are either both affirmative or negative: But if t be so considerable as to be greater than $3e$, the *Focal Distance* upon turning the *concave* side $\frac{-6re - 2pt}{3r - t + 3p}$ is affirmative, and the *Focal Distance* upon turning the *convex* side $\frac{-6er + 2rt}{3e - t - 3r}$ negative, and conse-

quently their Difference is $\frac{2rt - 12rp - 2st}{3r - t + 3e}$:

And if t were equal to $3e$, the *Focus* in this case upon turning the *convex* side to a *vastly distant* Object, falling exactly upon the *vertex* of the second Surface of the *Lens*, and consequently the *Focal Distance* being equal to nothing, the Difference will be the same with the *Focal Distance* upon turning the *concave* side, namely, $\frac{-6re - 6ee}{3r}$. After

the like manner, may be found from the general Rule, the Difference which would arise upon turning the different Surface of any sort of *Lens* towards other Rays besides *parallel*, whether *diverging* or *converging*; but the *Canons* for these Cases consist of so many Terms, and are of so little use, that they are not worth having.

There are three farther *Uses* to be made of the general Rule above deliver'd; The first is, from the *Lens* and *Focus*, where an Object is represented, being given, to determine the *Distance* of the Object from the *Lens*; or the *Lens* by which we wou'd form the Image of any Object, and the *Focus* where we would have it form'd, being given, to determine the Distance at which the Object should be placed before the *Lens*, that it may be represented in the given *Focus*.

That is, in the Equation $y = \frac{pd\varrho}{dr + d\varrho - pre}$,
 or $dry + d\varrho y - prey = pd\varrho r$, r , p , ϱ , and y being given, 'tis required to find d , and consequently we shall have $d = \frac{pr\varrho y}{ry + ey - pre}$;

where 'tis plain, that if r and ϱ be so great in respect to y , that pre exceeds $ry + ey$, d will be negative; and the Object cannot be represented in the Circumstances requir'd, unless by means of another *Lens*, we first make

make the Rays coming from the Object, diverging converge to a Point behind the first Surface of the *Lens* given, at the Distance of

$\frac{pr\rho y}{pr\rho - ry - \rho y}$: And if $pr\rho$ is equal to $ry + \epsilon y$, d will be infinite. Suppose the given *Lens* a *double convex* of Glass, and made of two Segments of equal Spheres, but of a thickness not considerable, and it is required to find at what Distance from the *Lens* a lucid Body should be placed, in order to have its Beams parallel after their Emergence from the *Lens*, and consequently its Light thrown upon Objects *vastly distant*, which may be thereby illuminated : In this case y being infinite, and r equal to ρ , and p equal to 2, we shall have the Distance requir'd $d = r$. But if t be considerable, we must find d from the Rule which takes in the thickness of the *Lens*, which gives us the exact value of

$$d = \frac{4rpt + 6r\rho y - 2ryt}{3ry - ty + 3\rho y - 6r\rho + 2\epsilon t} : \text{As, if}$$

for Example, the *double-convex* just mention'd, were an entire *Sphere* of Glass, and the same thing requir'd as before, y being, as we have already observ'd, infinite, and r equal to ρ , and moreover t equal to $2r$;

$$\text{this last Rule gives } d = \frac{6rry - 4rry}{3ry - 2ry + 3ry} =$$

$\frac{1}{2}r$, whereas by the former which neglects

the thickness, we have the Distance required twice as great, or $d = r$; a Difference very considerable, if the Spheres be of any bigness. So then, a lucid Body placed at the Distance of half the Radius from a Sphere of Glass, or at the Distance of the whole Radius from a *double-convex* of equal Spheres, whose thickness is inconsiderable, will illuminate Objects vastly distant. If the given *Lens* were a *Hemisphere* of Glass, and the same thing still requir'd; if the *convex* Surface be first, both y and ρ being in this case infinite, and t equal to r , the Distance of the lucid Body will be $d = \frac{6r\rho y}{3\rho y} = 2r$, but if the *plane* side be next the lucid Body, y and r being infinite, we have $d = \frac{6r\rho y - 2rty}{3ry} = \frac{2}{3}\rho$ or $\frac{4}{3}r$; there being as has been shewn before a Difference of $\frac{1}{3}t$ in the *Focal Distance* of a *plano-convex* expos'd to *parallel* Rays, occasion'd by turning the different sides of the *Lens*: If t had been neglected, we should have had $d = 2r$ in both cases. If we have an Object represented by a *double-concave* of Glass of equal Spheres, at a negative *Focus* the Distance of the Radius from the *Lens*, and it were requir'd to find the Distance of the Object, y , r , and ρ being all negative and equal, and t inconsiderable, we shall have $d =$

$d = \frac{-2rrr}{2rr - 2rr}$ infinite, and consequently the Object is *vastly distant*. The same thing may be done for all other Cases whatever, only remembering to make the proper Alterations according as r , ϵ , or y are negative or infinite, and t considerable or inconsiderable.

The second Use is, from one Surface (either the first or second) of a *Lens* being given already form'd, to find what degree of *convexity* the other Surface must have, in order to represent a given Object at a given *Focus*. That is, in the Equation before used $dry + dy - pr\epsilon y = pd\epsilon$, d , y , p and r being given, to find ϵ , or ϵ being given to find r : Whence we have

$$\epsilon = \frac{dry}{pdr - dy + pry}, \text{ and } r = \frac{dy}{pd\epsilon - dy + pey};$$

which will serve for any other *Lens*'s besides *double-convex*, and any other Rays besides *diverging* ones, by making such alterations as has been already directed. If the first Surface of a Glass *Lens* were *plane*, and it were requir'd to find what Degree of Convexity the second Surface must have, in order to represent an Object at a *Focus* just as far distant from the *Lens* as the Object it self: In this case d is equal to y , and

$$r \text{ infinite, and consequently } \epsilon = \frac{dry}{2pry} = \frac{1}{4}d,$$

so that the second Surface of the *Lens* must be

be made of a Segment of a Sphere, whose Radius is equal to $\frac{1}{2}$ of the Distance of the Object. If the Object to be represented at a given *Focus*, be *vastly distant*, d being in this case infinite, the Rule is abridg'd to

$p = \frac{ry}{pr - y}$; whence 'tis plain, that if in Glass y is greater than $2r$, or the given *Focus* be at a greater Distance from the *Lens* than twice the Radius of the given Surface, p will be negative, and the second Surface must be made *concave*; and if y be equal to $2r$, p is infinite, and the second Surface must be *plane*. If the thickness of the *Lens* be so great, that it ought to be consider'd, we must find p from the general Rule,

The third *Use* is, from the *Lens*, *Distance* of the Object, and *Focus* being given to determine the *Ratio of Refraction*. That is, in the Equation before us $dry + dpy - prey = pdrp$, d , r , p and y being given, to find

p , which gives us $p = \frac{dry + dpy}{drp + rpy}$: For p

being found, the value of $p = \frac{I}{I - R}$ gives the *Ratio I to R as I + p to p*. If a *double-convex Lens*, made of two Segments of the same Sphere, represents, or is required to represent a *vastly distant Object*, at a *Focus* the Distance $\frac{1}{2}r$, from the *Lens*; d being in this case infinite, and

and r equal to ρ , and y equal to $\frac{1}{2}r$, we shall have $p = \frac{3dr}{dr} = 3$, and consequently the *Ratio* I to R as 4 to 3; whence the *Lens* is made, or ought to be made of the *medium* of Water. If the *Focal Distance* were $\frac{1}{2}r$, we should have $p = \frac{1}{2}$, or I to R as 5 to 2, and consequently the *Lens* would be *Diamond*. If the *Focal Distance* were $4r$, then we have $p = 8$, or I to R as 9 to 8, and the *Lens* is *Glass*, and the *ambient medium* *Water*. But if we are very curious in determining the *Ratio* of *Refraction*, it is done more exactly when the *Lens* is form'd into an *Hemisphere*, or a *plano-Convex*, and receiving the *Rays* of the Sun upon its *plane* side, collects them in a Point at some Distance behind, which must be measured with great niceness; because in this case, our neglecting t occasions no Error at all. In this case, if the *Focal Distance* is equal to thrice the *Radius* of the *Sphere*, d and r being infinite, and y equal to 3ρ , 'twill be $p = 3$, or I to R, as 4 to 3; if y is equal to 2ρ , I is to R as 3 to 2; and if y is equal to ρ , then I is to R as 2 to 1.





V. At Prop. XX. Probl. XIII.

WHICH is, *To find the Distance at which an Object should be placed from a given Lens, so as that the Image form'd by the Lens may bear a given Proportion to the Object.*

The Author has given the *Construction*, but omitted the *Demonstration*; leaving the process of the *Calculation* which points out that *Construction*, as a tryal of Skill to the diligent Reader. But because my design in publishing the Book is to make it entirely easy; for fear it may prove too difficult or discouraging a Task to some who are either not skilful enough, or perhaps too lazy to go through with it, I have subjoin'd the following *Solution*.

[*Plate III. Fig. 6, 7.*] Let the given *Lens* *BD* be a *double-convex*, and call *AB* the *Radius* of the first *Surface* *a*, *CD* the *Radius* of the second *Surface* *b*, *BD* the *thicknes* of the *Lens* *C*, the *Proportion* of the *homologous* *Lines* of the *Object* and *Image* as *r* to *b*, and

and EB the Distance required z , at which the Object is to be placed before the *Lens* B , which we suppose made of Glass and the ambient *medium* Air. 'Tis plain there are two different Values of $EB(z)$, according as the *Focus* F is affirmative, or beyond the *Lens*, as at *Fig. 6.* Or negative, and on the same side with the Object, as at *Fig. 7.* Both which Cases shall be respectively consider'd, and included in the Demonstration.

By *Corol. Prop. XVIII.* the homologous Lines of the Object and Image are to another as their respective Distances from the *Lens*; wherefore r is to b , as $EB(z)$ the Distance of the Object required to DF the *Focal Distance* of the desir'd Image, which is consequently $\frac{bz}{r}$. But we shall have another value of this *Focal Distance* from *Prop. XV.* For if we look upon f as the *Focus* of Rays sent diverging from the Object at E , after their Refraction at the first Surface of the *Lens*, and F their *Focus* after both Refractions; and call the first *Focal Distance* Bf x , and the second DFy . Before we can find the *Focal Distance* DF , which determines the place of the *Image*, we must first find Bf . Now to find Bf , by *Prop.*

$$XV. \frac{EA}{3} \left(\frac{z+a}{3} \right) : AB(a) :: Ef \left(\frac{z+y}{z+y} \right) : Bf$$

G

$Bf(y)$; whence $zy + ay = \pm 3az + 3ay$,
and $Bf(y) = \frac{\pm 3az}{z - 2a}$. Having found Bf ,
the same Proposition gives us DF ,
for $\frac{fC}{2} \left(\frac{\pm 3az}{z - 2a} \right) = c \pm b =$

$$\frac{\pm 3az \mp cz \pm bz \mp 2ac \mp 2ab}{2z - 4a} : CD$$

$$(b) : : fF \left(\frac{3az}{z - 2a} - c \mp x = \frac{3az - cz \mp xz + 2ac \pm 2ax}{z - 2a} \right) : DF$$

(x) ; whence $\mp 3azx \pm czx \pm bzx$
 $\mp 2acx \mp 2abx = 6abz - 2bcz \mp 2bzx + 4abc \pm 4abx$, and $DF = \frac{6abz - 2bcz + 4abc}{\mp 3az \mp cz \pm 3bz \pm 2ac \mp 6ab}$. There-
fore comparing the two values of DF
together, we have this Equation $\frac{bz}{r} =$

$$\frac{6abz - 2bcz + 4abc}{\mp 3az \mp cz \pm 3bz \pm 2ac \mp 6ab}, \text{ or}$$

$$\mp 3abz \mp bcz \pm 3bz \pm 2abcz \mp 6abz = 6abrz - 2bcrz + 4abcr.$$

From whence we have $ZZ = \frac{6abZ - 2aCZ \pm 6arZ \mp 2rCZ \pm 4arC}{3a - C + 3b}$.

And if the thickness of the *Lens* be neglect-
ed as inconsiderable, all the Terms where

C occurs vanishing, we have $zz = \frac{6abz \pm 6arz}{3a + 3b}$, and consequently $EB(z) = \frac{2ab \pm 2ar}{a + b}$. Q. E. I.

[*Plat. II. Fig. 22.*] If then we would construct this Equation, we have this Proportion given us for that purpose, $a + b : b \pm r :: 2a : z$. Taking therefore A B equal to a , and B C in the same right Line equal to b ; from C draw at pleasure the indefinite right Line C M, upon which cut off C D equal to C B (b), and from D on either side take D M, or Dm such, that it may bear the same *Proportion* to D C, which the *Homologous* Lines of the Object do to those of the Image. Join A M or Am, to which thro' B draw B E or Be parallel; and twice M E or twice m e is the *Distance required* in the *Problem*. For A C ($a + b$) : M C or m C ($b \pm r$) :: 2AB ($2a$) : 2 ME or 2 m e (z). And consequently the *Construction* gives the true value of z , as before found. Q. E. D.

The same *Problem of magnifying or diminishing a given Object by a given Lens, in any assign'd proportion*, may be solved from the Equation above given, for finding the *Foci* of all sort of *Lens's*, which if express'd in the *Characteristicks* in present

use, is $r = \frac{pabz}{az + bz - pab}$. For supposing as before r to b expresses the proportion which the *Homologous Lines* of the Object are required to bear to those of the Image, and 'tis desir'd to find z the Distance of the Object from the *Lens*, which is necessary to perform the Conditions required: By *Corol. Prop. XVIII.* we have another value of the *Focal Distance* $x = \frac{bz}{r}$. Whence comparing

both together we have $\frac{bz}{r} = \frac{pabz}{az + bz - pab}$, and consequently if the *Focus* is to be affirmative $z = \frac{pab + par}{a + b}$, but if pab be greater than $az + bz$, then the *Focus* is negative, or on the same side with the Object, and this negative *Focal Distance* is $\frac{pabz}{-az - bz + pab}$, and consequently $z = \frac{par - par}{a + b}$; That is, if the *Lens* be Glass, the following Equation includes both cases, whether the Image is to be represented on the contrary or same side with the Object, $z = \frac{2ab \pm 2ar}{a + b}$: Where it is to be obser'd, that if the Image is represented on the contrary side by a *double-convex Lens*, or at an affirmative *Focus*, it may be made either equal

equal to, greater or less than the Object in what proportion we please; but if it is represented on the same side, or at a negative *Focus*, r must always be less than b , and consequently the Image may be shewn larger than the Object in all the degrees imaginable, but never less, nor equal; for when b is equal to r , and the *Focus* negative, z is $= 0$, and when b is less than r , z is negative and impossible.

Nor is this Solution confin'd to the case of *double-convexes* only, tho' made for *Lens*'s of that Figure, but will with proper Alterations extend equally to *Lens*'s of all other sorts whatever; only observing to change the Signs $+$ or $-$ with which the Radius of a *concave* Surface is affected, or making the Terms infinite where the Radius of a *plane* Surface occurs; because it has been shewn before, that the Radius of a *concave* Surface bears a contrary Sign to that of a *convex*, and the Radius of a *plane* Surface is infinite. If then the *Lens* be a *concavo-Convex* or *Meniscus* of Glass, and the first Surface concave, the Rule for both cases, whether the *Focus* is to be affirmative or negative, is $z = \frac{-2ab \mp 2ar}{-a + b}$; and if the second Surface be that which is *concave*, then for the affirmative and negative *Foci* respectively, the Rule becomes $z = \frac{-2ab \pm 2ar}{a - b}$. Where it is to be observ'd, that

that in the first case if the *Focus* be affirmative, a must be greater than b , or else z will be negative, and the Problem impossible; and if the *Focus* be negative, and b greater than r , then must a be still greater than b , but if b be less than r , then a must be less than b ; or else the Problem will be impossible: And in the last case, if the *Focus* be affirmative, and b greater than r , then b must be greater than a , and *vice versa*; and if the *Focus* be negative, b must still be greater than a . Which shews that in the case of a *Meniscus* turn'd on the *concave* side towards the Object, the Image can never be represented at all on the opposite side, unless the *concave* Surface be a Segment of a larger Sphere than the *convex*, and then it may be shewn in what Proportion to the Object we please; and if it is to be represented on the same side with the Object, and magnified, the Radius of *concavity* must be still larger than that of *convexity*, and *vice versa*, if the Image is to appear diminished. And in like manner may be understood what will happen upon turning the *convex* side of the *Meniscus* towards the Object. If the *Lens* be a *double-concave* of Glass, the *Focus* being in this case always negative, we have but one value of z , which is $\frac{2ab - 2ar}{a - b}$, affirmative only when r is greater than b ; which shews that

a double-concave can only diminish. If the *Lens* be a *plano-Convex* of Glass, 'twill be $z = 2b \pm 2r$; which shews that in this case the Object may either be magnified or diminished, if the *Focus* be affirmative, but only magnified if it be negative. It must be noted, that if the second Surface, whose Radius is b were suppos'd plane, we shou'd have $z = 2a$, because b not only stands for the Radius of that Surface, but also expresses the *ratio* of the Image to the Object, which is consequently in this case infinite, and the Image *vastly distant*. If the *Lens* be a *plano-Concave* of Glass, the *Focus* being always negative, we have only one value of z , which is $z = -2b + 2r$, which shews that a *plano-Concave* can only diminish.

It may be remark'd, That those Cases which make the Distance z negative, and the Problem impossible for the given *Lens*, may by means of another *Lens* be made practicable: If we first receive the Rays of the Object upon this second *Lens*, and before they are collected at the *Focus* throw them upon the given *Lens*, in such a manner as to make them fall *converging* to a Point behind the first Surface of the given *Lens*, at the Distance of the negative value of z .

If

If the *Lens* were a *double-convex* of *Wa-
ter*, p being $= \frac{R}{1-R}$, 'tis $z = \frac{3ab \pm 3ar}{a+b}$;
if of *Diamond* $z = \frac{\frac{2}{3}ab \pm \frac{1}{3}ar}{a+b}$; if of *Glass*
in an ambient *Medium* of *Water* $z = \frac{8ab \pm 8ar}{a+b}$; if of *Diamond* in an ambient
Medium of *Water* $z = \frac{\frac{8}{3}ab \pm \frac{8}{3}ar}{a+b}$, and con-
formably in all other Cases.

If the *Image* be desir'd just as great as the *Object*, then for a *double-convex* of *Glass*, r being equal to b , the Rule gives the Distance required $z = 2b \pm 2a$; where 'tis plain the Problem is always possible when the *Focus* is affirmative, but if the *Focus* is negative, 'tis only possible when b is greater than a . If the *Lens* were equally *con-
vex*, b being in this case equal to a , there will only be one value of z , and that for an affirmative *Focus* $z = 4a$: So that if the *Object* be placed at four times the Distance of the Radius from the *Lens*, the *Image* form'd at its *Focus* will be just as great as the *Object*. If the *Homologous Lines* of the *Image* were desir'd twice as large as those of the *Object*, r being in this case equal to $\frac{1}{2}b$, for a *double-convex* of *Glass* the Rule gives $z = 2b \pm a$, and if the *Lens* be equally *convex*

convex $z = 3a$ or a according as the *Focus* is affirmative or negative, so that an Object placed at the distance either of thrice or only once the Radius from the *Lens*, is represented at a *Focus*, either affirmative or negative, twice as large every way as the Object, or the Image in its whole content will be four times as large as the Object. If the *Homologous* Lines of the Image were desired twice as small as those of the Object, r being in this case equal to $2b$, the Rule gives for an equally convex *Lens* only one value of the Distance $z = 6a$, a *double-convex* being only capable of representing an Object diminish'd, when the *Focus* is affirmative. If the Image be desir'd an hundred times larger than the Object, or its *Homologous* Lines ten times as large, r being in this case equal $\frac{1}{10}b$, if the *Lens* be equally *convex*, the Rule gives the Distance $z = 2\frac{1}{3}a$, or $\frac{7}{3}a$, according as the *Focus* is affirmative or negative. And in general if m to n express the Proportion which r bears to b , the Rule laid down at first $\frac{pab \pm par}{a + b}$, becomes

$pb \pm p \frac{m}{n}a$, and if the Convexity be equal,

'tis $z = pa \pm p \frac{m}{n}a$. From all which it appears that with a *Lens* equally Convex on both sides, in order to magnify a given Object, the Distance is always something greater

ter than pa , when the *Focus* is affirmative, and always less than pa when the *Focus* is negative, unless the *Ratio* of the Image to the Object be infinitely great, and then n being infinite, 'tis $z = pa$. That is, in Glasses the Distance of the Object must always exceed twice the Radius, if the *Focus* is to be affirmative, or fall short of it if negative, and be equal to it when the Image is to be infinitely great, or vastly distant. And in order to diminish a given Object $\frac{m}{n}$ in this case exceeding *Unity*, the Distance, which has but one Case here, is always greater than $2pa$; and the more z exceeds $2pa$ the more the Object is diminished, and *vice versa*, till z becoming equal to $2pa$, the Object and Image are likewise equal.

All this is easily observ'd in that common Experiment of the *Camera Obscura*: Where the Rays propagated from External Objects are received by a *Lens*, and transmitted into the Room, and do there paint upon a white Sheet placed at the *Focal Distance* of those Rays from the *Lens*, the Images of their respective Objects, in Colours scarce less lively than those of the Objects themselves. And hence this Problem of magnifying or diminishing a given Object, may be of great use in *painting*; for by admitting the Image of any Object by means of a *Lens* into a dark Chamber, in what *Ratio* to

to the Life we please, 'tis easy to hit the proportion of every part with great exactness, which is otherwise but seldom done; especially in such Pieces where the Figures are either much greater or much less than the Life. 'Tis true, these Images that are represented in the *dark Chamber* by a single *Lens* appear inverted; but may be made erect by using a second *Lens* after the following manner. Place the *Object* at such a Distance from the first *Lens* in the Window, that the *Image* form'd by that, may be just as big as the Life; then beyond the place of this *Image* fix the given *Lens* at the Distance requir'd α ; this will form an *Image* of the former *Image* which shall be in the *Ratio* assign'd, and also erect, but something less lively. What has been above deliver'd is likewise of service in the Construction of the *Magick Lanthorn*, and other Optical Machines, where the *Images* of any *Objects* are to be represented monstrously larger or less than the Life.





VI. At Prop. XXIV. Probl. XVII.

W^Hich is, *To make a concave Speculum of Glass of a given Thickness; the Radius of whose concavity is also given, in such a manner, that parallel Rays reflected from the first Surface of the Speculum may meet in the same Point of its Axis, with those that pass refracted into the Speculum, and are reflected from the second Surface, and again refracted at their Emer-
sion from the Speculum.* Or the first Surface of a *concave Speculum* being already form'd, to determine of what Sphere 'tis necessary to take a Segment to form the second Surface in such manner, that an *Object vastly distant* may be represented by *Reflection* from both Surfaces in one and the same Place, or that the two Images may be united, and consequently be made more lively. The Author has given two several *Construc-
tions* of the *Problem*; the first exact, and

the other only near the Truth, but more expeditious: But for Reasons already mention'd he has omitted their *Demonstration*, which is as follows.

[*Plat. III. Fig. 8.*] Let AB be the given Radius of the first Surface, and BD the given Thickness of the *Speculum*, and CD the Radius required of the hindmost Surface necessary to perform the Conditions of the *Problem*. Suppose the Point f in fD , the Axis of the *Speculum* to be the *Focus* of the *parallel Rays*, after *Refraction* at their entrance into the first Surface, φ their *Focus* after *Reflection* from the second Surface, and F their *Focus* after *Refraction* again at their *Emersion* from the *Speculum*; 'tis required that the Point F should be the same with the *Focus* of the same *parallel Rays* after their *Reflection* from the first Surface. Call the given Radius ABa , the given Thickness BDe , the first *Focal Distance* Bf , the second $D\varphi x$, the last *Focal Distance* BFy , and the required Radius of the hindmost Surface CDz . We must find these *Focal Distances* one after another, in order to determine the last BF , which must be equal to the *Focal Distance* of *parallel Rays* reflected from a *concave Speculum*. For the first then Bf , by *Prop. XIV.* $Bf(v):fA(v-a)::I:R::$ (in Glass) $3:2$; whence $Bf(v) = 3a$. And for the second

second $D\varphi$, by Prop. IV. $D\varphi(x) : \varphi C$
 $(z-x) :: Df(3a+c) : fC(3a+c-z)$;
 whence $3ax+cx-zx=3az+cz-3ax-cx$,
 and consequently $D\varphi(x) = \frac{3az+cz}{6a+2c-z}$.

Lastly, To determine the Focal Distance
 BF after both Refractions at the first Surface, and the Reflection from the last, by
 Propofit. XV. the Emersion being out of

Glass into Air, $\frac{\varphi A}{2} \left(a+c - \frac{3az-cz}{6a+2c-z} \right)$

$$= \frac{6aa+8ac-4az+2cc-2cz}{12a+4c-2z} : AB$$

$$(a) :: \varphi F \left(\frac{3az+cz}{6a+2c-z} - y - c = \frac{3az+2cz-6ay-2cy+zy-6ac-2cc}{6a+2c-z} \right)$$

BF(y); whence $6aay+8acy-4azy+2ccy-2czy=6aaaz+4accz-12aay-4acy+2azy-12aac-4acc$, and consequently $BF(y) = \frac{3aaaz+2accz-6aac-2acc}{9aa+6ac-3az+cc-cz}$.

Now this must be equal to the Focal Distance of parallel Rays reflected from a concave Speculum, which by Prop. III. is just half the Radius, whence we have another value of BF(y) = $\frac{a}{2}$.

And

And comparing both together we have

$$\frac{3aa z + 2ac z - 6aac - 2acc}{9aa + 6ac - 3az + cc - cz} = \frac{a}{2}, \text{ or}$$

$$9az + 5cz - 18ac - 5cc = 9aa, \text{ and consequently the Radius required } CD(z) = \frac{9aa + 18ac + 5cc}{9a + 5c}.$$

Q. E. I.

Fig. 3. If then we would *Construct* this value of $CD(z)$, we have the following Proportion given us for that purpose $9a + 5c$:

$$\sqrt{9aa + 18ac + 5cc} :: \sqrt{9aa + 18ac + 5cc} :: z.$$

Wherefore making the Rectangular Triangle MLK in such manner, that LM shall be equal to $9a + 5c$, and LK equal to $\sqrt{9aa + 18ac + 5cc}$, and then drawing from K the Perpendicular SK , the Right Line LS is the Radius required. For by *Element. VI. 8.* $LM (9a + 5c) : LK (\sqrt{9aa + 18ac + 5cc}) :: LK (\sqrt{9aa + 18ac + 5cc}) : LS(z)$. And consequently the *Construction* gives the true value of $CD(z)$ as before found. *Q.E.D.*

Fig. 4. The Radius required is also capable of another *Construction*, for making an actual *Division* of $9aa + 18ac + 5cc$ by $9a + 5c$, the *Quotient* is $a + c \div \frac{4ac}{9a + 5c}$; And if $AB(a)$ be sufficiently great in respect to $BD(c)$, the Term $5c$ in the *Nominator*

nominator of the *Fraction* may be neglected, and then it becomes $z = a + c + \frac{2}{3}c$, whose excess above the Truth is not at all sensible. Wherefore taking DO equal to $\frac{2}{3}$ of BD(c), and making AC equal to DO, CD ($a + c + \frac{2}{3}c$) is very nearly the true value of the *Radius required*. Q. E. D.

To shew how near this last value of $z = a + c + \frac{2}{3}c$ is to the Truth, if we suppose the Thickness of the *Speculum* to be $\frac{1}{5}$ of the given Radius of the first Surface, which is very considerable: In this case the exact value of $z = a + c + \frac{4ac}{9a + 5c}$ is but $\frac{2}{225}a$ or nearly $\frac{1}{45}a$ less than its value found by neglecting $\frac{5c}{a}$ in the other Equation $z = a + c + \frac{2}{3}c$, a Difference not all considerable in *Physical Matters*.

If c be supposed equal to $\frac{1}{3}a$, which is still a far greater Supposition, even in this Case, the value of the Radius z taken from the last Equation $z = a + c + \frac{2}{3}c$, is but $\frac{1}{625}a$ or very nearly $\frac{1}{125}a$ greater than the Truth, which is an excess not very sensible, unless a be extremely great. But if the Thickness be greater than in this last Supposition, it will be convenient to take the value of the Radius required from the exact Equation

$z = a + c + \frac{4ac}{9a + 5c}$. As, if c be equal to $\frac{1}{2}a$, the Radius required is $1\frac{1}{4}\frac{1}{6}a$, which is

is about $\frac{1}{2}a$ less than what it would be if $5c$ were neglected. If c be equal to a , the Radius required is $2\frac{2}{3}a$, which is about $\frac{1}{6}a$ less than it would be if $5c$ were neglected. If the Thickness be so considerable as to be equal to thrice the given Radius of the first Surface, then the Radius required is $4\frac{1}{2}a$.

Hence likewise if the Radius of the last Surface be given, together with the Thickness of the *Speculum*, we may find the Radius of *concavity* necessary to unite the two Images of a vastly distant Object made by *Reflection* from both Surfaces. For if c be not very great, we shall have $a = z - c - \frac{4}{9}c$, as near the Truth as need be required in Practice: For if we had the exact value of a , we cou'd not in practice grind the *Speculum* to the due Concavity, even so near as the value just now given. If c be considerable, the value of a must be found from an Equation of an higher Degree $9aa - 9az + 18ac = 5cz - 5cc$; which if it be contracted, by putting p for $2c - z$,

$$\text{will give } a = \mp \frac{1}{2}p \pm \frac{\sqrt{5cz - 5cc}}{9} + \frac{1}{4}pp,$$

the Sign of $\frac{1}{2}p$ being either $-$ or $+$ according as $2c$ is greater or less than z . After the same manner, having the Radius of both Surfaces given, we may find what

I

Thick.

Thickness of the *Speculum* is necessary to unite the two Images of a *vastly distant Object*, form'd by *Reflection* from both Surfaces, by means of the following Equation, $5cc + 18ac - 5cz = 9az - 9aa$, which if it be contracted, by putting q for $3\frac{1}{3}a - z$, will give $c = \mp \frac{1}{2}q \pm \sqrt{9az - 9aa + \frac{1}{4}q^2}$.





VII. A more particular Account of
MICROSCOPES and TELESCOPES,
from Mr. *Huygens*.

PROP. I.

TO explain the Effects and Uses of single Microscopes, and the manner of making little Spheres and Lens's.

[Fig. 9, 10.] Let N be the *Lens* QRQ, the Object at its *Focus* R, O the Eye very near the *Lens*. The Rays coming from R will after their Refraction fall parallel upon the Eye, and consequently make distinct Vision. For the Fabrick of the Eye, having its *Focal Distance* just at the Bottom of it upon the *Retina*, requires that the Rays from each single Point should fall nearly parallel in order to be there collected: that is, that the Basis of each Cone of Rays flowing from every Point of any Object, which Basis is the *Pupil* of the Eye, should bear so small a proportion to the Length of the Cone, as

that those Cones may be look'd upon as little Cylinders. (The Distance indeed requisite for distinct Vision is not limited to a point, but is indulged in larger Bounds ; because Nature has furnish'd us with the power of contracting the Pupil as the Object comes nearer, and so diminishing the Basis of each Cone in proportion, and consequently of preserving distinct Vision, but this is only to a certain and that no very great Degree.) But the Object QRQ will appear in the same magnitude, as if the *Lens* N were removed, and a Plate with a small hole in it substituted in its place, namely under the Angle QAQ. So that all the interposed *Lens* does in this case, is only making distinct Vision which would without the *Lens* be confused. But since at the distance suppose of 8 Inches from the Object, a naked Eye has then distinct Vision ; the apparent Image may be said to be so much magnify'd as those 8 Inches exceed the little Space NR, or the *Focal* Distance of the *Lens* N : which if it be equal to $\frac{1}{2}$ of an Inch, the appearance of the Image seen distinctly by the *Microscope*, is to that seen distinctly by the naked Eye as 40 to 1. Therefore the less the *Focal* Distance of the little *Lens* N is, the greater will its Effects be in dilating the Image of a small Object ; tho' there are some Inconveniences (to be mention'd afterwards) which here offer themselves,

themselves, and forbid our going beyond some certain Limits. And the same thing happens to little *Spheres*, which may be us'd for *Lens's*, and might otherwise be made as little as we please. But these small *Spheres* are inferiour to little *Lens's* upon this Account, that for the same degree of magnifying, if both be made of Glass, the *Lens's* are three times more distant from the Object than the *Spheres*; and by that means leave a sufficient Space for the lateral Light to enter, and make the Colours of the Object visible; whereas otherwise we are forced to turn the *Microscope* directly against the Light, and can only than discern distinctly such Objects as by their thinness are pellucid.

The Effects of a little *Sphere*, and what has been said concerning the three times less Distance is thus demonstrated. *Fig. 11. 9.*

Let there be a Glass *Sphere* whose Centre is K, and its Axis A B, in which produced on both sides the Eye is plac'd at D, and the Object at C, each of the Distances A D, B C being taken equal to half the Radius A K; And consequently the Point C is the *Focus* where Rays falling parallel to the Axis A B, upon the *Sphere* at A H are after emersion collected, as is shewn at A D D E N D. IV. Wherefore an Object plac'd at C will send Rays upon the *Sphere*, which will after Refraction be receiv'd parallel by the Eye, and consequently

consequently make distinct Vision. But by *Prop. XIV.* if we take the Point F such, that FA may be equal to the Radius AK, the Point OF is the *Focus* towards which parallel Rays after Refraction at the first Surface BG do tend in their passage thro' the *Sphere*, and from which they are diverted after Refraction at their Emerson and collected at D. Make GE parallel to the Axis, and comprehending the Portion of the Object CE, and draw the right Line ED. The Ray then EG being refracted at G, proceeds according to GF, and being again refracted at H, goes on to meet the Eye at D. Wherefore the Line CE is seen under the Angle ADH, which would appear to the naked Eye under the Angle CDE, which I say is but the half of the former.

For because AF is double of AD, the Angle ADH is double of AFH. But DF is parallel to EG, because GE is both parallel to FD, and to be look'd upon as equal to it or to the right Line BC; because CE is to be a Line very small with respect to the Diameter of the *Sphere*. Therefore the Angle ADH is also double of the Angle CDE, and consequently equal to the Angle CKE. From whence it is plain that to the Eye placed at D, the Line CE will appear under the same Angle, in which it wou'd appear to the naked Eye seeing from the Point K.

Whence

Whence if the Diameter of the little *Sphere* A B were $\frac{1}{2}$ of an Inch, we should have K C $= \frac{1}{6}$ of an Inch, which is to the Distance of 8 Inches in the proportion of 1 to 128; so that the increas'd Magnitude of the Object wou'd be as 128 to 1, which is indeed very considerable. But if N R the *Focal Distance* of the *Lens* be equal to the right Line K C, we have shewn that by this means the Object R Q would be seen in the same Magnitude as if the Eye were plac'd at N without the *Lens*; Nor in the using this *Lens* will the apparent Magnitude be any ways changed, in whatsoever part of the Axis R N produc'd the Eye be plac'd. Therefore 'tis plain the same degree of Magnifying, and the same Effect every way is perform'd equally by the *Lens* N and the little *Sphere* A B. And it is moreover manifest that the Distance R N, being taken equal to K C, is equal to thrice B C, *Q. E. D.*

We are next to explain the manner how little Spheres and Lens's may be prepar'd and fitted for Use.

The less *Spheres* are to be, the easier they are prepar'd after the following manner. Take the smallest Fragments of Glafs, and hold them to the lower part of the Flame of a Candle where the bluish Colour is discernible, that they may grow red hot; and then if they be taken up by the fineſt Steel-Wire that can be got, and dexterously turn'd, they will be form'd into Globules,

Globules, which are large enough if equal to a Grain of Mustard-Seed. Out of several thus prepar'd, you will find some very good; which may be tried by including them in a Brass-plate, and is thus done. Take a Plate of the thinnest Brass the breadth of a Finger, and twice as much in Length, and bend it double; perforate this Rectangle in the middle with the Point of a Needle, and rub the opposite holes smooth with a Whetstone, that no roughness may remain about the Edges, and black them with the smoak of a Candle, that no brightness may continue within. Put the little *Sphere*, still adhering to the Steel Wire, into the Holes within the Brass Plate, and fix it there by fastning the two Leaves of the Plate together. After this manner you may make several *Microscopes* with great ease, out of which you may make choise of those that are best.

The principal Use of this sort of *Microscopes*, is to look at Corpuscles that are pellucid. And they are placed in a *Machine* made in such manner, that by turning a Screw they may approach to or recede from the Object, and so be brought to the due distance, which is requisite for distinct Vision. And to this it conduces very much, that the too great Light be restrain'd, and only admited thro' the Hole, which is about four times the distance from the Object. For
by

by this means the Aperture of the *Lens* is better limited, than by the breadth of the contiguous Hole, which there is no necessity at all for straitning. The Eye must be brought as close to the little *Sphere* as may be, that it may comprehend the greater Space.

The Corpuscles or Drops of Liquors which are to be look'd at, are put upon a little circular plane of Glass, which is made to slide laterally every way, that we may bring every part of the Object to be view'd successively. Some attract the Liquor to be examin'd into *capillary* Tubes of Glass so small as scarce to admit an Hair, which has likewise its Uses. But in using those little *Lens*'s before mention'd, care must be taken that while by means of another *Lens* on one side we cast Light upon the Object, the Hole of the apperture may be exactly limited, by trying how much it may lie open without being an Hinderance to distinct Vision. For here the Points of Corpuscles emit Rays of Light, and are so many *Radiant* Points, which is quite otherwise in those pellucid Corpuscles that are look'd at thro' little *Spheres*, where the Objects intercept the Light, not emit it.

The Effects of this sort of little *Lens*'s and *Spheres* are very wonderful, as may be seen from those Experiments with them which have been made publick, and from which

our Knowledge of Nature has receiv'd very great Light and Information. By these the Circulatory Motion of the Blood has been put beyond controversy, which our *Lewen-hoek*, the most diligent Observer of these Matters, has shewn me in the Tail of an Eel to my very great satisfaction and delight. For the Blood appears pellucid, and consisting of reddish Globules, and runs thro' the Channels of the Arteries, which are continu'd to the Veins with a very rapid Motion. Which without doubt might be observ'd in all other Animals, if we cou'd find out such parts in them as are pervious to the Light. He put the live Eel into a Glass Tube half full of Water, to which he externally applied the *Microscope* at that part where the extremity of the Tail touch'd the Tube.

'Tis also very pleasant to observe the *Animalcula* that swim in drops of Water, in which we have infus'd Ginger, Pepper, or something else of an hot powerful Odour for some Days. They are of various Forms, and some less than others; their Motions are wonderful and for their size sufficiently quick, nor is the Instrument apparent by which they perform them, for they have neither Legs nor Arms, nor do they bend their Bodies like Fishes. For the little Eels in Vinegar, which are much larger than they, swim like those in the River, in which

it

it is very much to be wonder'd that they shou'd generate little ones of themselves. For I saw one which had four young Ones within it, (for they are altogether pellucid) and after it had been kept in the Tube for some Hours brought them all forth, every one of which did afterwards swim by it self.

It is very probable that those *Animalcula* which I have said move about in Water, are invited thither out of the Air by the O-dour of the Infusion. For the same Figures appear upon macerating several things in Water, but if the Vessel be cover'd none at all appear. Nor is it difficult to conceive how they shou'd be supported in the Air, when they are so much smaller than the finest Dust that is. So that perhaps we draw many thousands of them into our Lungs every time we fetch our Breath, without knowing it. Nor wou'd it be useles to observe at what time of the Year they appear in greatest Numbers, and whether they encrease in a vitiated Air. Milk appears to consist of small pellucid Globules swimming in a Liquor likewise pellucid, but of a different refraction ; and hence it is that it appears white, tho' it contains no other matter but what is perfectly transparent, and without Colour.

I omit those many wonderful Forms of minute Insects ; the Wings of Butter-flies

and Gnats, cover'd with little Feathers ; the Powders observ'd in the middle of Flower-Tops, which are nothing else but little transparent Bladders fill'd with that matter, of which the Bees make their Honey, and which they carry between their Legs into their Hives. But what ought to be look'd upon as the most wonderful and astonishing of all, is that an immense multitude of *Animalcula* are discover'd to swim in the *Senen Masculinum*, after the manner of little Fishes, almost of the same Figure with a Frog newly form'd and yet without Legs. Which *Animalcula*, I make no question, enter the *Ova Muliebria*, and are the Rudiments of what is brought forth from thence. There are several Considerations which confirm this Opinion, nor is it any great Objection that out of so great a Multitude either few or only one of them comes to maturity and grows to be an Animal ; since the same abundance and superfluous fruitfulness is equally observable in most Seeds of Trees and Herbs, as of Firr, Poppies, &c.

These *Animalcula* by reason of their wonderful smallness (for even ten Thousand of them are not equal to the smallest grain of Sand) ought to be look'd at thro' such Glass Globules as have the greatest power of Magnifying.



PROP. II.

To explain the Effects of Compound
MICROSCOPES.

WE come now to speak of *compound Microscopes*, by the help of which such Objects as are not transparent are look'd at, and their true Colours discover'd, and that much better and more commodiously than thro' single *Lens's*.

[*Fig. 12. 13.*] Suppose the *Microscope* be a double one, consisting of two *Lens's*, one less A, and another greater B. Why we dispose them so, we shall afterwards explain. And let B be the *ocular Lens* nearest the Eye, plac'd suppose at C; A the *Object Lens* nearest the Object placed suppose at E; and ABC the common Axis of both *Lens's*. There will be two Cases, as may be seen represented in the two Figures to which this *Prop.* refers. In the first, Rays proceeding from a single point E of the Object and falling upon the *Lens* A, are refracted by it and again united in the Point P, and there inter-

intersecting one another and proceeding towards the *Lens B*, are by it refracted and made parallel, and so enter the Eye at C and by that means make distinct Vision. 'Tis necessary therefore that AE the distance of the Object shou'd be greater than AQ the *Focal* distance of the *Lens A*. And the *Focus* P must be found by *Prop. XVI.* or by making EQ, EA, EP, in a continual Proportion. But the *Lens B* is to be so plac'd, that its *Focus* on that side towards A, may fall exactly upon the Point P, in order that the Rays may be made parallel after refraction at the *Lens B*. All which is easily done by what has been before demonstrated. The other Figure (13.) represents the several Rays DAG, FAH, EAB, proceeding from different Points of the Object. A is the middle point of the Lens, and AP, AB, AC, are made in a continual proportion in order to determine the place of the Eye C; for by this means however small the Aperture of the little *Lens A* may be, the whole *Lens B* will nevertheless be fill'd with the Image of the Object, because the Rays falling from A upon the whole *Lens B*, are collected in the Point C.

[*Fig. 13.*] But the proportion of the apparent Magnitude to the true, will be found by drawing the right Line CF. For the proportion required will be the same with that which the Angle BCH bears to the Angle ECF; which proportion is compounded of

the

the proportion of the Angle BCH to BAH , and of the Angle BAH or $EA F$ to the Angle ECF . But the first of these is the same with the proportion of the right Line AB to BC , and the last that of CE to EA , because Angles that are small are look'd upon to be to one another as their Tangents. Therefore the proportion of the apparent Magnitude to the true, will be compounded of the proportions AB to BC or AP to PB (for AP , AB , AC are in a continual proportion) and CE to EA . But that the Effects of the *Microscope* maybe more exactly estimated, the Angle BCH is rather to be compared with the Angle, under which the right Line EF would be seen at the distance of 8 Inches from the Eye, that is with the Angle ELF , LE being taken equal to 8 Inches, according to what has been said before of magnifying by a single *Lens*. And consequently the proportion of magnifying must be here understood to be compounded of the proportion of the Angle BCH to BAH , and BAH or $EA F$ to ELF ; that is, of the proportion of AP to PB , and of EL , a line 8 Inches long, to the right Line EA . For if the *Microscope* were of so great length, that for example, CE should be two foot long, that is equal to thrice the right Line LE ; and the apparent Magnitude to the true, were found by the former Reasoning to be as 90 to 1, yet it is not really any greater than as 30 to

to 1; because the right Line E F wou'd only appear 30 times greater by the Assistance of the *Microscope*, than it wou'd if view'd by the naked Eye at the distance of 8 Inches. For we are not to consider how much, by means of the *Microscope*, we magnify an Object at the distance of two Feet; but how much greater it is made, than when view'd at that distance, to which we bring our Eye when we desire to look at any thing more curiously.



Of



Of the LIGHT and APERTURE of
MICROSCOPES.

Upon the *Aperture* of *Microscopes* all their Effects and Virtue entirely depend; so that from hence it is that we are to learn to what degree the magnifying of Objects may be brought; which no Body that I know of, has hitherto determined. And it will be found that we may here proceed *ad infinitum*, as shall be shown in *Telescopes*, not indeed in a single *Microscope* of one little *Lens*, but in those which are made by a combination of more than one.

In *Microscopes* made of a single *Lens*, it is to be observ'd, that if their *Focal Distance* be about half an Inch or greater, there will be no occasion for limiting the *Aperture* in order to make distinct Vision; because the very narrowness of the *Pupil* of the Eye excludes, as much as there is occasion, those Rays that disturb Vision, and as much as they wou'd be excluded if the *Lens* were made to have a less *Aperture*. But in smaller *Lens's* where this Limitation of the *Aperture* is necessary, the Rule is that the *Diameters*

ameters of those Apertures shou'd be in the same proportion with the *Focal Distances* of the respective *Lens's*, in order to have the Object seen by both equally distinct. But the *Light* or *Brightness* will be in a duplicate proportion of those *Focal Distances*; so that the more convex the *Lens* is, the greater indeed, but then the more obscurely will every thing be seen,

[*Fig. 14.*] Let *P* be a small *Lens*, whose *Axis* is *TBF*, *PD* the *Semidiameter* of the *Aperture*, which Experience teaches to be the greatest that can be admitted, and that less than the *Pupil of the Eye*, *F* the *extreme Focus of Red Ray*, (which are least refrangible) proceeding parallel to the *Axis*, in which Point suppose the *Object* to be placed, and *B* the *Focus of Violet colour'd Rays*, which are most refrangible. The same things being supposed in a smaller *Lens p*, the *Semidiameter* of whose *Aperture pd* is to the *Focal Distance pf* in the same proportion as in the greater; I say the *Object* will be seen equally distinct in both.

For since the Ray *ED* parallel to the *Axis* falling upon the *Lens P* is refracted unequal and divided into its *extreme Colours* by the Angle *FDB*, so that the *extreme Colour Red* passes to *F*, and the *extreme Violet-Colour* to *B*; it will happen on the contrary, that a Ray *FD* proceeding from the *Object* will be divided into it's extreme

treme Colours by an Angle E D K, equal to F D B. Therefore in both Cases F D B is the *Angle of Aberration*, upon which depends the Aberration of the Rays in the bottom of the Eye, as shall be shewn when we come to speak of *Telescopes*. But since from the Nature of this *Aberration*, P F is to F B, as pf to fb ; and also by construction P D is to P F, as pd to pf ; it follows that the Angles as well P F D, $pf d$, as P B D, $pb d$ are equal. Wherefore the difference of the Antecedents, P F D, P B D, is equal to the difference of the Consequents $pf d$, $pb d$; that is, the Angle F D B is equal to the Angle $fd b$, and consequently the *Aberrations* in the bottom of the Eye are in both Cases equal, and by that means Vision equally distinct.

Moreover because the Angles P F D, $pf d$ are equal, it is plain that the same quantity of Rays in both cases proceeds from the same points of the Object F and f or any others, upon the *Lens's*, and from thence to the Eye. But the Breadth of the Object in the Bottom of the Eye is in the smaller *Lens* so much greater, as P F is greater than pf , as has been before demonstrated; and the apparent Surfaces are in a duplicate proportion of those Breadths. Therefore the same quantity of Lucid Rays expended towards illustrating each Surface, will make that which is least

the clearest by so much as the other Surface is greater, that is, in a duplicate proportion of $P F$ to pf , which was the last thing to be demonstrated.

Since therefore the same perfection of Vision which is to be found in larger *Lens's*, cannot be had in more convex *Lens's* without diminishing the Brightness of the Object at the same time ; It follows that we cannot proceed in magnifying as much as we please, unless a greater Light be borrow'd somewhere else to illustrate the Object. Nor will this be of any great Benefit, because the Latitude at the Pupil of the Eye, or the little Cylinder of Rays flowing from every point of the Object, and which has here the same Latitude with the Aperture, cannot be contracted farther than the fifth or sixth part of a *Line* ; so that even this limits the Efficacy of these little *Lens's*.

The Effects of more compounded *Microscopes* will easily be accounted for, after the same manner. And indeed a full Consideration of *Prop. XXI.* and *XXII.* is sufficient for explaining the Effects of all sorts of compound *Microscopes*.



Of T E L E S C O P E S.

P R O P. III.

A Telescope made by a convex and concave Lens represents vastly distant Objects distinct and erect; and magnifies them according to the Proportion of the Focal Distance of the convex Lens to the Focal Distance of the concave Lens.

Fig. 15. Let AO be the common Axis of both Lens's; and A the extreme *Convex Lens*, whose *Focus* of parallel Rays proceeding from the vastly distant *Object* is suppos'd to be at O. Let D be the *Concave Lens*, which is so placed between the *Lens* A and its *Focus* O, that the same Point O may also be the *Focus* of the *Concave Lens*, where Rays falling parallel from the side of O would be collected. And first suppose the Eye of the Spectator placed next to this *Lens*.

The Rays then proceeding parallel from each Point of the vastly distant *Object*, and falling upon the *Lens*, those which proceed from

from that Point of the *Object*, which is in the Axis produced, would be collected at the Point O ; but they are again made parallel by means of the Lens D. We would have the Rays fall parallel upon the Eye, that the Telescope may be fitted for those who have good Eyes ; For we shall speak afterwards of short-sighted Eyes. In like manner the Rays proceeding from those Points of the *vastly distant Object* which are out of the Axis, would be all collected at respective Points near O ; but these also by *Refraction* at the *Lens D* are again made parallel, tho' something Oblique to the Axis A D, which Rays, to avoid Confusion, are not express'd in the Figure. Therefore the Rays which proceed from the *vastly distant Object* being made to fall parallel upon the Eye, will make *distinct Vision* ; and since those Rays that proceed from the *Object* go on to meet the Eye in the same order, 'tis plain the apparent Position of the *Object* will be the same with the true, or the *Object* will be *erect*.

Fig. 16. The *Lens's* A C and D, and the Point O being placed as before, find by Prop. XVI. the Point P, to which Rays tending, will by *Refraction* at the *Lens A C* be collected at D the Centre of the *Lens D* : Which Point is also found, by making DP a third proportional to D O, D A, and taking it on the same side with D O.

DO. Suppose the Ray ECP to be one of those which proceed from the extreme right side of the *vastly distant Object*, which imagine to be the Moon, and its Centre to be placed in the Axis DA produced. It is plain that this Ray will come to the Eye in the Right Line CDF, because it passes thro' the Centre of the *Lens* D, whose middle Thickness may be neglected as inconsiderable, and its two Surfaces in that place look'd upon as parallel. But we have shewn before that all the Rays proceeding from each Point of the Moon, will by means of such a Telescope fall parallel upon the Eye. Wherefore the Eye will receive all the Rays from that Point E of the Moon in such manner, as that they shall be parallel to the Ray CDF; and consequently will see that Point of the Moon in the place to which the Right Line DC tends, which tending to the same side of the Axis, on which that point of the Moon is situated, from whence the Rays proceeded, 'tis plain the Object will appear *Erect*. Moreover the Angle ADC determines the Semidiameter of the Moon, as encreas'd by the Telescope. But the Angle CPA is that which determines its Semidiameter, as seen by the naked Eye, because we before supposed the Ray ECP to proceed from the extreme right side of the Moon, and the Ray HAP from its Centre: For tho' the Point P is beyond the Eye,

Eye, and the Eye sees from the Point O, yet the Moon being an Object vastly distant will appear under the same Angle to the naked Eye, whether it be view'd from the Point P or O. Therefore the Moon will appear magnified, according to the Proportion of the Angle ADC to APC, which Proportion may here be look'd upon as the same with that of PA to DA. But because by Construction, DO is to DA as DA to DP; by inverting and compounding the Proportion AO will be to OD, as PA to AD. Wherefore the apparent Magnitude will be to the true as AO to OD. *Q. E. D.*

It appears from hence that the apparent Magnitude is the same, in whatsoever place behind the *Lens* D the Eye is situated.

Fig. 17. Let the *Lens's* AC and D be placed as before, and let AQ be taken in their Axis produced equal to AO. And out of those Rays which proceed from a Point of the right side of the Moon, let us consider the Ray RQC passing thro' the Point Q, (for some one will pass thro' it) and meeting with the *Lens* AC in C. It will afterwards become parallel to the Axis AD, and when refracted again at the concave *Lens*, will diverge, as if it came from the Point L, and will tend to the Eye in the Right Line LIF, so as that the Distance LD may be equal to DO, because

cause L is in that Case the *Focus* of parallel Rays falling upon the *Lens* D.

The Proportion of the encreas'd Magnitude is hence easily collected. For because the Rays proceeding from the right extreme of the Moon, after having passed both *Lens*'s, come parallel to the Pupil G F, and consequently they become parallel to the Ray L I F, which we know to be one of them; that Point of the Moon will appear in the Right Line I L, and consequently the Semidiameter of the Moon will be comprehended in the Angle I L D. But the Angle in which the Semidiameter wou'd appear to the naked Eye, either from D or from Q, is R Q H, or C Q A. Therefore the Proportion of the encreas'd Magnitude is the same with that of the Angle D L I to A Q C; that is, because of A C, D I equal, as A Q to L D. But A Q is equal to A O, and L D is equal to D O. Therefore the Proportion of the apparent Magnitude to the true is as A O to O D. Q. E. D.

To determine what will be the Amplitude of the visual Angle, or of the Space which is represented at one view by a Telescope consisting of a Convex and Concave Lens.

M

Fig.

Fig. 18. The Amplitude of the *visual Angle* in these *Telescopes*, depends chiefly upon the Magnitude of the Pupil of the Eye, which is confirm'd by Experiment. For if applying your Eye to the *Telescope*, you first shut it that the Pupil may be dilated as it usually is in the Dark, and then open it on a sudden; at first view you will discern Objects in a larger Orb than a little while afterwards, the Orb being presently contracted as soon as the Pupil is contracted by the Brightness of the Light. But if you place a Plate perforated with a small Hole before the Eye, you will discern every Object in a lesser Orb.

If you make the Hole extremely small, the lucid Orb will not be contracted in proportion to the smallness of the Hole; but its Amplitude will then be limited by the Aperture of the Convex *Lens*, and consequently will not be diminished beyond a certain Degree, except the Convex *Lens* be also more contracted. The Reason of which is very easy to be explain'd. For if E F be the convex *Lens*, and B the concave, to which the Pupil of the Eye apply'd has first the Magnitude C D: Draw from the opposite Points C, D in the Circumference of the Pupil, through the Centre of the *Lens* A the right Lines C A H, D A G. These will determine the *visual Angle*, under

der which that part of any Object which is seen at one view is comprehended: Because Rays coming from the Points G, H thro' the Centre of the *Lens A*, penetrate without Inflection to C and D; therefore that part of the Object which is comprehended within the Angle G A H, cannot but send Rays to the Eye, even tho' the Pupil were a little narrower than D B C. For drawing G A K so as to make A K equal to A O, and joining E K; If so be E K fall upon the Pupil, the Object comprehended under the Angle G A H will be discern'd, but the extreme Points towards which the right Lines A G, A H tend, will be seen but obscurely, because only a small part of the Rays, which they cast upon the *Lens E F* enter the Pupil. And hence it happens, that how much foever the Aperture of the *Lens E F* is contracted, the Amplitude of the *visual Angle* is nevertheless not at all, or extremely little diminished, so the Orb of the Pupil be not contracted. But this breadth of the Pupil being diminished, and reduced as it were to a Point, the Amplitude of the *visual Angle* is the same with that of the Angle E P F; E F being supposed the Aperture of the convex *Lens*, and the Point P found by Prop. XVI. or by making B O (the Distance of the concave *Lens* from the *Focus*

of the convex) BA and BP in a continual Proportion. For no Rays, transmitted through the *Lens A* can arrive at the Point of the Eye B, but such as before they fall upon that *Lens* tend towards the Point P. The greatest Angle EPF of which Rays is determin'd by the Aperture of the *Lens A*,

This is the *Telescope* which was first found out by *Galileus*, and still retains his Name: And is the same with a common Perspective Glass.





P R O P. IV.

A Telescope made of two Convex Lens's represents vastly distant Objects distinct but inverted, and magnifies them according to the Proportion of the Focal Distance of the Exteriour or Object Lens, to the Focal Distance of the Interiour or Ocular Lens.

[Fig. 19, 20.] Let *AC* be the exteriour convex *Lens*, *D* the interiour, *AD* the common *Axis* of both, and *O* the *Focus* of the *Lens* *AC*. Let the other Convex *D* be so placed, that the same Point *O* may be also its *Focus*, or the Point of concourse of parallel *Rays* coming from the side of *G*, where the *Eye* is supposed. We are to shew that all this being supposed, vastly distant Objects will be seen distinct, and inverted and magnified, according to the Proportion of *AO* to *OD*.

And

And here we must make use of two several Figures, as in the preceding *Proposition*, in the first of which the Rays coming parallel to the Axis H A, are by the Refraction of the Lens A C collected at its *Focus* O, and from thence tending farther to the *Lens* D, are by it again made parallel to the Axis A D, and so come to the Eye placed at G. And as in the preceding *Prop.* we must again consider this composition of parallel Rays, as coming from a single Point of the vastly distant Object, which is placed in the Axis H A D, as suppose from the Centre of the Moon; and the like parallel Rays coming from every other Point of the Object upon the *Lens* A C, as suppose from the extreme right side of the Moon, which are inclin'd to the former, and being thereby refracted, are collected in a Point of the Axis near O, where intersecting themselves, and proceeding to the *Lens* D, they are again made parallel, (that is, only among themselves respectively) and so arrive at the Eye. Whence 'tis plain, Vision will be made distinct.

The other Figure shews the inverted Situation, and the Proportion of the encreas'd Magnitude of the *Object*. Where the convex *Lens*, A C and D, and their common *Focus* O being placed as before; and moreover, as in the second Demonstration of preceding *Prop.* the Distance A Q being made

made equal to $A O$; the remaining part of the Demonstration will proceed much after the same manner. For if out of the Rays which proceed from a Point in the extreme right side of the Moon, we choose one $R Q C$ passing thro' the Point Q : That, after refraction at the *Lens* $A C$ will pass in $C I$ parallel to $A D$, and being again refracted by the *Lens* D , will tend along the right Line $I F L$ to the Point L , taken in such manner that the Distance $D L$ is equal to $D O$. But because the Rays from the extreme right side of the Moon, after refraction at both *Lens*'s, arrive parallel at the Eye, as has been said before, and $I F L$ is one of them; It follows that they will all fall parallel to $I F L$ upon the Eye, and that Point of the Moon will be seen in a Place, according to the right Line $F I$; which since it tends to the opposite side to that from whence those Rays came, 'tis plain that the situation of the Moon will appear inverted, so that the right side will be changed to the left, and the upper Parts to the lower. Moreover since the Centre of the Moon will be seen in the right Line $D A$, $I L D$ will be the apparent Angle of the Semidiameter of the Moon. But to the naked Eye, that Semidiameter is comprehended under the Angle $H Q R$, or $A Q C$. Therefore the *Ratio* of the apparent Magnitude to the true, is as the Angle $D L I$ to $A Q C$, that is as $A Q$ to $D L$; because $A C$,

DI

DI are equal, that is as AO to OD.
Q. E. D.

And here likewise it appears that it signifies nothing to the apparent Magnitude, wheresoever the Eye is placed behind the *Lens* D. But that it may comprehend most at one view, it is convenient it should be placed at or near the Point L: Because it appears, that altho' the Breadth of the *Pupil* be ever so little, yet the whole *Lens* D while it does not exceed the Aperture of the *Lens* A C (for it is usually confined within this Measure) will be seen full of the Object.

This is the *Telescope* most commonly used, to look at Celestial Bodies.



P R O P.



PROP. V.

TO explain the Construction of a Telescope compounded of four Convexes, by means of which Objects are seen erect and very amply.

[Fig. 21, 22.] Telescopes made of two Convexes, because of their Inverting the position of the Object, are seldom used, except in observing the Stars, the position of which is not regarded. The proportion in which this Sort magnifies the Object has already been demonstrated. But if we wou'd have these Images again made *erect*, and at the same time a great share of them be represented to the Eye at one View very *amply*, we must use 3, 4, 5 or more *Lens's*. Which however are not to be multiply'd without Cause, because the Matter of each of them and the Reflexion of their several Surfaces divert part of the Rays. But we cannot obtain the desir'd Effect perfectly, with fewer than 4 *Lens's*. For altho' in the same Length of the *Telescope* both an *erect* Situation and the same degree of magnifying, and an equal share of the Object may be had as well with 3 as 4 *Lens's*;

N com.

composition of 3 *Lens*'s is much more convenient than that of 4; because in that, the two Ocular *Lens*'s, or at least that which is next the Eye must be made of larger Segments of a Sphere, with respect to its Diameter, or to the *Focal* distance, if the same Magnitude of the visual Angle be requir'd. And hence the Objects come to be colour'd, and right Lines, at the Margins of the Aperture, appear Curve. Therefore we must make our Telescope of 4 *Lens*'s, which is done after the following Manner.

The exterior or Object *Lens* is A, whose *Focal* Distance is A B, and in the same Axis are placed three Ocular *Lens*'s, C, D and E, all equal to one another, the inmost of which is placed beyond the *Focus* B, by its *Focal* distance BC, and the next D is placed beyond C, by twice that distance BC, and the last as far from D as that was from C, and lastly the Eye must be placed beyond this last by the distance BC.

There is here again occasion for two Figures, in the first of which are represented Rays proceeding from a single Point of the vastly distant Object: Which, 'tis plain to any who understand what has gone before, first fall as it were parallel upon the *Lens* A, and are by it collected at its *Focus* B, and thence diverging fall upon the *Lens* C, which makes them again parallel and throws them upon the *Lens* D, which collects them at its

Focus

Focus H, the middle Point of the distance *D E*, from whence proceeding on to the *Lens E* they are by it made a third time parallel, and being receiv'd so by the Eye *F*, they make distinct Vision by being collected at its Focus which is in the bottom of the Eye.

The other Figure considers the *Proportion of Magnifying*, which is, *That which A B the Focal Distance of the Object Lens bears to B C the Focal Distance of one of the Ocular Lens's*. And demonstrates likewise the Amplitude of the Visual Angle. For the Apertures of the three Ocular *Lens's* being supposed equal, which must not exceed the Aperture of the Object *Lens A*; draw *M Q, N R* parallel to the common *Axis*, and comprehending the Diameters of the Apertures of the *Lens's E* and *D*. And also *K O, L P* parallel to the same *Axis*, and comprehending *K L* the Aperture of the *Lens C*: and taking *A G* equal to *A B*, draw the right Lines, *O G U, P G T* intersecting one another in *G*. Now it is evident the Latitude of the Object which is seen by the naked Eye from the Point *G*, and consequently from *F* also, the distance of the Object being as it were infinite, wou'd appear comprehended in the Angle *T G V*; if seen thro' the Telescope wou'd appear comprehended in the Angle *M F N*: And consequently the proportion of the apparent Magnitude to the true, is as the Angle *M F N* to the Angle *T G V*

TGV or PGO; that is, PO and MN being equal, as the distance AG to the distance EF; that is, as AB the *Focal* distance of the Object *Lens* to BC the Focal distance of one of the Ocular *Lens*'s. Q. E. D.

It appears moreover that the Visual Angle MFN comprehends the same Latitude of the Object, with a Telescope made of two *Lens*'s only A and C; for that share of the Object which is comprehended in the Angle TGV, wou'd be seen thro' that Telescope in the Angle KSL equal to the Angle MFN.

This incomparable Composition of *Lens*'s was found out by I know not whom at *Rome*, and may be much improv'd by placing an *Annulus* or Ring either at H, the common *Focus* of the *Lens*'s D and E, or at B the common *Focus* of the *Lens*'s A and C; which is especially of very great Use in measuring the Diameters of Planets. For this *Annulus* does therefore exactly circumscribe the Circle of the apparent Images, because it cuts off those irregular Rays which are not collected near enough to B or H, and consequently are not by means of the succeeding *Lens*'s sent parallel to the Eye, which distinct Vision requires; And the Colours likewise near the Margins are by this Contrivance taken away, which without it are not well to be avoided.

It may seem a little strange that the Colours of the Iris arise no more in this Telescope by the Refraction of so many Ocular Lens's, than in that where there is but one ; but to any one that will consider it, the Reason will be very Obvious. For the *Lens* QR corrects and takes away those Colours, which the *Lens* KL produc'd, their Spherical Surfaces being equal by Construction.

Of the APERTURE of the LENS's.

Since the proportion of magnifying in *Telescopes* made of two *Lens*'s has been shewn to be that which the *Focal Distance* of the Object *Lens* bears to the *Focal Distance* of the Ocular *Lens*, it may be thought perhaps that however short the *Telescope* be, the Object may be magnify'd in any assigned proportion. But there are two Causes which make this impossible ; One is, that the *Aperture* of the Object *Lens* remaining the same, the more we magnify the Object by using a less Convex Ocular *Lens*, the more obscure we make them appear. The other is, that it represents them less distinct. And if we expect a Remedy by encreasing the *Aperture*, the Confusion will be the more encreased. What belongs to the *Brightness* or *Obscurity* will be easily understood by attentively considering the Image of any Object painted upon the Bottom of the Eye ; which the greater

greater it is made, whether by means of the Refraction of *Lens's*, or only by approaching nearer, in so much greater plenty must the Rays from every Point be receiv'd within the Eye, in order that the same *Brightness* may still remain. For if looking at an Object with the naked Eye, you approach to it twice as near, the Image at the Bottom of the Eye will be twice greater in Diameter, and four times in Area. But four times more Rays do also from every Point of it enter the Pupil of the Eye; because the Angle made by the Cone of Rays becomes twice as large; and therefore it is that the same *Brightness* of the Image is perceiv'd at both Distances, which is the Contrivance of Nature. But if a *Telescope* were to be made which shou'd magnify the Diameter of any Object ten times, and represent it as bright as when it is look'd at with the naked Eye; the Diameter of the Aperture of the Object Lens ought to be ten times greater than the Diameter of the Pupil, altho' no part of the Rays were intercepted by the Reflection of the Surfaces of each *Lens*, or by the Colour of the Glass. For by this means, when the Surface of the Object is magnified an hundred times we have also an hundred times more Light than was receiv'd by the naked Pupil.

But a much less Measure of *Brightness* suffices for *Telescopes*; for those which we use in the Day-time, are not too obscure, if they

they have but $\frac{1}{2}$ or $\frac{1}{3}$ of that *Brightness* which is usually a perceiv'd by the naked Eye. But those longer ones with which we observe the Moon and the Planets, require not above half this last *Brightness*, because the Eye is moved with a much less Brightness in the Night than in the Day. So that in a *Telescope* 30 Feet long, which magnifies the Diameters of the Planets 109 times, and wou'd consequently require the Diameter of the Aperture of the Object *Lens* 109 times greater than the Diameter of the Pupil, that is of about 11 Inches, if we suppose the Diameter of the Pupil to be $\frac{1}{10}$ of an Inch; we find that an Aperture of 3 Inches in Diameter suffices, which admits less than $\frac{1}{3}$ of that *Brightness* which wou'd be admitted by an Aperture of 11 Inches.

The Proportions between the *Focal Distances* of the Object *Lens*, (which is likewise the Length of the *Telescope*) the *Aperture* of the same Object *Lens*, the *Focal distance* of the Ocular *Lens*, and the apparent Magnified Diameter of the Object, for *Telescopes* from the Length of 1 Rhinland Foot to 100, are express'd in the Table following.

T A B L E

88 TABLE for TELESCOPES.

The Focal Distance of the Object Lens, or the Length of the Telescope.	The Diameter of the Aperture of the Object Lens.	The Focal Distance of the Ocular Lens.	The Proportion of Magnifying, consider'd as to Diameter.
<i>Rbinland Feet.</i>	<i>Incches, & Decimals.</i>	<i>Incches, & Decimals.</i>	
1.	0,55.	0,61.	20.
2.	0,77.	0,85.	28.
3.	0,95.	1,05.	34.
4.	1,09.	1,20.	40.
5.	1,23.	1,35.	44.
6.	1,34.	1,47.	49.
7.	1,45.	1,60.	53.
8.	1,55.	1,71.	56.
9.	1,64.	1,80.	60.
10.	1,73.	1,90.	63.
13.	1,97.	2,17.	72.
15.	2,12.	2,33.	77.
20.	2,45.	2,70.	89.
25.	2,74.	3,01.	100.
30.	3,00.	3,30.	109.
35.	3,24.	3,56.	118.
40.	3,46.	3,81.	126.
45.	3,67.	4,04.	133.
50.	3,87.	4,26.	141.
55.	4,06.	4,47.	148.
60.	4,24.	4,66.	154.
65.	4,42.	4,86.	161.
70.	4,58.	5,04.	166.
75.	4,74.	5,21.	172.
80.	4,90.	5,39.	178.
85.	5,05.	5,56.	183.
90.	5,20.	5,72.	189.
95.	5,34.	5,87.	194.
100.	5,48.	6,03.	199.



P R O P. VI.

TO explain the Manner of fitting a *Telescope* for observing *Eclipses* of the Sun, and discovering the Spots in its Surface, and to determine how great its Image will be represented.

[Fig. 23, 25.] A *Telescope* is found to be of great Use in observing *Eclipses* of the Sun, and also in discovering the Spots which are said to be in its Surface; by receiving the Image form'd by both it's *Lens*'s upon a white Plane, from which the Light is every other way excluded. In order to explain which Invention we must first demonstrate the Position of the *Lens*'s which is necessary to form the Image of the Sun, as clear and distinct as may be.

Let A B be the convex *Lens* next the Sun, whose *Focus* is E. The other is D, either concave or convex, for either of these sorts of *Telescopes* will do the Business, tho' a *Telescope* of two *Convexes* is the most convenient, because we make it represent the

O Images

Images erect, while by the other sort we invert them. Let the point K be the *Focus* of the *Lens* D, where Rays coming from the side of H parallel are after Refraction by it collected, and in H suppose the white Plane placed in order to receive the Image of the Sun. Which that it may appear distinct and nicely terminated, 'tis necessary that the Rays which proceed from any one Point of the Sun, and which fall parallel upon the *Lens* A B, shou'd again be collected in one point upon the Plane. Wherefore the Distance between the *Lens*'s A B and D ought to be something greater than in the common Disposition of the *Telescope*, or than when it is fitted for a good Sight; and the position of the *Lens* B ought to be such that the Rays which wou'd otherwise tend to the *Focus* E of the *Lens* A B, may be diverted and brought to H, which may be done by Prop. XVI. or by taking E K, E D, E H in a continual Proportion. But in the common Disposition of the *Telescope*, the *Focus* K is required to coincide with the *Focus* E, as has been shewn above. So that here the distance of the *Lens*'s is increas'd by the Space E K, which will always be so much less, as the Distance E H is increas'd. For the Distance D K which is the given *Focal* Distance of the *Lens* D, is divided in such manner in E, that H E is to E D, as the same E D to E K.

[Fig.]

[Fig. 24, 26.] How great the Diameter of the Image of the Sun will appear upon the Plane H may thus be determin'd. Draw from the Centre of the *Lens* A B to the *Lens* D the right Lines B P, B Q, comprehending an Angle equal to that, under which the Sun's Diameter appears without a *Telescope*; And make B G a third Proportional to B K, B D, and join G P, G Q, which will consequently give G the *Focus* of Rays diverging from B after refraction at the *Lens* D, which produce till they meet the Plane placed at H in the Points L, M. I say L M will be the Diameter of the Sun's Image represented upon the Plane L H M. For produce P B, Q B towards O and N. Therefore since from the extreme Point of the real Diameter of the Sun on the right Hand side, Rays are sent upon the whole *Lens* A B, which are all to be look'd upon as parallel among themselves, and to the right Line O B; one of them will proceed along the right Line O B, and penetrating the *Lens*, go on in the right Line B P, because B is the Centre of the *Lens*, whose Thickness is here neglected. For the same Reason, one of the parallel Rays from the left extreme of the Sun's Diameter, will proceed along the right Line N B Q. But moreover, both will be refracted in such manner by the *Lens* D, that diverging from

their *Focus G*, they will proceed along the right Lines *PL, QM*, which are the right Lines *GP, GQ* produced. Therefore 'tis plain the Point in the right Hand extreme of the Sun will be represented at *L*, and the opposite Point in the left extreme at *M*. For since the Image of the Sun is required distinct, it is necessary, that where one Ray proceeding from any Point of it falls upon the Plane, all the rest which proceed from the same Point should be collected there also. Therefore the Diameter of the Image is *LM*; And by a *Telescope* made of a Convex and Concave, the Image is inverted, and by one made of two Convexes it is represented erect. *Q. E. D.*

But it must be observ'd, that the greater *J. M* the Image of the Sun is, the *Lens's* *AB* and *D* remaining the same, the less clear and distinct will it be. For if all the Rays descending from the Sun upon the *Lens AB*, should possess a Space in the place *LHM* equal to the breadth of the *Lens AB*; that is, if they were to form the Image of the Sun equal to the Aperture of the *Lens AB*, this Image would be as clear as if the Plane were enlightned by the Sun without the Interposition of *Lens's*. No respect being had to those Rays which the *Lens's* reflect, or by reason of their imperfect Transparency do not transmit, which perhaps occa-

occasions a loss of above half the whole Number of Rays. But if the Sun's Image be made larger, which is necessary to be done in Observations of this Nature, it will then be so much the more obscure. But Experience is the best Judge to determine in what Magnitude it will be most convenient to represent the Sun's Image in these Observations ; by trying first one, and then another Distance of the Plane from the Telescope. Where it is to be observ'd, that as we encrease this Distance, the Distance between the *Lens's* A B and D ought to be a little diminish'd, in order to preserve the distinctness of the Image ; the Reason of which has already been given.



